the OBSERVER'S HANDBOOK 1976



sixty-eighth year of publication

the ROYAL ASTRONOMICAL SOCIETY of CANADA

editor: JOHN R. PERCY

THE ORIGINS OF THE ROYAL ASTRONOMICAL SOCIETY OF CANADA

In the mid-nineteenth century, in the bustling Lake Ontario port city of Toronto, there were no professional astronomers. However, many inhabitants of the city were keenly interested in sciences and current developments in them. King's College, which grew into the University of Toronto, had been started in 1842. In 1849 it had 36 undergraduates attending, and had graduated a total of 55 students in the three faculties of arts, law and medicine. The Toronto Magnetic Observatory had been established in 1840. Its early directors and observers were officers and soldiers in garrison. Some of them, such as Captain J. F. Lefroy, contributed much to the cultural life of the city. Out of this body of interest came the Canadian Institute established in 1849 "to promote those pursuits which are calculated to refine and exalt a people".

Besides holding weekly meetings, the Canadian Institute accumulated an outstanding library. There many hours were spent in study by Andrew Elvins who had come to Canada from Cornwall in 1844. In 1860 he moved to Toronto, with a population then of 44,000, and became chief cutter in a well known clothing store on King Street. While the Canadian Institute held discussion meetings of all sciences, Elvins wished to concentrate on astronomy. For this purpose he gathered together a few like-minded friends.

On December 1, 1868 The Toronto Astronomical Club met for the first time, at the Elvins' home, "having for its object the aiding of each other in the pursuit of astronomical knowledge". The thousands of meteor sightings of the Leonid showers made in Toronto in November 1867 and 1868 had doubtless encouraged the project. In May, 1869 the word "Club" was changed to "Society". Written records were kept for the first year, until the secretary moved away. After that, the group met only sporadically, but by the distribution of materials Elvins kept interest alive.

As the century wore on, Elvins, who lived till 1918, acquired more kindred spirits, some of them influential and prominent. As a result, on March 10, 1890 the organization was incorporated as The Astronomical and Astrophysical Society of Toronto. In May, 1900 chiefly through the efforts of one of the important early members George E. Lumsden, the name was changed to The Toronto Astronomical Society. On March 3, 1903 through legal application the name took on its current form, The Royal Astronomical Society of Canada. For many years the Society had its offices and library in the Canadian Institute buildings, and held meetings there.

Early in the 1890's, Dr. Clarence A. Chant of the University of Toronto became deeply interested in the Society. The impetus which he gave to it until his death in 1956 still lingers. During its first fifteen years the Society published annually volumes containing its Transactions and Annual Report. In 1907 Dr. Chant started The Journal of the Royal Astronomical Society of Canada, and this Handbook, called then "The Canadian Astronomical Handbook". It is a remarkable fact that at the time of his death Dr. Chant had been the Editor of both the Journal and the Handbook for exactly 50 years. During this period he received generous assistance from many of the Society's members. At times the Journal was published monthly, but currently it is bi-monthly.

The change of name in 1903 led immediately to the concept that the Society should not be limited to Toronto, but should become national in scope. The second Centre to be established was that of Ottawa in 1906, where the Dominion Observatory was being established. Now the Society has 18 Centres from sea to sea across Canada, as listed elsewhere in this Handbook. The growth in membership to nearly 3000 also shows its flourishing state.

HELEN SAWYER HOGG

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252 College Street, Toronto M5T 1R7, Canada

editor: JOHN R. PERCY

THE ROYAL ASTRONOMICAL SOCIETY OF CANADA Incorporated 1890 – Royal Charter 1903 – Federally Incorporated 1968

The National Office of the Society is located at 252 College Street, Toronto, Ontario M5T 1R7; the business office, reading room and astronomical library are housed here.

Membership is open to anyone interested in astronomy and applicants may affiliate with one of the eighteen Centres across Canada established in St. John's, Halifax, Quebec, Montreal, Ottawa, Kingston, Hamilton, Niagara Falls, London, Windsor, Winnipeg, Saskatoon, Edmonton, Calgary, Vancouver, Victoria and Toronto, or join the National Society direct.

Publications of the Society are free to members, and include the JOURNAL (6 issues per year) and the OBSERVER'S HANDBOOK (published annually in November). Annual fees of \$12.50 (\$7.50 for persons under 18 years) are payable October 1 and include the publications for the following calendar year.

Algol, Minima/35-57, 98 Anniversaries, Festivals/3 Appulses/60 Asteroids/74-75 Calendar/Inside Back Cover Clusters/106-107 Comets/80 Constellations/5 Coordinates/4 Delta Cephei/98, 109 Eclipses/59 Galaxies/108-109 Greek Alphabet/4 Impact Craters/81-82 Julian Day Calendar/11 Jupiter, General/29 Jupiter, Belts, Zones/79 Jupiter, Long. Central Meridian/76 Jupiter, Satellites, Configurations/ 35 - 57Jupiter, Satellites, Phenomena/77-79 Mars, General/28 Mars, Long. Central Meridian/73 Mars, Map/73 Mercury/28 Messier's Catalogue/105 Meteors/80-81 Miscellaneous Ast. Data/6 Moon, Observation/32 Moon, Map/72 Moon, Rise and Set/22-27 Nebulae, Galactic/104 Neptune/31 Occultations, Grazing/68-71

INDEX

Occultations, Lunar/61-66 Occultations, Planetary/60 Planets, General/28-31 Planets, Elements/8 Pluto/31 Precession/84 Radio Sources/110 Satellites, Solar System/9 Saturn/29 Saturn. Satellites/83 Sky Month By Month/33-57 Solar System, Elements/8 Solar System. Satellites/9 Star Maps/111-116 Stars, Brightest/86-96 Stars, Clusters/106-107 Stars, Double/97 Stars, Names/85 Stars, Nearest/102-103 Stars, Variable/98-101, 109 Sun, Ephemeris/7 Sun, Physical Observations/58 Sun, Rise and Set/14-20 Sun, Sunspots/59 Symbols, Terminology/4 Time, Correction to Sundial/7 Time, Radio Time Signals/11 Time, Solar, Sidereal etc./10 Time Zones, Map/12 Twilight/21 Uranus/29 Venus/28 Visiting Hours/13

THE OBSERVER'S HANDBOOK 1976

THE OBSERVER'S HANDBOOK for 1976 is the sixty-eighth edition. I wish to thank all those who assisted in its preparation: those whose names appear in the various sections, those mentioned below, and especially my editorial assistant, John F. A. Perkins.

There are several major changes in this year's HANDBOOK. You will notice that the HANDBOOK no longer carries advertisements. The decision to discontinue advertisements was made by the Council of the Society, in order for the HANDBOOK to qualify (under federal law) as a non-commercial publication. We thank our advertisers for their support over the years.

The pages which were previously occupied by advertisements have been used to ease the crowding of material in the various sections, and to accommodate new material. There is now a set of six new star maps, drawn by John Perkins, and two extra pages of material on variable stars, provided by Janet Mattei of the A.A.V.S.O. Dr. John F. Heard has expanded his descriptions of "The Sky Month By Month", and Dr. Helen S. Hogg has provided a capsule history of the Society, on the inside front cover.

My thanks go to Dr. Ian Halliday for revising the page of "Miscellaneous Astronomical Data", to Dr. Cecil Costain for providing information on standard time zones, to Drs. Donald MacRae, C. T. Bolton and R. F. Garrison for revising the section on "The Brightest Stars", and to many readers of the HANDBOOK for spotting errors and inaccuracies and for making helpful suggestions. Special thanks also go to Maude Towne and Isabel Williamson, who have calculated the tables of moonrise and moonset for many years and who are now retiring from this onerous but important task. The David Dunlap Observatory and Erindale College, University of Toronto, have once again provided financial, technical and moral support for the HANDBOOK.

Finally, my deep indebtedness to H.M. Nautical Almanac Office (particularly Leslie V. Morrison, Gordon E. Taylor and Superintendent Dr. G. A Wilkins) and to the *American Ephemeris* is gratefully acknowledged.

JOHN R. PERCY

New Year's Day Thur.	Jan.	1	Pentecost (Whit Sunday)	June	6
EpiphanyTues.	Jan.	6	Trinity Sunday	June	13
Accession of Queen			Corpus Christi	June	17
Elizabeth (1952)Fri.	Feb.	6	St. John Baptist		
Septuagesima Sunday	Feb.	15	(Mid-Summer Day)Thur.	June	24
Quinquagesima			Dominion DayThur	July	1
(Shrove) Sunday	Feb.	29	Birthday of Queen Mother	•	
St. David Mon.	Mar.	1	Elizabeth (1900)Wed.	Aug.	4
Ash Wednesday	Mar.	3	Labour DayMon.	Sept.	6
St. Patrick	Mar.	17	Jewish New Year	-	
Palm Sunday	Apr.	11	(Rosh Hashanah)Sat.	Sept.	25
First Day of Passover Thur.	Apr.	15	St. Michael	•	
Good Friday	Apr.	16	(Michaelmas Day)Wed.	Sept.	29
Easter Sunday	Apr.	18	Yom KippurMon.	Oct.	4
Birthday of Queen	-		ThanksgivingMon.	Oct.	11
Elizabeth (1926) Wed.	Apr.	21	All Saints' Day Mon.	Nov.	1
St. GeorgeFri.	Apr.	23	Remembrance DayThur.	Nov.	11
Rogation Sunday	May	23	First Sunday in Advent	Nov.	28
Victoria Day Mon.	May	24	St. AndrewTues.	Nov.	30
Ascension Day Thur.	May	27	Christmas DaySat.	Dec.	25

ANNIVERSARIES AND FESTIVALS, 1976

SYMBOLS AND ABBREVIATIONS

SUN, MOON AND PLANETS (The Moon generally

- \odot The Sun
- New Moon
- Full Moon
- ♀ Venus
- First Quarter
 Last Quarter
- 🕀 Earth
 - o[¬] Mars

₿ Mercury

SIGNS OF THE ZODIAC

Υ	Aries 0°	Ω Leo.	120°	\mathbf{x}	Sagittarius240°
Я	Taurus	M Virge	o150°	て	Capricornus 270°
Д	Gemini 60°	≏ Libra	a180°	***	Aquarius300°
69	Cancer	m Scor	pius 210°	Ж	Pisces

THE GREEK ALPHABET

Α, α	Alpha	I, i Iota	P,ρ Rho
Β, β	Beta	К, к Карра	Σ, σ Sigma
Γ, γ	Gamma	Λ, λ Lambda	T,τ Tau
Δ, δ	Delta	Μ, μ Μu	τ, υ Upsilon
Ε, ε	Epsilon	N, v Nu	Φ, φ Phi
Ζ, ζ	Zeta	Ξ,ξ Χί	X, χ Chi
Η, η	Eta	O, o Omicron	Ψ, ψ Psi
Θ, θ,	9 Theta	Π, π Ρί	Ω, ω Omega

CO-ORDINATE SYSTEMS AND TERMINOLOGY

Astronomical positions are usually measured in a system based on the *celestial* poles and celestial equator, the intersections of the earth's rotation axis and equatorial plane, respectively, and the infinite sphere of the sky. *Right ascension* (R.A. or α) is measured in hours (h), minutes (m) and seconds (s) of time, eastward along the celestial equator from the vernal equinox. Declination (Dec. or δ) is measured in degrees (°), minutes (′) and seconds (′) of arc, northward (N or +) or southward (S or -) from the celestial equator toward the N or S celestial pole. One hour of time equals 15 degrees.

Positions can also be measured in a system based on the *ecliptic*, the intersection of the earth's orbit plane and the infinite sphere of the sky. The sun appears to move eastward along the ecliptic during the year. *Longitude* is measured eastward along the ecliptic from the vernal equinox; *latitude* is measured at right angles to the ecliptic, northward or southward toward the N or S ecliptic pole. The vernal equinox is one of the two intersections of the ecliptic and the celestial equator; it is the one at which the sun crosses the celestial equator moving from south to north.

Objects are *in conjunction* if they have the same longitude or R.A., and are *in opposition* if they have longitudes or R.A.'s which differ by 180° . If the second object is not specified, it is assumed to be the sun. For instance, if a planet is "in conjunction", it has the same longitude as the sun. At *superior conjunction*, the planet is more distant than the sun; at *inferior conjunction*, it is nearer.

If an object crosses the ecliptic moving northward, it is at the *ascending node* of its orbit; if it crosses the ecliptic moving southward, it is at the *descending node*.

Elongation is the difference in longitude between an object and a second object (usually the sun). At conjunction, the elongation of a planet is thus zero.

- 24 Jupiter
- b Saturn
- O Uranus
- Ψ Neptune
- P Pluto

THE CONSTELLATIONS

LATIN NAMES WITH PRONUNCIATIONS AND ABBREVIATIONS

Andromeda,

× 1 × /= 1	A 1	4 1
an-drom e-da	. And	Andr
Antila, ant 11- <i>a</i>	. Ant	Anti
Apus, a pus	. Aps	Apus
Aquarius, a-kwar 1-us	. Aqr	Aqar
Aquila, ak'wi-i a	. Aqi	Aqii
Ara, a r <i>a</i>	. Ara	Arae
Aries, a ri-ez	. Arı	Arie
Auriga, o-ri ga	. Aur	Auri
Bootes, bo-o'tez	. Boo	Boot
Caelum, se lum	. Cae	Cael
Camelopardalis,	~	~ 1
ka-mel o-par da-lis	. Cam	Caml
Cancer, kan ser	. Cnc	Canc
Canes Venatici,	~ .	~
ka'nez ve-nat'i-si	.CVn	CVen
Canis Major,	~	~
kā'nīs mā'jēr	. CMa	СМај
Canis Minor,		
kā'nīs' mī'nēr	. CMi	CMin
Capricornus,		
kăp'rĭ-kôr'n <i>ŭ</i> s	. Cap	Capr
Carina, ka-rī'na	. Car	Cari
Cassiopeia, kăs'ĭ-ō-pē'ya'.	. Cas	Cas
Centaurus, sĕn-tô'rŭs	. Cen	Cent
Cepheus, sē'fūs	Cep	Ceph
Cetus, sē't <i>ŭ</i> s	. Cet	Ceti
Chamaeleon, ka-mē'lē-ŭn.	.Cha	Cham
Circinus, sûr'sĭ-n <i>ŭ</i> s	. Cir	Circ
Columba, kō-lŭm'ba	. Col	Colm
Coma Berenices,		
kō'ma bĕr'ē-nī'sēz	Com	Coma
Corona, Australis,		
kō-rō'na ôs-trā'lĭs	CrA	CorA
Corona Borealis,		
ka-rō na bō'rē-ā'lĭs	CrB	CorB
Corvus, kôr′v <i>ŭ</i> s	Crv	Corv
Crater, krā'tēr	Crt	Crat
Crux, krŭks	Cru	Cruc
Cygnus, sĭg′n <i>ŭ</i> s	Cyg	Cygn
Delphinus, děl-fi'n <i>ü</i> s	Del	Dlph
Dorado, dō-rā'dō	Dor	Dora
Draco, drā'kō	Dra	Drac
Equuleus, ē-kwoo'lē- <i>ū</i> s	Equ	Eaul
Eridanus, ē-rīd'a-nūs	Erî	Erid
Fornax, fôr'năks	For	Forn
Gemini, jĕm'ĭ-nī	Gem	Gemi
Grus, grūš	Gru	Grus
Hercules, hûr'kū'lēz	Her	Herc
Horologium,		
hŏr′ō-lô′jĭ-ŭm	Hor	Horo
Hydra, hi'dra	Hya	Hyda
Hydrus, hī'dr <i>ū</i> s	Hyi	Hydi

Indus in'düs	Ind	Indi
Lacerta $l_{a-s\hat{u}r'ta}$	Lac	Lacr
Lao 15'ō	Lac	Laci
	I MC	LEUII
Leo Millor, le o III her		
Lepus, le p u s	. Lep	Leps
Libra, li bra	. Lib	Libr
Lupus, lū′p <i>ū</i> s	. Lup	Lupi
Lynx, lĭngks	. Lyn	Lync
Lyra, lī'ra	. Lyr	Lyra
Mensa, měn'sa,	. Men	Mens
Microscopium		
mī'krō-skō'nĭ- <i>ŭ</i> m	Mic	Micr
Monoceros m-ōnŏs'êr-ás	Mon	Mono
Musoo mus/ka	Mue	Muso
Normo pôr/ma	Nor	Norm
	. INOF	Norm
Octans, or $tanz$. Oct	Octn
Ophiuchus, of 'i-ūk <i>u</i> s	. Oph	Ophi
Orion, ō-rī' <i>ŏ</i> n	. Ori	Orio
Pavo, Pā'vō	. Pav	Pavo
Pegasus, pěg'a-sŭs	. Peg	Pegs
Perseus, pûr sūs	. Per	Pers
Phoenix, fē'nīks	Phe	Phoe
Pictor nik'têr	Pic	Pict
Pisces nis 'āz	Pec	Pisc
Discis Austrinus	.1 50	1 150
riscis Austrilius,	Do A	DecA
pis is os-tri n u s	. PSA	PSCA
Puppis, pup is	. Pup	Pupp
Pyxis, pik'sis	. Pyx	Pyxi
Reticulum,	•	
rē-tĭk′ū-l <i>ŭ</i> m	. Ret	Reti
Sagitta, sa-jĭt'a	. Sge	Sgte
Sagittarius, săj'ĭ-tā'rĭ-ŭs.	. Sgr	Sgtr
Scorpius, skôr pĭ-ŭs	. Sco	Scor
Sculptor, skulp'ter	Scl	Scul
Scutum skū't <i>u</i> m	Sct	Scut
Serpens sûr/něnz	Ser	Sern
Serpens, sur penz	Sov	Sovt
Tourne tâ/ně	. SCA	Tour
	. rau	Taur
relescopium,	T .1	T-1 -
tel'e-sko'pi-um	. Iel	I ele
Triangulum,		
trī-ăng'gū-l <i>ŭ</i> m	. Tri	Tria
Triangulum Australe,		
trī-ăng'gū-l <i>ŭ</i> m ôs-trā'lē.	. Tra	TrAu
Tucana, tū-kā'na	. Tuc	Tucn
Ursa Major.		
ûr'sa mā'iēr	UMa	UMai
Ursa Minor		Jinaj
ûr'sa mi'nêr	IIM	IIMin
Vala $v\bar{a}'la$	Vel	Velr
Vince	. v Cl 	Vina
Virgo, vur go	. VIC	Virg
volans, vo lanz	. VOI	voin
Vulpecula, vůl-pěk'ū-la	. Vul	Vulp

ā fāte; ā chāotic; ă tăp; ă finăl; à åsk; a idea; â câre; ä älms; au aught; ē bē; e crēate; ĕ ěnd; ě angěl; ẽ makêr; ī tīme; ĭ bǐt; ĭ anĭmal; ō nōte; ō anatōmy; ŏ hŏt; ă ăccur; ô ôrb; ōō mōōn; oo book; ou out; ū tūbe; ū unite; ŭ sŭn; ŭ sŭbmit; û hûrl.

MISCELLANEOUS ASTRONOMICAL DATA

```
UNITS OF LENGTH
     1 Angstrom unit = 10^{-8} cm
                                                                       1 micrometre, \mu = 10^{-4} cm = 10^{4}A.
                                                                       1 \text{ cm} = 10 \text{ mm} = 0.39370 \dots \text{ in}
     1 inch
                           = exactly 2.54 centimetres
                                                                       1 \text{ m} = 10^2 \text{ cm} = 1.0936 \dots \text{ yd}
     1 vard
                           = exactly 0.9144 metre
                                                                      1 \text{ km} = 10^5 \text{ cm} = 0.62137 \dots \text{ mi}
                           = exactly 1.609344 kilometres
     1 mile
     1 astronomical unit = 1.4960 \times 10^{13} cm = 1.496 \times 10<sup>8</sup> km = 9.2956 \times 10<sup>7</sup> mi
     1 light-year = 9.461 \times 10^{17} cm = 5.88 \times 10^{12} mi = 0.3068 parsecs
                          = 3.086 \times 10^{18} cm = 1.917 \times 10^{13} mi = 3.262 l.y.
     1 parsec
     1 megaparsec = 10<sup>6</sup> parsecs
UNITS OF TIME
                                = 23h 56m 04.09s of mean solar time
    Sidereal day
    Mean solar day
Synodic month
                              = 24h 03m 56.56s of mean sidereal time
                                = 29d \ 12h \ 44m \ 03s = 29^{d} \ 5306
                                                                         Sidereal month = 27d \ 07h \ 43m \ 12s
     Tropical year (ordinary) = 365d \ 05h \ 48m \ 46s = 365!2422
                                                                                           = 2743216
                               = 365d \ 06h \ 09m \ 10s = 365d \ 2564
     Sidereal year
    Eclipse year
                                = 346d \ 14h \ 52m \ 52s \ = \ 34646200
THE EARTH
    Equatorial radius, a = 6378.164 km = 3963.21 mi: flattening, c = (a - b)/a = 1/298.25
     Polar radius. b = 6356.779 \text{ km} = 3949.92 \text{ mi}
                                  = 111.133 - 0.559 cos 2\phi km = 69.055 - 0.347 cos 2\phi mi (at lat. \phi)
     1° of latitude
                                   = 111.413 \cos \phi - 0.094 \cos 3\phi \, \text{km} = 69.229 \cos \phi - 0.0584 \cos 3\phi \, \text{mi}
     1° of longitude
     Mass of earth
                                  = 5.976 \times 10^{24} \text{ kgm} = 13.17 \times 10^{24} \text{ lb}
     Velocity of escape from \oplus = 11.2 \text{ km/sec} = 6.94 \text{ mi/sec}
EARTH'S ORBITAL MOTION
     Solar parallax = 8^{\prime\prime}.794 (adopted)
     Constant of aberration = 20^{\prime\prime}.496 (adopted)
     Annual general precession = 50^{\circ}.26; obliquity of ecliptic = 23^{\circ} 26^{\circ} 35^{\circ} (1970)
     Orbital velocity = 29.8 km/sec = 18.5 mi/sec
     Parabolic velocity at \oplus = 42.3 \text{ km/sec} = 26.2 \text{ mi/sec}
SOLAR MOTION
     Solar apex, R.A. 18h 04m, Dec. + 30°; solar velocity = 19.75 km/sec = 12.27 mi/sec
THE GALACTIC SYSTEM
     North pole of galactic plane R.A. 12h 49m, Dec. + 27.^{\circ}4 (1950)
     Centre of galaxy R.A. 17h 42.4m, Dec. - 28° 55' (1950) (zero pt. for new gal. coord.)
     Distance to centre \sim 10,000 parsecs; diameter \sim 30,000 parsecs
     Rotational velocity (at sun) \sim 250 km/sec
     Rotational period (at sun) \sim 2.46 \times 10^8 years
     Mass \sim 1.4 \times 10^{11} solar masses
EXTERNAL GALAXIES
     Red Shift \sim +75 km/sec/megaparsec \sim 14 miles/sec/million l.y.
RADIATION CONSTANTS
     Velocity of light, c = 2.997925 \times 10^{10} \text{ cm/sec} = 186,282.1 \text{ mi/sec}
     Frequency, v = c/\lambda; v in Hertz (cycles per sec), c in cm/sec, \lambda in cm
     Solar constant = 1.950 gram calories/square cm/minute = 1.36 \times 10^6 cgs units
     Light ratio for one magnitude = 2.512 ...; log ratio = exactly 0.4
     Stefan's constant = 5.66956 \times 10^{-5} cgs units
MISCELLANEOUS
     Constant of gravitation, G = 6.6727 \times 10^{-8} cgs units
     Mass of the electron, m = 9.1096 \times 10^{-28} gm; mass of the proton = 1.6727 \times 10^{-24} gm
     Planck's constant, h = 6.6262 \times 10^{-27} erg sec
     Absolute temperature = T^{\circ} K = T^{\circ} C + 273^{\circ} = 5/9 (T^{\circ} F + 459^{\circ})
     1 \text{ radian} = 57^{\circ}.2958
                                       \pi = 3.141,592,653,6
                  = 3437'.75
                                         No, of square degrees in the sky = 41,253
                  = 206.265''
                                         1 \text{ gram} = 0.03527 \text{ oz}
```

SUN-EPHEMERIS AND CORRECTION TO SUN-DIAL

Date	Apparent R.A. 0h E.T.	Apparent Dec. 0h E.T.	Corr. to Sun-dial 12h E.T	Date	Apparent R.A. 0h E.T.	Apparent Dec. 0h E.T.	Corr. to Sun-dial 12h E.T
Jan. 1 4 7 10 13 16 19 22 25 28 31	h m s 18 42 10 18 55 25 19 08 36 19 21 43 19 34 44 19 47 40 20 03 00 20 13 14 20 25 51 20 38 21 20 50 44	$\begin{array}{c} \circ & -23 & 05.2 \\ -22 & 49.8 \\ -22 & 30.2 \\ -22 & 06.8 \\ -21 & 39.4 \\ -21 & 08.2 \\ -20 & 33.4 \\ -19 & 55.1 \\ -19 & 13.4 \\ -18 & 28.6 \\ -17 & 40.7 \\ \end{array}$	$ \begin{array}{c} m & s \\ + & 3 & 16 \\ + & 4 & 40 \\ + & 6 & 01 \\ + & 7 & 17 \\ + & 8 & 29 \\ + & 9 & 34 \\ + & 10 & 33 \\ + & 11 & 26 \\ + & 12 & 12 \\ + & 12 & 51 \\ + & 13 & 23 \end{array} $	July 2 5 8 11 14 17 20 23 26 29	h m s 6 44 33 6 56 55 7 09 15 7 21 30 7 33 42 7 45 49 7 57 52 8 09 50 8 21 43 8 33 30	$\begin{array}{c} \circ & \cdot \\ +23 & 02.7 \\ +22 & 47.7 \\ +22 & 29.2 \\ +22 & 07.1 \\ +21 & 41.7 \\ +21 & 12.9 \\ +20 & 40.8 \\ +20 & 05.7 \\ +19 & 27.5 \\ +18 & 46.4 \end{array}$	
Feb. 3 6 9 12 15 18 21 24 27	21 02 59 21 15 07 21 27 08 21 39 01 21 50 48 22 02 27 22 14 01 22 25 29 22 36 51	-16 50.0 -15 56.7 -15 00.9 -14 02.8 -13 02.7 -12 00.7 -10 56.9 - 9 51.5 - 8 44.9	$\begin{array}{c} + 13 & 48 \\ + 14 & 05 \\ + 14 & 15 \\ + 14 & 15 \\ + 14 & 17 \\ + 14 & 13 \\ + 14 & 02 \\ + 13 & 45 \\ + 13 & 22 \\ + 12 & 54 \end{array}$	Aug. 1 4 7 10 13 16 19 22 25 28 31	8 45 12 8 56 48 9 08 19 9 19 45 9 31 05 9 42 20 9 53 31 10 04 37 10 15 40 10 26 39 10 37 34	$\begin{array}{c} +18 & 02.5 \\ +17 & 16.0 \\ +16 & 27.0 \\ +15 & 35.6 \\ +14 & 41.9 \\ +13 & 46.2 \\ +12 & 48.4 \\ +11 & 48.9 \\ +10 & 47.6 \\ + & 9 & 44.8 \\ + & 8 & 40.6 \end{array}$	$ \begin{array}{c} + 6 \\ + 6 \\ + 6 \\ + 5 \\ + 5 \\ + 5 \\ + 4 \\ + 4 \\ + 3 \\ + 2 \\ + 1 \\ + 2 \\ + 1 \\ + 1 \\ + 9 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 \\ + 0 $
Mar. 1 4 7 10 13 16 19 22 25 28 31	22 48 09 22 59 21 23 10 30 23 21 35 23 32 37 23 43 36 23 54 33 0 05 29 0 16 25 0 27 20 0 38 15	$\begin{array}{c} - 7 & 37.0 \\ - 6 & 28.2 \\ - 5 & 18.5 \\ - 4 & 08.3 \\ - 2 & 57.5 \\ - 1 & 46.5 \\ - 0 & 35.4 \\ + & 0 & 35.7 \\ + & 1 & 46.6 \\ + & 2 & 57.2 \\ + & 4 & 07.2 \end{array}$	$\begin{array}{r} +12 \ 21 \\ +11 \ 43 \\ +11 \ 01 \\ +10 \ 16 \\ +9 \ 28 \\ +8 \ 37 \\ +7 \ 44 \\ +6 \ 51 \\ +5 \ 56 \\ +5 \ 50 \\ +4 \ 08 \end{array}$	Sept. 3 6 9 12 15 18 21 24 27 30	10 48 27 10 59 17 11 10 05 11 20 51 11 31 37 11 42 23 11 53 09 12 03 55 12 14 43 12 25 33	$\begin{array}{r} + 7 & 35.2 \\ + & 6 & 28.7 \\ + & 5 & 21.2 \\ + & 4 & 12.9 \\ + & 3 & 03.9 \\ + & 1 & 54.5 \\ + & 0 & 44.6 \\ - & 0 & 25.5 \\ - & 1 & 35.7 \\ - & 2 & 45.7 \end{array}$	- 0 44 - 1 44 - 2 46 - 3 49 - 4 53 - 5 57 - 7 01 - 8 04 - 9 05 - 10 05
Apr. 3 9 12 15 18 21 24 27 30	0 49 12 1 00 09 1 11 08 1 22 10 1 33 13 1 44 20 1 55 31 2 06 46 2 18 05 2 29 29	$\begin{array}{r} + 5 & 16.6 \\ + 6 & 25.1 \\ + 7 & 32.6 \\ + 8 & 39.0 \\ + 9 & 44.0 \\ + 10 & 47.6 \\ + 11 & 49.6 \\ + 12 & 49.9 \\ + 13 & 48.3 \\ + 14 & 44.6 \end{array}$	$\begin{array}{r} + & 3 & 15 \\ + & 2 & 23 \\ + & 1 & 32 \\ + & 0 & 44 \\ - & 0 & 043 \\ - & 1 & 21 \\ - & 1 & 55 \\ - & 2 & 25 \\ - & 2 & 50 \end{array}$	Oct. 3 6 9 12 15 18 21 24 27 30	12 36 25 12 47 20 12 58 18 13 09 20 13 20 27 13 31 39 13 42 57 13 54 21 14 05 51 14 17 27	$\begin{array}{r} -3 55.5 \\ -5 04.8 \\ -6 13.6 \\ -7 21.7 \\ -8 28.8 \\ -9 34.9 \\ -10 39.8 \\ -11 43.2 \\ -12 45.0 \\ -13 45.1 \end{array}$	$\begin{array}{r} -11 & 02 \\ -11 & 56 \\ -12 & 47 \\ -13 & 34 \\ -14 & 16 \\ -14 & 52 \\ -15 & 23 \\ -15 & 48 \\ -16 & 07 \\ -16 & 19 \end{array}$
May 3 6 9 12 15 18 21 24 27 30	2 40 58 2 52 31 3 04 09 3 15 52 3 27 41 3 39 34 3 51 33 4 03 36 4 15 45 4 27 57	$\begin{array}{c} +15 & 38.8 \\ +16 & 30.6 \\ +17 & 19.9 \\ +18 & 06.7 \\ +18 & 50.6 \\ +19 & 31.8 \\ +20 & 09.9 \\ +20 & 45.0 \\ +21 & 16.8 \\ +21 & 45.3 \end{array}$	- 3 10 - 3 26 - 3 36 - 3 42 - 3 42 - 3 38 - 3 28 - 3 13 - 2 54 - 2 30	Nov. 2 5 8 11 14 17 20 23 26 29	14 29 10 14 41 01 14 52 59 15 05 04 15 17 17 15 29 38 15 42 07 15 54 43 16 07 26 16 20 16	$\begin{array}{c} -14 \ 43.1 \\ -15 \ 39.0 \\ -16 \ 32.6 \\ -17 \ 23.7 \\ -18 \ 12.1 \\ -18 \ 57.6 \\ -19 \ 40.1 \\ -20 \ 19.4 \\ -20 \ 55.4 \\ -21 \ 27.7 \end{array}$	$\begin{array}{r} -16 & 24 \\ -16 & 22 \\ -16 & 13 \\ -15 & 56 \\ -15 & 31 \\ -14 & 58 \\ -14 & 58 \\ -14 & 18 \\ -13 & 30 \\ -12 & 36 \\ -11 & 35 \end{array}$
June 2 5 8 11 14 17 20 23 26 29	4 40 14 4 52 33 5 04 56 5 17 21 5 29 47 5 42 15 5 54 44 6 07 12 6 19 41 6 32 08	$\begin{array}{r} +22 \ 10.5 \\ +22 \ 32.1 \\ +22 \ 50.2 \\ +23 \ 04.7 \\ +23 \ 15.5 \\ +23 \ 22.7 \\ +23 \ 25.8 \\ +23 \ 21.8 \\ +23 \ 14.1 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	Dec. 2 5 8 11 14 17 20 23 26 29	16 33 11 16 46 12 16 59 18 17 12 29 17 25 43 17 38 59 17 52 18 18 05 38 18 18 57 18 32 15	$\begin{array}{r} -21 \ 56.5 \\ -22 \ 21.4 \\ -22 \ 42.4 \\ -23 \ 59.4 \\ -23 \ 12.4 \\ -23 \ 12.4 \\ -23 \ 25.7 \\ -23 \ 26.0 \\ -23 \ 22.1 \\ -23 \ 14.0 \end{array}$	$\begin{array}{r} -10 & 28 \\ - & 9 & 16 \\ - & 7 & 59 \\ - & 6 & 38 \\ - & 5 & 13 \\ - & 3 & 45 \\ - & 2 & 16 \\ - & 0 & 46 \\ + & 0 & 43 \\ + & 2 & 12 \end{array}$

PRINCIPAL ELEMENTS OF THE SOLAR SYSTEM MEAN ORBITAL ELEMENTS

	Mean Distance from Sun (a)		Period Revoluti	Period of Revolution		- In- clina-	Long. of	Long. of Peri-	Mean Long. at	
		millions	Sidereal	Syn-	city	tion	Node	helion	Epoch	
Planet	A. U.	of km	(P)	odic	(e)	(i)	(83)	(π)	(L)	
				days		•	0	۰.	•	
Mercury	0.387	57.9	88.0d.	116	.206	7.0	47.9	76.8	222.6	
Venus	0.723	108.1	224.7	584	.007	3.4	76.3	131.0	174.3	
Earth	1.000	149.5	365.26		.017	0.0	0.0	102.3	100.2	
Mars	1.524	227.8	687.0	780	.093	1.8	49.2	335.3	258.8	
Jupiter	5.203	778.	11.86y.	399	.048	1.3	100.0	13.7	259.8	
Saturn	9.539	1427.	29.46	378	.056	2.5	113.3	92.3	280.7	
Uranus	19.18	2869.	84.01	370	.047	0.8	73.8	170.0	141.3	
Neptune	30.06	4497.	164.8	367	.009	1.8	131.3	44.3	216.9	
Pluto	39.44	5900.	247.7	367	.250	17.2	109.9	224.2	181.6	

These elements, for epoch 1960 Jan. 1.5 E.T., are taken from the *Explanatory* Supplement to the American Ephemeris and Nautical Almanac.

(Object	Equa- torial Di- ameter km	Ob- late- ness	$Mass \oplus = 1$	Mean Den- sity water =1	Sur- face Grav- ity $\oplus = 1$	Rotation Period	Incli- nation of Equa- tor to Orbit °	Albedo
0	Sun	1,392,000	0	332,958	1.41	27.9	25 ^d -35 ^d †		
Œ	Moon	3,476	0	0.0123	3.36	0.16	$27^{d} 07^{h} 43^{m}$	6.7	0.067
₿	Mercury	4,865	0	0.055	5.46	0.38	58 ^d 16 ^h	<7	0.056
Ŷ	Venus	12,110	0	0.815	5.23	0.90	243 ^d (retro.)	~179	0.76
\oplus	Earth	12,756	1/298	1.000	5.52	1.00	23 ^h 56 ^m 04 ^s	23.4	0.36
ď	Mars	6,788	1/192	0.107	3.93	0.38	24 37 23	24.0	0.16
24	Jupiter	143,000	1/16	318.0	1.33	2.64	9 50 30	3.1	0.73
þ	Saturn	121,000	1/10	95.2	0.69	1.13	10 14	26.7	0.76
ô	Uranus	47,000	1/16	14.6	1.56	1.07	10 49	97.9	0.93
Ψ	Neptune	50,900	1/50	17.3	1.54	1.08	16	28.8	0.62
Р	Pluto	5,500?	2	0.11	5?	0.6?	6 ^d 9 ^h 17 ^m	?	0.14?

PHYSICAL ELEMENTS

†Depending on latitude. For the physical observations of the sun, p. 58, the sidereal period of rotation is 25.38 m.s.d.

SATELLITES OF THE SOLAR STSTEN	SATELLITES	S OF THE	SOLAR	SYSTEM
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		Die	Mean D from P	istance lanet	Re [.] I	voluti Period	on	Orbit	
Name	Vis. Mag.	km	km/1000	arc sec	d	h	m	°	Discovery
SATELLITE OF T	HE EART	тн							
Moon	-12.7	3476	384.5		27	07	43	18–29	
SATELLITES OF	Mars								
Phobos	11.6	23	9.3	26	0	07	39	1.0	A. Hall, 1877
Deimos	12.7	13	23.5	63	1	06	18	1.3	A. Hall, 1877
SATELLITES OF	Jupiter								
V Amalthea	13.0	(200)	180	59	0	11	57	0.4	E. Barnard 1892
I Io	5.0	3640	422	138	1	18	28	0	Galileo, 1610
II Europa	5.3	3100	671	220	3	13	14	0	Galileo, 1610
III Ganymede	4.6	5270	1,070	351	7	03	43	0	Galileo, 1610
IV Callisto	5.6	4990	1,885	618	16	16	32	0	Galileo, 1610
XIII Leda	20	< 10	10,206	3350	210	14		28.8	C. Kowal, 1974
VI Himalia	14.7	(150)	11,470	3765	250	14		27.6	C. Perrine, 1904
VII Elara	16.0	(50)	11,740	3850	259	16		24.8	C. Perrine, 1905
X Lysithea	18.8	< 20	11,850	3888	263	13		29.0	S. Nicholson, 1938
XII Ananke	18.3	< 20	21,200	6958	631	02		147	S. Nicholson, 1951
XI Carme	18.6	< 20	22,560	7404	692	12		164	S. Nicholson, 1938
VIII Pasiphae	18.1	< 20	23,500	7715	738	22		145	P. Melotte, 1908
IX Sinope	18.8	< 20	23,700	7779	758			153	S. Nicholson, 1914
SATELLITES OF	Saturn								
Janus	14	(300)	160	26	0	17	59	0.0	A. Dollfus, 1966
Mimas	12.1	(400)	-187	30	0	22	37	1.5	W. Herschel, 1798
Enceladus	11.8	(500)	238	38	1	08	53	0.0	W. Herschel, 1789
Tethys	10.3	(950)	295	48	1	21	18	1.1	G. Cassini, 1684
Dione	10.4	1100	378	61	2	17	41	0.0	G. Cassini, 1684
Rhea	9.7	1600	526	85	4	12	25	0.4	G. Cassini, 1672
Titan	8.4	5800	1,221	197	15	22	41	0.3	C. Huygens, 1655
Hyperion	14.2	(320)	1,481	239	21	06	38	0.4	G. Bond, 1848
Iapetus	11.0v	1500	3,561	575	79	07	56	14.7	G. Cassini, 1671
Phoebe	16.5	(200)	12,960	2096	550	11		150	W. Pickering, 1898
SATELLITES OF	Uranus								
Miranda	16.5	(400)	128	9	1	09	56	0	G. Kuiper, 1948
Ariel	14.4	(1400)	192	14	2	12	29	0	W. Lassell, 1851
Umbriel	15.3	(1000)	267	20	4	03	38	0	W. Lassell, 1851
Titania	14.0	(1800)	438	33	8	16	56	0	W. Herschel, 1787
Oberon	14.2	(1600)	587	44	13	11	07	0	W. Herschel, 1787
SATELLITES OF 1	NEPTUNE	3							
Triton	13.6	4000	354	17	5	21	03	160.0	W. Lassell, 1846
Nereid	18.7	(600)	5600	264	359	10	_	27.4	G. Kuiper, 1949

Apparent magnitude and mean distance from planet are at mean opposition distance. The inclination of the orbit is referred to the planet's equator; a value greater than 90° indicates retrograde motion.

Apparent magnitudes and most of the diameters are from data presented at the IAU Colloquium on Planetary Satellites, Cornell University, August 1974, as printed in *Astronomy* magazine. I thank Mr. R. T. Dickinson for providing this data. The diameters of the smaller satellites are very uncertain, because they depend on assumptions about the albedo of the satellite.

TIME

Any recurring event may be used to measure time. The various times commonly used are defined by the daily passages of the sun or stars caused by the rotation of the earth on its axis. The more uniform revolution of the earth about the sun, causing the return of the seasons, defines ephemeris time. The atomic second has been defined; atomic time has been maintained in various labs, and an internationally acceptable atomic time scale is under discussion.

A sundial indicates *apparent solar time*, but this is far from uniform because of the earth's elliptical orbit and the inclination of the ecliptic. If the real sun is replaced by a fictitious mean sun moving uniformly in the equator, we have *mean* (solar) *time*. *Apparent time* – *mean time* = *equation of time*. This is the same as *correction to sundial* on page 7, with reversed sign.

If instead of the sun we use stars, we have *sidereal time*. The sidereal time is zero when the vernal equinox or first point of Aries is on the meridian. As the earth makes one more rotation with respect to the stars than it does with respect to the sun during a year, sidereal time gains on mean time 3^m 56^s per day or 2 hours per month. Right Ascension (R.A.) is measured east from the vernal equinox, so that the R.A. of a body on the meridian is equal to the sidereal time.

Sidereal time is equal to mean solar time plus 12 hours plus the R.A. of the fictitious mean sun, so that by observation of one kind of time we can calculate the other. Local Sidereal time may be found approximately from Standard or zone time (0 h at midnight) by applying the corrections for longitude (p. 14) and sundial (p. 7) to obtain apparent solar time, then adding 12 h and R.A. sun (p. 7). (Note that it is necessary to obtain R.A. of the sun and correction to sundial at the standard time involved.)

Local mean time varies continuously with longitude. The local mean time of Greenwich, now known as *Universal Time* (UT) is used as a common basis for timekeeping. Navigation and surveying tables are generally prepared in terms of UT. When great precision is required, UT1 and UT2 are used differing from UT by polar variation and by the combined effects of polar variation and annual fluctuation respectively.

To avoid the inconveniences to travellers of a changing local time, *standard time* is used. The earth is divided into 24 zones, each ideally 15 degrees wide, the zero zone being centered on the Greenwich meridian. All clocks within the same zone will read the same time.

In Canada and the United States there are 9 standard time zones as follows: Newfoundland (N), $3^h 30^m$ slower than Greenwich; 60th meridian or Atlantic (A), 4 hours; 75th meridian or Eastern (E), 5 hours; 90th meridian or Central (C), 6 hours; 105th meridian or Mountain (M), 7 hours; 120th meridian or Pacific (P), 8 hours; 135th meridian or Yukon (Y), 9 hours; 150th meridian or Alaska-Hawaii, 10 hours; and 165th meridian or Bering, 11 hours slower than Greenwich.

The mean solar second, defined as 1/86400 of the mean solar day, has been abandoned as the unit of time because random changes in the earth's rotation make it variable. The unit of time has been redefined twice within the past two decades. In 1956 it was defined in terms of Ephemeris Time (ET) as 1/31,556,925.9747 of the tropical year 1900 January 0 at 12 hrs. ET. In 1967 it was redefined as 9,192,631,770 periods of the radiation corresponding to the transition between the two hyperfine levels of the ground state of the cesium 133 atom. Ephemeris Time is required in

celestial mechanics, while the cesium resonator makes the unit readily available. The difference, ΔT , between UT and ET is measured as a small error in the observed longitude of the moon, in the sense $\Delta T = ET - UT$. The moon's position is tabulated in ET, but observed in UT. ΔT was zero near the beginning of the century, but in 1976 will be about 47 seconds.

RADIO TIME SIGNALS

National time services distribute co-ordinated time called UTC, which on January 1, 1972, was adjusted so that the time interval is the atomic second. The resulting atomic time gains on mean solar time at a rate of about a second a year. An approximation to UT1 is maintained by stepping the atomic time scale in units of 1 second on June 30 or December 31 when required so that the divergence from mean solar time (DUT1 = UT1 - UTC) does not exceed 0.6 second. The first such "leap second" occurred on June 30, 1972. These changes are coordinated through the Bureau International de l'Heure (BIH), so that most time services are synchronized to the tenth of a millisecond.

DUT1 is identified each minute on CHU and WWV by a special group of split or double pulses. The number of such marker pulses in a group gives the value of DUT1 in tenths of a second. If the group starts with the first (not zero) second of each minute, DUT1 is positive and mean solar time is ahead of the transmitted time; if with the 9th second DUT1 is negative, and mean solar time is behind.

Radio time signals readily available in Canada include:CHU Ottawa, Canada3330, 7335, 14670 kHzWWV Fort Collins, Colorado 2.5, 5, 10, 20, 25 MHzWWVH Maui, Hawaii2.5, 5, 10, 15 MHz.

JULIAN DAY CALENDAR, 1976

Jan.	1	. 2442779	May 1	 . 2442900	Sept.	1	2443023
Feb.	1	. 2442810	June 1	 . 2442931	Oct.	1	2443053
Mar.	1	. 2442839	July 1	 . 2442961	Nov.	1	2443084
Apr.	1	. 2442870	Aug. 1	 2442992	Dec.	1	2443114

The Julian day commences at noon, so that J.D. 2442779 = Jan. 1.5 U.T. 1976 = 12 hours U.T., Jan. 1, 1976.

The Julian date is commonly used by astronomers to refer to the time of astronomical events, because it avoids some of the annoying complexities of the civil calendar. The Julian day corresponding to a given date is the number of days which have elapsed since Jan. 1, 4713 B.C.

This system was introduced in 1582 by Josephus Justus Scaliger under the name of the Julian period. The Julian period lasts 7980 years, and is the least common multiple of three cycles: the solar cycle of 28 Julian years, the lunar (or Metonic) cycle of 19 Julian years, and the Roman indiction cycle of 15 years. On Jan. 1, 4713 B.C., all three cycles began together. For more information, see "The Julian Period", by C. H. Cleminshaw in the *Griffith Observer*, April 1975

MAP OF STANDARD TIME ZONES



PRODUCED BY THE SURVEYS AND MAPPING BRANCH, DEPARTMENT OF ENERGY, MINES AND RESOURCES, OTTAWA, CANADA, 1973.

Note: Since the preparation of the above map, the standard time zones have been changed so that all parts of the Yukon Territory now observe Pacific Standard Time. The Yukon Standard Time Zone still includes a small part of Alaska, as shown on the above map.

VISITING HOURS AT SOME CANADIAN OBSERVATORIES Compiled by Marie Litchinsky

Burke-Gaffney Observatory, Saint Mary's University, Halifax, Nova Scotia B3H 3C3. October-April: Saturday evenings 7:00 p.m.

May-September: Saturday evenings 9:00 p.m.

David Dunlap Observatory, Richmond Hill, Ontario L4C 4Y6.

Wednesday mornings throughout the year, 10:00 a.m.

Saturday evenings, April through October (by reservation, tel. 884-2112).

Dominion Astrophysical Observatory, Victoria, B.C. V8X 3X3.

May-August: Daily, 9:15 a.m.-4:15 p.m.

Sept.-April: Monday to Friday, 9:15 a.m.-4:15 p.m.

Public observing, Saturday evenings, April-October inclusive.

Dominion Radio Astrophysical Observatory, Penticton, B.C. V2A 6K3

Sunday, July and August only (2:00–5:00 p.m.).

PLANETARIUMS

- The Calgary Centennial Planetarium, Mewata Park, Calgary, Alberta T2P 2M5 Winter: Wed.-Fri., 7:15 and 8:45 p.m. Sat.-Sun., 1:45 (children), 3:00, 7:15 and 8:45 p.m. (Closed Christmas Day, New Year's Day and Good Friday.)
- Summer: Daily except Tues., 1:45 (children), 3:00, 4:15, 7:15 and 8:45 p.m.
- Dow Planetarium, 1000 St. Jacques St. W., Montreal, P.Q.
- In English: Tues.-Fri., 12:15 p.m.; Sat. 1:00 and 3:30 p.m.; Sun. 2:15 p.m. Evenings (except Mon.) 8:15 p.m.
- *In French:* Tues.-Sat., 2:15 p.m., also Sat. 4:30 p.m., Sun. 1:00, 3:30 and 4:30 p.m. Evenings (except Mon.) 9:30 p.m.
- H. R. MacMillan Planetarium, 1100 Chestnut Street, Vancouver, B.C. V6J 3J9.
 - Sept.-June: Tues.-Wed. 3:00 and 7:30 p.m.; Thurs. 7:30 p.m.; Fri. 7:30 and 9:00 p.m.; Weekends and Holidays, 1:00, 2:30, 4:00, 7:30 and 9:00 p.m.
 - July-August: Mornings (Tues. to Sat.) 11:30 a.m. Afternoons (Tues. to Sun.) 1:00, 2:30, 4:00 p.m.

Evenings (Mon. to Sun.) 7:30 and 9:00 p.m.

- Manitoba Museum of Man & Nature Planetarium, 190 Rupert Ave., Winnipeg, Man. R3B 0N2.
 - Sept.-June: Tues.-Fri. 3:15 and 8:00 p.m.; Sat. and holidays, 1:00, 2:30, 4:00, 7:30 and 9:00 p.m.; Sun. 1:00, 2:30 and 4:00 p.m. Closed Mondays except holidays.
 - July-August: Mon. 2:00 and 3:30 p.m. (except holidays); Tues.-Fri. 11:00 a.m., 2:00, 3:30, 7:30 and 9:00 p.m.; Sat., Sun. and holidays 1:00, 2:30, 4:00, 7:30 and 9:00 p.m.

McLaughlin Planetarium, 100 Queen's Park, Toronto, Ont. M5S 2C6.

Tues.-Sun. 1:30, 3:00 and 7:30 p.m. Holidays 1:30 and 3:00 p.m.

Theatre closed on Mondays, except on holidays.

- *McMaster University*, School of Adult Education, GH-122, Hamilton, Ont. Group reservations only.
- Queen Elizabeth Planetarium, Edmonton, Alberta. Winter: Tues.-Fri. 8:00 p.m., Sat. 3:00 p.m., Sun. and holidays 3:00 and 8:00 p.m. Summer: Mon.-Sat. 3:00 and 8:00 p.m., Sun, and holidays 3:00 and 8:00 p.m.
- Seneca College Planetarium, 1750 Finch Ave. East, Willowdale, Ont. M2N 5T7. Group reservations only.
- The University of Manitoba Planetarium, 394 University College, 500 Dysart Rd., Winnipeg, Man. R3T 2M8.

Telephone 474–9785 for times of public shows and for group reservations.

TIMES OF RISING AND SETTING OF THE SUN AND MOON

The times of sunrise and sunset for places in latitudes ranging from 30° to 54° are given on pages 15 to 20, and of twilight on page 21. The times of moonrise and moonset for the 5 h meridian are given on pages 22 to 27. The times are given in Local Mean Time, and in the table below are given corrections to change from Local Mean Time to Standard Time for the cities and towns named.

The tabulated values are computed for the sea horizon for the rising and setting of the upper limb of the sun and moon, and are corrected for refraction. Because variations from the sea horizon usually exist on land, the tabulated times can rarely be observed.

The Standard Times for Any Station

To derive the Standard Time of rising and setting phenomena for the places named, from the list below find the approximate latitude of the place and the correction in minutes which follows the name. Then find in the monthly table the Local Mean Time of the phenomenon for the proper latitude on the desired day. Finally apply the correction to get the Standard Time. The correction is the number of minutes of time that the place is west (plus) or east (minus) of the standard meridian. The corrections for places not listed may be obtained by converting the longitude found from an atlas into time ($360^\circ = 24$ h).

CA	CANADIAN CITIES AND TOWNS						AMERICAN CITIES		
	Lat.	Corr.		Lat.	Corr.			Lat.	Corr.
Athabasca Baker Lake Brantford Calgary Charlottetown Churchill Edmonton Fredericton Gander Glace Bay Granby Goose Bay Granby Guelph Halifax Hamilton Hull Kapuskasing Kingston Kitchener London Medicine Hat Montreal Moose Jaw Niagara Falls North Bay Ottawa Owen Sound Penticton	55° 64 50 31 46 59 54 45 54 45 45 45 45 45 45 45 45 45 45	$\begin{array}{r} +33M\\ +24C\\ +40E\\ +21E\\ +36M\\ +12A\\ +17E\\ +12A\\ +17E\\ +22A\\ +23M\\ +22A\\ +20E\\ +23E\\ +22E\\ +23E\\ +22E\\ +23E\\ +22E\\ +23E\\ +22E\\ +22E\\$	Peterborough Port Harrison Prince Albert Prince Rupert Quebec Regina St. Catharines St. Hyacinthe Saint John, N.B. St. John's, Nfid. Sarnia Saskatoon Sault Ste. Marie Shawinigan Sherbrooke Stratford Sudbury Sydney The Pas Timmins Toronto Three Rivers Thunder Bay Trail Truro Vancouver Victoria Whitehorse Windsor Windsor Windsor Winnipeg Yellowknife	44 59 53 47 47 47 47 43 47 45 43 47 47 43 47 43 47 43 47 43 47 45 43 47 45 43 47 45 43 47 45 43 47 45 43 47 54 43 46 43 47 54 54 54 54 54 54 54 54 54 54 54 54 54	+13E +13E +13E +41P -15E +17E +58C +17E +58C +17E +24A +01A +02E +37E -09E +24E +13A +24A +45C +13P +24E +13B +12P +13P +12P +13P +22C +38M		Atlanta Baltimore Birmingham Boston Buffalo Chicago Cincinnati Cleveland Dallas Denver Detroit Fairbanks Flagstaff Indianapolis Juneau Kansas City Los Angeles Louisville Memphis Miami Milwaukee Minneapolis New Orleans New York Omaha Philadelphia Phoenix Pitsburgh St. Louis San Francisco Seattle Washington	34° 39 33 42 43 42 43 42 43 42 43 40 42 55 40 42 55 40 42 55 40 42 43 35 26 53 40 42 43 43 43 43 43 43 43 42 43 40 42 43 40 42 55 40 43 40 42 40 40 40 40 40 40 40 40 40 40 40 40 40	+37E +06E -13C +15E +15E +15E +26E +27C 00M +32E +27C 00M +32E +27C 00C +38E +27C 00C +38E +27C 00C -04E +28E +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +28C +20C +28C +28C +20C +28C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +28C +20C +20C +20C +20C +20C +20C +20C +20

Example—Find the time of sunrise at Owen Sound, on February 12.

In the above list Owen Sound is under " 45° ", and the correction is +24 min. On page 15 the time of sunrise on February 12 for latitude 45° is 7.06; add 24 min. and we get 7.30 (Eastern Standard Time).

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Jan.	15 25 4	5 29 5 35 5 38	18 21 18 25 18 32	5 37 5 42 5 45	18 14 18 18 18 25	5 44 5 50 5 53	18 06 18 10 18 18	5 52 5 57 6 00	17 57 18 02 18 10	5 59 6 04 6 07	17 51 17 55 18 04

TWILIGHT-BEGINNING OF MORNING AND ENDING OF EVENING

The above table gives the local mean time of the beginning of morning twilight, and of the ending of evening twilight, for various latitudes. To obtain the corresponding standard time, the method used is the same as for correcting the sunrise and sunset tables, as described on page 12. The entry——in the above table indicates that at such dates and latitudes, twilight lasts all night. This table, taken from the American Ephemeris, is computed for *astronomical* twilight, i.e. for the time at which the sun is 108° from the zenith (or 18° below the horizon).

MOONRISE	AND	MOONSET,	1976;	LOCAL	MEAN	TIME

DATE	Latitu	de 30°	Latitu	de 35°	Latitu	de 40°	Latitu	de 45°	Latitu	de 50°	Latitu	de 54°
	Mo	oon	Mo	bon	Mo	bon	Ma	oon	Mo	oon	Mo	oon
	Rise	Set										
Jan.	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m
1	06 43	17 34	06 54	17 24	07 06	17 12	07 20	16 59	07 37	16 42	07 54	16 25
2	07 30	18 33	07 39	18 24	07 50	18 14	08 02	18 03	08 17	17 49	08 31	17 35
3	08 12	19 30	08 19	19 24	08 28	19 16	08 37	19 08	08 49	18 57	09 01	18 47
4	08 50	20 26	08 55	20 22	09 01	20 17	09 08	20 11	09 17	20 04	09 25	19 57
5	09 24	21 20	09 28	21 18	09 31	21 15	09 36	21 13	09 40	21 09	09 45	21 06
6 7 8 9 D 10	09 57 10 29 11 01 11 34 12 09	22 13 23 05 23 57 	09 58 10 28 10 58 11 29 12 02	22 13 23 07 00 01 00 56	09 59 10 27 10 54 11 23 11 54	22 13 23 10 00 06 01 03	10 01 10 25 10 50 11 16 11 45	22 13 23 13 00 12 01 12	10 02 10 24 10 45 11 08 11 34	22 13 23 16 00 19 01 22	10 04 10 22 10 40 11 00 11 23	22 13 23 19
11	12 48	01 43	12 39	01 51	12 29	02 00	12 17	02 11	12 03	02 25	11 49	02 38
12	13 31	02 37	13 21	02 47	13 09	02 58	12 55	03 11	12 39	03 27	12 22	03 43
13	14 19	03 32	14 08	03 43	13 55	03 55	13 40	04 10	13 22	04 28	13 04	04 46
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15	16 10	05 20	15 59	05 31	15 47	05 43	15 33	05 58	15 15	06 16	14 58	06 34
16 ♀	17 11	06 10	17 02	06 20	16 51	06 31	16 39	06 44	16 24	07 00	16 09	07 16
17	18 15	06 58	18 08	07 06	18 00	07 15	17 50	07 26	17 38	07 39	17 27	07 51
18	19 20	07 42	19 15	07 48	19 10	07 55	19 03	08 03	18 56	08 12	18 48	08 21
19	20 25	08 24	20 23	08 28	20 20	08 31	20 17	08 36	20 14	08 41	20 10	08 46
20	21 30	09 04	21 30	09 05	21 31	09 06	21 31	09 07	21 32	09 09	21 33	09 10
21 22 23 @ 24 25	22 34 23 38 00 42 01 45	09 44 10 24 11 06 11 51 12 39	22 37 23 44 00 50 01 55	09 42 10 20 11 00 11 42 12 29	22 41 23 50 00 59 02 06	09 40 10 15 10 52 11 32 12 17	22 45 23 58 01 10 02 19	09 38 10 09 10 43 11 21 12 04	22 50 00 07 01 23 02 35	09 35 10 03 10 33 11 07 11 47	22 55 00 16 01 35 02 51	09 33 09 56 10 23 10 53 11 30
26	02 46	13 31	02 57	13 20	03 09	13 07	03 24	12 52	03 42	12 34	04 00	12 15
27	03 43	14 26	03 54	14 15	04 07	14 02	04 23	13 47	04 41	13 28	05 00	13 10
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31 ()	06 47	18 15	06 53	18 10	07 00	18 04	07 08	17 57	07 18	17 48	07 28	17 40
Feb.	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m
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5	09 34	22 40	09 30	22 46	09 25	22 52	09 19	22 59	09 12	23 08	09 06	23 16
6 7 8 ♪ 9 10	10 08 10 45 11 25 12 09 12 58	23 33 00 26 01 19 02 13	10 02 10 37 11 15 11 59 12 47	23 40 00 35 01 30 02 24	09 55 10 28 11 04 11 46 12 35	23 48 	09 46 10 17 10 52 11 32 12 20	23 58 00 57 01 55 02 51	09 37 10 04 10 36 11 15 12 02	00 10 01 12 02 12 03 09	09 27 09 52 10 21 10 58 11 43	00 21 01 26 02 29 03 28
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15€€	18 06	06 16	18 03	06 21	17 59	06 26	17 54	06 32	17 48	06 40	17 43	06 47
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20	23 38	09 50	23 47	09 42	23 57	09 33		09 22		09 09		08 57
21 22 @ 23 24 25	00 40 01 38 02 33 03 22	10 38 11 29 12 23 13 19 14 16	00 50 01 50 02 43 03 32	10 28 11 18 12 12 13 08 14 06	01 02 02 02 02 56 03 43	10 17 11 06 11 59 12 56 13 55	00 09 01 16 02 17 03 11 03 57	10 04 10 51 11 44 12 42 13 42	00 24 01 34 02 36 03 29 04 13	09 48 10 33 11 26 12 24 13 26	00 39 01 51 02 54 03 47 04 29	09 33 10 16 11 07 12 06 13 11
26	04 06	15 12	04 15	15 04	04 24	14 55	04 36	14 45	04 50	14 32	05 03	14 19
27	04 46	16 08	04 53	16 02	05 01	15 55	05 10	15 47	05 21	15 37	05 31	15 28
28	05 23	17 03	05 28	16 59	05 33	16 54	05 40	16 49	05 47	16 43	05 55	16 36
29 🍘	05 57	17 56	06 00	17 54	06 03	17 52	06 07	17 50	06 11	17 47	06 15	17 44

DATE	Latitude 30°		Latitu	atitude 35° Latitude 40°		Latitude 45°		Latitude 50°		Latitude 54°		
	Moon		Mo	Moon Moon		Moon		Moon		Moon		
	Rise Set		Rise	ise Set Rise Set		Rise Set		Rise Set		Rise Set		
Mar.	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m
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3	07 35	20 33	07 31	20 38	07 27	20 43	07 23	20 49	07 17	20 56	07 12	21 03
4	08 08	21 25	08 03	21 32	07 57	21 39	07 50	21 48	07 41	21 58	07 33	22 09
5	08 44	22 18	08 37	22 26	08 28	22 35	08 19	22 47	08 07	23 00	07 56	23 13
6 7 8∌ 9 10	09 22 10 04 10 50 11 41 12 36	23 10 00 03 00 55 01 45	09 13 09 54 10 40 11 30 12 26	23 20 00 14 01 06 01 56	09 03 09 43 10 27 11 18 12 14	23 31 00 26 01 18 02 08	08 52 09 29 10 13 11 03 12 00	23 44 00 40 01 33 02 22	08 38 09 13 09 55 10 45 11 43	00 00 00 57 01 51 02 40	08 24 08 57 09 37 10 27 11 26	00 16 01 15 02 09 02 57
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16 17 18 19 20	19 05 20 14 21 22 22 28 23 30	06 12 06 56 07 42 08 31 09 23	19 09 20 20 21 30 22 38 23 41	06 10 06 51 07 35 08 22 09 12	19 13 20 27 21 40 22 49 23 53	06 08 06 46 07 27 08 11 09 00	19 17 20 35 21 51 23 03	06 06 06 40 07 18 08 00 08 47	19 22 20 44 22 04 23 19 	06 03 06 33 07 07 07 45 08 30	19 27 20 54 22 17 23 35	06 00 06 27 06 56 07 31 08 13
21 22 @ 23 24 25	00 27 01 19 02 05 02 46	10 17 11 14 12 11 13 08 14 04	00 38 01 29 02 14 02 54	10 06 11 03 12 01 13 00 13 57	00 51 01 41 02 24 03 02	09 54 10 51 11 50 12 50 13 50	00 08 01 06 01 55 02 36 03 12	09 39 10 36 11 37 12 39 13 41	00 26 01 24 02 12 02 51 03 23	09 21 10 18 11 21 12 25 13 30	00 44 01 42 02 28 03 05 03 35	09 03 10 01 11 04 12 12 13 20
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Apr. 1 2 3 4 5	h m 06 45 07 22 08 03 08 47 09 35	h m 20 13 21 05 21 57 22 49 23 39	h m 06 38 07 14 07 53 08 37 09 25	h m 20 20 21 14 22 08 23 00 23 50	h m 06 30 07 04 07 42 08 25 09 12	h m 20 29 21 25 22 19 23 12	h m 06 22 06 53 07 29 08 11 08 58	h m 20 39 21 37 22 33 23 27	h m 06 11 06 40 07 14 07 53 08 40	h m 20 52 21 52 22 50 23 44	h m 06 01 06 27 06 59 07 36 08 22	h m 21 04 22 07 23 07 00 02
6 7 8 9 10	10 28 11 24 12 23 13 24 14 27	00 27 01 13 01 56 02 38	10 17 11 14 12 15 13 18 14 24	00 37 01 21 02 03 02 43	10 05 11 03 12 05 13 11 14 20	00 02 00 48 01 31 02 11 02 48	09 51 10 50 11 55 13 03 14 15	00 16 01 02 01 43 02 20 02 54	09 34 10 35 11 42 12 54 14 10	00 34 01 18 01 56 02 31 03 01	09 17 10 19 11 29 12 45 14 04	00 51 01 34 02 10 02 41 03 08
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16 17 18 19 20	21 13 22 15 23 11 00 01	07 09 08 05 09 03 10 02 11 01	21 24 22 26 23 22 00 10	06 59 07 54 08 52 09 52 10 52	21 36 22 39 23 34 00 21	06 48 07 42 08 39 09 40 10 42	21 50 22 53 23 48 00 34	06 35 07 27 08 25 09 26 10 30	22 07 23 12 00 05 00 49	06 19 07 09 08 07 09 09 10 15	22 25 23 30 00 22 01 04	06 04 06 52 07 49 08 52 10 01
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29	05 22	19 01	05 14	19 09	05 06	19 19	04 55	19 31	04 43	19 45	04 31	19 59
30	06 02	19 53	05 53	20 03	05 42	20 15	05 30	20 28	05 15	20 44	05 01	21 00

DATE	Latitu	ide 30°	Latitu	ide 35°	Latitu	ide 40°	Latitu	ude 45°	Latitu	ide 50°	Latitu	ide 54°
	Mo	oon	Mo	oon	Mo	bon	Mo	bon	Mo	oon	Mo	Don
	Rise	Set	Rise	Set	Rise	Set	Rise	Set	Rise	Set	Rise	Set
May 1 2 3 4 5	h m 06 45 07 33 08 24 09 18 10 15	h m 20 45 21 36 22 24 23 10 23 54	h m 06 35 07 22 08 13 09 08 10 06	h m 20 56 21 47 22 35 23 19 	h m 06 23 07 10 08 01 08 57 09 56	h m 21 08 21 59 22 46 23 30	h m 06 10 06 55 07 46 08 43 09 45	h m 21 23 22 13 23 00 23 42 	h m 05 53 06 37 07 29 08 27 09 31	h m 21 40 22 31 23 17 23 56 	h m 05 36 06 20 07 11 08 11 09 18	h m 21 58 22 49 23 33 00 11
6 7 ♪ 8 9 10	11 14 12 14 13 17 14 20 15 26	00 35 01 15 01 54 02 35	11 07 12 10 13 14 14 21 15 29	00 01 00 40 01 18 01 55 02 33	10 59 12 05 13 12 14 22 15 33	00 09 00 46 01 22 01 56 02 31	10 50 11 59 13 09 14 22 15 37	00 19 00 54 01 26 01 57 02 28	10 40 11 52 13 06 14 23 15 42	00 31 01 02 01 31 01 58 02 26	10 29 11 45 13 03 14 24 15 47	00 43 01 10 01 35 01 59 02 23
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16 17 18 19 20 C	21 51 22 39 23 21 23 59 	07 46 08 47 09 46 10 44 11 40	22 01 22 47 23 28 	07 35 08 37 09 39 10 38 11 36	22 12 22 57 23 36 00 09	07 23 08 26 09 29 10 31 11 32	22 26 23 09 23 45 00 16	07 08 08 13 09 19 10 23 11 26	22 42 23 22 23 55 00 23	06 51 07 57 09 06 10 14 11 20	22 58 23 36 00 06 00 31	06 33 07 42 08 53 10 04 11 14
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June 1 2 3 4 5 €	h m 08 10 09 08 10 08 11 08 12 10	h m 21 54 22 35 23 15 23 54	h m 08 01 09 01 10 02 11 05 12 09	h m 22 02 22 41 23 19 23 55	h m 07 50 08 52 09 57 11 02 12 09	h m 22 11 22 48 23 23 23 57	h m 07 38 08 43 09 50 10 58 12 08	h m 22 21 22 56 23 29 23 59	h m 07 24 08 31 09 41 10 54 12 08	h m 22 34 23 06 23 35 	h m 07 09 08 19 09 33 10 49 12 07	h m 22 47 23 15 23 41
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11 😨	18 39	04 27	18 50	04 17	19 02	04 04	19 17	03 50	19 35	03 32	19 53	03 15
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30	08 02	21 16	07 56	21 21	07 49	21 26	07 41	21 33	07 31	21 40	07 22	21 47

DATE	Latitu	de 30°	Latitu	de 35°	Latitu	de 40°	Latitu	de 45°	Latitu	de 50°	Latitu	ide 54°
	Mo	oon	Mo	oon	Mo	oon	Mc	oon	Mo	oon	Mo	oon
	Rise	Set	Rise	Set	Rise	Set	Rise	Set	Rise	Set	Rise	Set
July	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m
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16 17 18 19 @ 20	22 12 22 45 23 20 23 56	10 01 10 54 11 46 12 38 13 31	22 10 22 42 23 14 23 48 	10 01 10 56 11 51 12 45 13 39	22 09 22 37 23 08 23 40	10 02 11 00 11 56 12 53 13 49	22 06 22 33 23 00 23 30 	10 03 11 03 12 03 13 02 14 00	22 04 22 27 22 51 23 19 23 50	10 04 11 08 12 10 13 12 14 13	22 01 22 21 22 43 23 07 23 36	10 05 11 12 12 18 13 23 14 27
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Aug. 1 2 D 3 4 5	h m 11 06 12 10 13 14 14 16 15 15	h m 22 37 23 23 00 13 01 06	h m 11 12 12 18 13 23 14 26 15 25	h m 22 31 23 15 00 03 00 55	h m 11 18 12 26 13 34 14 38 15 38	h m 22 24 23 05 23 52 00 43	h m 11 25 12 37 13 46 14 52 15 52	h m 22 16 22 54 23 38 00 29	h m 11 33 12 49 14 02 15 10 16 10	h m 22 06 22 41 23 22 	h m 11 42 13 01 14 17 15 27 16 28	h m 21 56 22 28 23 07 23 54
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9€	18 25	05 01	18 30	04 54	18 37	04 47	18 44	04 37	18 53	04 26	19 02	04 15
10	19 02	05 59	19 06	05 54	19 10	05 49	19 15	05 43	19 20	05 35	19 25	05 28
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26	06 45	19 11	06 44	19 11	06 43	19 10	06 42	19 09	06 41	19 09	06 40	19 08
27	07 50	19 53	07 52	19 50	07 54	19 46	07 57	19 42	08 00	19 38	08 03	19 33
28	08 56	20 36	09 00	20 31	09 05	20 24	09 11	20 17	09 18	20 09	09 25	20 00
29	10 02	21 22	10 08	21 14	10 16	21 05	10 25	20 55	10 36	20 43	10 47	20 32
30	11 07	22 11	11 15	22 01	11 25	21 51	11 37	21 38	11 51	21 23	12 05	21 09
31	12 10	23 03	12 20	22 53	12 31	22 41	12 45	22 27	13 01	22 10	13 18	21 53

DATE	Latitude 30°		Latitude 35°		Latitude 40°		Latitude 45°		Latitude 50°		Latitude 54°	
	Moon		Moon		Moon		Moon		Moon		Moon	
	Rise Set		Rise Set		Rise Set		Rise Set		Rise Set		Rise Set	
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26 27 28 29 € 30 31	09 49 10 46 11 36 12 22 13 02 13 39	20 40 21 41 22 41 23 40 	09 59 10 56 11 45 12 29 13 08 13 42	20 30 21 31 22 33 23 33 	10 11 11 07 11 56 12 37 13 14 13 46	20 18 21 20 22 23 23 26 00 27	10 25 11 21 12 08 12 47 13 21 13 51	20 03 21 07 22 12 23 17 00 21	10 43 11 37 12 22 12 59 13 30 13 57	19 46 20 51 21 58 23 06 00 13	11 00 11 54 12 37 13 10 13 38 14 02	19 29 20 35 21 45 22 56

DATE	Latitude 30°		Latitude 35°		Latitude 40°		Latitude 45°		Latitude 50°		Latitude 54°	
	Moon		Moon		Moon		Moon		Moon		Moon	
	Rise Set		Rise Set		Rise Set		Rise Set		Rise Set		Rise Set	
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3	15 20	03 18	15 18	03 20	15 15	03 22	15 11	03 24	15 07	03 27	15 04	03 29
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18	02 57	14 49	03 00	14 45	03 04	14 40	03 08	14 35	03 13	14 28	03 18	14 22
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THE PLANETS FOR 1976

MERCURY

Mercury is the smallest planet and the closest to the sun. Like our moon, it has very little atmosphere, and its surface is covered with impact craters. However, Mercury lacks the vast plains or *maria* which are so conspicuous on the moon. The orbit of Mercury is well within that of the earth, and the planet appears to move quickly (hence its name) from one side of the sun to the other, several times in the year. Its greatest elongation (angular distance from the sun) varies from 18° to 28°, and on such occasions it can be seen by the unaided eye for about two weeks. Despite its considerable brilliance it is always viewed in the twilight sky and one must look carefully to see it.

The following table lists the greatest elongations east (evening sky) and west (morning sky) during the year. Only those marked * are particularly favourable.

J. T.)	Elong. East	Mag.	Date (U.T.)	Elong. West	Mag.	
7,05 ^h	19°	-0.3	Feb. 16, 15 ^h	26°	+0.2	
28, 02 ^h	21°	+0.4	June 15,09 ^h	23°	+0.8	
26, 10 ^h 20, 10 ^h	27° 20°	+0.5 -0.3	*Oct. 7, 16 ^h	18°	-0.1	
	J.T.) 7, 05 ^h 28, 02 ^h 26, 10 ^h 20, 10 ^h	J.T.) Elong. East 7,05 ^h 19° 28,02 ^h 21° 26,10 ^h 27° 20,10 ^h 20°	J.T.)Elong. EastMag. $7,05^{h}$ 19° -0.3 $28,02^{h}$ 21° $+0.4$ $26,10^{h}$ 27° $+0.5$ $20,10^{h}$ 20° -0.3	J.T.)Elong. EastMag.Date (U.T.) $7,05^{h}$ 19° -0.3 Feb. 16, 15^{h} $28,02^{h}$ 21° $+0.4$ June 15, 09^{h} $26,10^{h}$ 27° $+0.5$ *Oct. 7, 16^{h} $20,10^{h}$ 20° -0.3 *Oct. 7, 16^{h}	J.T.)Elong. EastMag.Date (U.T.)Elong. West $7,05^{h}$ 19° -0.3 Feb. 16, 15^{h} 26° $28,02^{h}$ 21° $+0.4$ June 15, 09^{h} 23° $26,10^{h}$ 27° $+0.5$ *Oct. $7, 16^{h}$ $20,10^{h}$ 20° -0.3 *	

VENUS

Since the orbit of Venus lies within that of the earth, its apparent motion is like Mercury's but is much slower and more stately. At inferior conjunction, it comes within 50 million km of the earth, and its nearness and its reflective cloud layer make it the brightest of the planets. In size and structure, it is much like the earth, but it has a thick layer of clouds and a dense, hot atmosphere of carbon dioxide. It is visible to the unaided eye in the daytime, if one knows where to look. In a small telescope, it displays a series of phases like those of the moon.

Venus is easily identified by its great brilliance, though it is not as conspicuous in 1976 as in other years. It was at greatest elongation west (morning sky) late in 1975, and in 1976 it will gradually close in on the sun until mid-May, when it becomes too close to the sun for observation. From late July onwards, it is visible in the evening sky, reaching greatest elongation east in early 1977. Venus is in conjunction with Mars on September 10.

MARS

Since the orbit of Mars is outside that of the earth, its planetary phenomena are quite different from those of Mercury and Venus. At intervals of about 780 days (the synodic period), Mars can be seen in opposition to the sun. Its distance from earth is then smallest and (if Mars is at perihelion) can be as small as 56 million km. Such close approaches occur at intervals of 15 to 17 years; the most recent was in 1971.

The atmosphere of Mars is thin and consists mainly of carbon dioxide; some surface features are distinctly visible in a good telescope. Mariner spacecraft have photographed most of the surface. Much of it is covered by shallow impact craters, but there are also volcanoes, canyons and wind-swept deserts.

Mars—always conspicuous because of its reddish colour—is visible in the evening until early October, when it becomes too close to the sun for observation. On January 1, Mars is in Taurus, and has a magnitude of -1.2 and an apparent diameter of 30 arc sec. On April 7, at 20^h E.S T., it occults ε Gem (see "Planetary Appulses"). On the evening of May 5, still in Gemini, it makes a fine grouping with Saturn and the crescent moon. On May 11, at 21^h E.S.T., it is 1.3° N. of Saturn, the latter being the brighter. On July 5, at 13^h E.S.T., it is 0.7° N. of Regulus in Leo.

JUPITER

Jupiter, the giant of the sun's family, is a fine object for the telescope. Belts of clouds may be observed, interrupted by irregular spots which may be short-lived or persist for weeks. The "great red spot" has been visible for centuries. The flattening of the planet, due to its fast rotation, is conspicuous, and the phenomena of its satellites provide a continual interest. The four largest satellites are of great interest to professional astronomers, who are studying them as intensively as they studied the planets a few years ago.

In early 1976, Jupiter dominates the region of Pisces, in the evening sky. Conjunction occurs on April 27, after which Jupiter moves into the morning sky. In September and October, it is in Taurus, directly between the Hyades and the Pleiades. At opposition on November 18, it has a magnitude of -2.4 and an apparent diameter of about 95 arc sec. By the end of the year, it is visible in the evening sky, sharing the spotlight with Venus.

SATURN

Saturn was the outermost planet known until modern times and, with its unique system of rings, is one of the finest of celestial objects in a good telescope. The plane of the rings makes an angle of 27° with the plane of the planet's orbit, and twice during the planet's revolution period of $29\frac{1}{2}$ years, the rings appear to open out widest; then they slowly close until, midway between the maxima, the rings are presented edgewise to the sun or the earth, at which times they are invisible. The rings were open widest in 1973, the southern face being visible (see also "Saturn and its Satellites").

Saturn is at opposition on January 20, when its magnitude is -0.1 and its apparent diameter is 18.5 arc sec. Moving westward through Cancer, it enters Gemini in late winter. On the evening of May 5, it makes a fine grouping with Mars and the crescent moon, and on May 11, at 21^h E.S.T., it is 1.3° S. of Mars. Conjunction occurs on July 29, after which Saturn moves into the morning sky.

URANUS



Although Uranus at opposition can be seen with the unaided eye under a clear, dark sky, it was apparently unknown until 1781, when it was accidentally discovered



(telescopically) by William Herschel. It can easily be seen with binoculars, and in a telescope it shows a small, greenish, almost featureless disc. Jupiter, Saturn, Uranus and Neptune are rather similar in the sense that their interiors consist mainly of hydrogen and helium; their atmospheres consist of these same elements, and simple compounds of hydrogen.

Throughout most of 1976, Uranus is in Virgo and passes near λ Vir in early January, mid-March and mid-October (see map). At opposition on April 25, its magnitude is + 5.7 and its apparent diameter is 3.9 arc sec.

NEPTUNE

The discovery of Neptune in 1846, after its existence in the sky had been predicted from independent calculations by Leverrier in France and Adams in England, was regarded as the crowning achievement of Newton's theory of universal gravitation. Actually, Neptune had been seen—but mistaken for a star—several times before its "discovery".

In 1976, Neptune is far south on the ecliptic, in Ophiuchus (see map) and not wellplaced for northern observers. At opposition on June 2 (E.S.T.), its magnitude is +7.7 and its apparent diameter is 2.5 arc sec.



PLUTO

Pluto, the most distant known planet, was discovered at the Lowell Observatory in 1930, as a result of an extensive search started two decades earlier by Percival Lowell. The faint star-like image was first detected by Clyde Tombaugh by comparing photographs taken on different dates. Little is known about the exact mass, radius and density of this planet. It varies in brightness with a period of 6.4 days, apparently due to its rotation.

At opposition on March 30, its astrometric position is R.A. (1950) $13^{h}04^{m}5$, Dec. (1950) $+12^{\circ}06'$, its apparent magnitude is +14, and its distance from earth is 4440 million km.

THE OBSERVATION OF THE MOON

During 1976 the ascending node of the moon's orbit moves from Libra into Virgo (Ω from 229° to 210°). See p. 61 for occultations of stars.

The sun's selenographic colongitude is essentially a convenient way of indicating the position of the sunrise terminator as it moves across the face of the moon. It provides an accurate method of recording the exact conditions of illumination (angle of illumination), and makes it possible to observe the moon under exactly the same lighting conditions at a later date.

The sun's selenographic colongitude is numerically equal to the selenographic longitude of the sunrise terminator reckoned eastward from the mean centre of the disk. Its value increases at the rate of nearly 12.2° per day or about $\frac{1}{2}$ ° per hour; it is approximately 270°, 0°, 90° and 180° at New Moon, First Quarter, Full Moon and Last Quarter respectively. (See the tabulated values for 0 h U.T. starting on p. 35.)

Sunrise will occur at a given point *east* of the central meridian of the moon when the sun's selenographic colongitude is equal to the eastern selenographic longitude of the point; at a point *west* of the central meridian when the sun's selenographic colongitude is equal to 360° minus the western selenographic longitude of the point. The longitude of the sunset terminator differs by 180° from that of the sunrise terminator.

The sun's selenographic latitude varies between $+1\frac{1}{2}^{\circ}$ and $-1\frac{1}{2}^{\circ}$ during the year.

By the moon's libration is meant the shifting, or rather apparent shifting, of the visible disk. Sometimes the observer sees features farther around the eastern or the western limb (libration in longitude), or the northern or southern limb (libration in latitude). The quantities called the earth's selenographic longitude and latitude are a convenient way of indicating the two librations. When the libration in longitude, that is the selenographic longitude of the earth, is positive, the mean central point of the disk of the moon is displaced eastward on the celestial sphere, exposing to view a region on the west limb. When the libration in latitude, or the selenographic latitude of the earth, is positive, the mean central point of the disk of the moon is displaced towards the south, and a region on the north limb is exposed to view.

In the Astronomical Phenomena Month by Month the dates of the greatest positive and negative values of the libration in longitude are indicated by l in the column headed "Sun's Selenographic Colongitude," and their values are given in the footnotes. Similarly the extreme values of the libration in latitude are indicated by b .

THE SKY MONTH BY MONTH

By John F. Heard

Introduction—In the monthly descriptions of the sky on the following pages, positions of the sun and planets are given for 0 h Ephemeris Time, which differs only slightly from Standard Time on the Greenwich meridian. The times of transit at the 75th meridian are given in *local mean time*; to change to Standard Time, see p. 14. Estimates of altitude are for an observer in latitude 45° N.

The Sun—The values of the equation of time are for noon E.S.T. on the first and last days of the month. For times of sunrise and sunset and for changes in the length of the day, see pp. 15–20. See also p. 7.

The Moon—Its phases, perigee and apogee times and distances, and its conjunctions with the planets are given in the "Astronomical Phenomena Month by Month" For times of moonrise and moonset, see pp. 22–27.

The Planets—Further information in regard to the planets, including Pluto, is found on pp. 28–31. For the configurations of Jupiter's satellites, see "Astronomical Phenomena Month by Month", and for their eclipses, see p. 77.

In the Configurations of Jupiter's Satellites, O represents the disk of the planet, d signifies that the satellite is on the disk, * signifies that the satellite is behind the disk or in the shadow. Configurations are for an inverting telescope.

The configurations have been read from diagrams in the American Ephemeris and Nautical Almanac. Where two satellites are nearly coincident, it is difficult to tell the correct order of the satellites from the diagram. Such ambiguous cases are indicated by bold face type: thus 12304 may actually be 13204. An hour's observation usually reveals the correct configuration, because the apparent motion of the innermost satellites is much faster than that of the outermost. Also, the four satellites differ slightly in apparent magnitude.

Satellites move from east to west across the face of the planet, and from west to east behind it. Before opposition, shadows fall to the west, and after opposition, to the east.

Minima of Algol—The times of mid-eclipse are given in "Astronomical Phenomena Month by Month" and are calculated from the ephemeris

heliocentric minimum = 2440953.4677 + 2.8673285 E and are rounded off to the nearest ten minutes.

THE SKY FOR JANUARY 1976

During this month which follows the winter solstice the days are lengthening as we know, but, as you may notice from a table of sunrise and sunset, the lengthening is not symmetrical; in Toronto, for example, sunrise gets earlier by about 14 minutes, but sunset gets later by about 35 minutes. The explanation? A hint: look at what is happening to the equation of time. Contrast the situation in February.

An hour after sunset at mid-month the summer triangle (Altair, Vega, Deneb) is sinking into the west and Orion, "pride of the winter skies", is rising in the east. Jupiter, Mars and Saturn, lined up between the meridian and the eastern horizon, trace out the ecliptic for us. Have a try at observing Mercury as an evening star early in the month. And if you are awake just before dawn on a clear night you will surely see Venus.

The bright asteroid Ceres is well placed for observation during January, moving through Taurus about halfway between the Pleiades and the Hyades. Look for it with binoculars, referring to the information on pp. 74–75

The Sun—During January the sun's R.A. increases from 18 h 42 m to 20 h 55 m and its Decl. changes from 23° 05' S. to 17° 24' S. The equation of time changes from -3 m 22 s to -13 m 25 s. The earth is at perihelion on the 4th at a distance of 147,102,000 km (91,405,000 miles) from the sun.

Mercury on the 1st is in R.A. 20 h 00 m, Decl. $22^{\circ} 27'$ S., and on the 15th is in R.A. 20 h 46 m, Decl. $16^{\circ} 52'$ S During the first half of the month it may be seen as an evening star very low in the south-west just after sunset. Greatest eastern elongation is on the 7th at which time Mercury is about 11 degrees above the south-western horizon at sunset. By the 23rd it is in inferior conjunction.

Venus on the 1st is in R.A. 15 h 51 m, Decl. 17° 52' S., and on the 15th it is in R.A. 17 h 01 m, Decl. 21° 00' S., mag. -3.6, and transits at 9 h 28 m. It is prominent in the south-eastern sky for nearly three hours before sunrise.

Mars on the 15th is in R.A. 4 h 53 m, Decl. $25^{\circ} 43'$ N., mag -0.7, and transits at 21 h 14 m. Now a month past opposition and closest approach, Mars has faded by nearly a magnitude but still outshines any nearby stars. In Taurus, it is well up in the east at sunset and sets about three hours before dawn. On the 20th it is stationary in right ascension and resumes eastward motion.

Jupiter on the 15th is in R.A. 1 h 04 m, Decl. 5° 29' N., mag. -2.0, and transits at 17 h 26 m. In Pisces, it is nearly on the meridian at sunset and sets at about midnight.

Saturn on the 15th is in R.A. 8 h 09 m, Decl. $20^{\circ} 26'$ N., mag -0.1, and transits at 0 h 34 m. In Cancer, it rises about as the sun sets. Opposition is on the 20th.

Uranus on the 15th is in R.A. 14 h 19 m, Decl. 13° 19' S., and transits at 6 h 39 m.

Neptune on the 15th is in R.A. 16 h 47 m, Decl. 20° 52' S., and transits at 9 h 10 m.
1976				JANUARY E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 19 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	d	h	m		hm		o
Thur.	1	09	40	M New Moon		40132	257.73
Fri.	2				5 50	3104d	269.92
Sat.	3	01		Mercury 7° S. of Moon		32014	282.11
Sun.	4	02		Ouadrantid meteors		3024*	294.30
		06		Earth at perihelion			
Mon.	5			F	2 40	13024	306.48 ^t
Tues.	6					20134	318.66
Wed.	7	00		Mercury greatest elong, E. (19°)	23 30	21034	330.83
Thur.	8	07		Venus 7° N. of Antares		01234	343.00
		12		Moon at apogee (404,410 km)			
Fri.	9	07		Jupiter 4° S. of Moon		13024	355.17
		07	40	First Quarter			
Sat.	10				20 20	32041	7.33
Sun.	11			Mercury at ascending node		3410*	19.48
		23		Venus 0.4° N. of Neptune			
Mon.	12	21		Ceres 0.3° N. of Moon. Occ'n.		43102	31.62
Tues.	13	00		Juno stationary	17 10	42013	43.76
		15		Mercury stationary			
		22		Mars 5° N. of Moon			
Wed.	14					42103	55.90
Thur.	15			Mercury at perihelion		40123	68.03 ¹
Fri.	16	23	47	😌 Full Moon	14 00	41032	80.16
Sat.	17	08		Saturn 5° N. of Moon		432O1	92.28
Sun.	18					3410*	104.41
Mon.	19				10 50	3O42d	116.54
Tues.	20	06		Saturn at opposition		2O34*	128.66
		08		Moon at perigee (366,880 km)			
		10		Ceres stationary			
		15		Mars stationary			
Wed.	21	14		Pluto stationary		21034	140.80
Thur.	22				7 40	O1234	152.94
Fri.	23	01		Mercury in inferior conjunction		10324	165.09
		02		Spica 0.1° N. of Moon. Occ'n. ¹			
		18	04	Last Quarter			
Sat.	24	02		Uranus 2° N. of Moon		32014	177.24
Sun.	25				4 30	312O4	189.40
Mon.	26			Mercury at greatest hel. lat. N.		30124	201.57
		16		Neptune 0.6° S. of Moon. Occ'n.			
Tues.	27					340**	213.74
Wed.	28	03		Venus 2° S. of Moon	1 20	421O3	225.93
Thur.	29					40123	238.11 ¹
Fri.	30				22 00	41O32	250.30
Sat.	31	01	20	New Moon		432O1	262.49

ASTRONOMICAL PHENOMENA MONTH BY MONTH

¹Jan. 2, +4.81°; Jan. 15, -5.88°; Jan. 29, +4.86°. ⁵Jan. 5, -6.65°; Jan. 19, +6.54°.

¹Visible in Central and S. America, S. Atlantic, S. Africa.

THE SKY FOR FEBRUARY 1976

An hour after sunset at mid-month the Great Square of Pegasus is "getting away" in the west; we won't see it easily after this month. Taurus with its brightest star Aldebaran and with the Pleiades is not far from the zenith. Near Aldebaran is Mars; identify them. Which is the brighter now? Watch how Mars' distance eastward from Aldebaran increases at an accelerating rate; by the end of April it will be near Castor and Pollux. South of Taurus is Orion, now best positioned for evening observing. Take a look at the centre star of Orion's sword with binoculars and admire the great Orion Nebula.

The Sun—During February the sun's R.A. increases from 20 h 55 m to 22 h 48 m and its Decl. changes from 17° 24' S. to 7° 37' S. The equation of time changes from -13 m 34 s to a maximum of -14 m 17 s on the 12 th and then to -12 m 30 s at the end of the month.

Mercury on the 1st is in R.A. 19 h 41 m, Decl. $18^{\circ} 13'$ S., and on the 15th is in R.A. 20 h 05 m, Decl. $19^{\circ} 44'$ S. It is in greatest western elongation on the 16th, so at about this time it might be seen low in the south-east just before sunrise. However, this is an unfavourable elongation, Mercury being only about 9 degrees above the horizon at sunrise.

Venus on the 1st is in R.A. 18 h 30 m, Decl. 22° 19' S., and on the 15th it is in R.A. 19 h 44 m, Decl. 21° 03' S., mag. -3.4, and transits at 10 h 09 m. It rises about an hour and a half before the sun and is only about 12 degrees above the south-eastern horizon at sunrise.

Mars on the 15th is in R.A. 5 h 08 m, Decl. $25^{\circ} 42'$ N., mag. +0.2, and transits at 19 h 29 m. In Taurus, it is nearing the meridian at sunset and sets about three hours after midnight. Again it has faded by nearly a magnitude in the past month.

Jupiter on the 15th is in R.A. 1 h 21 m, Decl. 7° 21' N., mag. - 1.8, and transits at 15 h 42 m. In Pisces, it is well past the meridian at sunset and sets before midnight.

Saturn on the 15th is in R.A. 7 h 59 m, Decl. 20° 59' N., mag 0.0, and transits at 22 h 17 m. In Cancer, it is well up in the east at sunset.

Uranus on the 15th is in R.A. 14 h 20 m, Decl. 13° 25' S., and transits at 4 h 42 m.

Neptune on the 15th is in R.A. 16 h 50 m, Decl. 20° 57' S., and transits at 7 h 12 m.

1976				FEBRUARY E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 20 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	d	h	m		hm		0
Sun.	1					43120	274.68
Mon.	2				18 50	43012	286.88
Tues.	3	13		Mercury stationary		40**d	299.07
Wed.	4					24103	311.25
Thur.	5	08 22		Moon at apogee (405,100 km) Jupiter 4° S. of Moon	15 40	O1243	323.44
Fri.	6			-		10324	335.61
Sat.	7					23014	347.79
Sun.	8	05	05	First Quarter	12 30	32104	359.96
Mon.	9					30124	12.12
Tues.	10	11		Mars 5° N. of Moon		13O24	24.27
		20		Uranus stationary			
Wed.	11				9 20	2O34d	36.42
Thur.	12					O143*	48.57 ¹
Fri.	13	14		Saturn 5° N. of Moon		14O23	60.71
Sat.	14				6 10	423O1	72.84
Sun.	15	11	43	😨 Full Moon		43210	84.98 ^b
Mon.	16	10		Mercury greatest elong. W. (26°)		43012	97.11
		17		Juno 0.02° S. of Moon. Occ'n.			5
Tues.	17	05		Moon at perigee (361,310 km)	3 00	43102	109.24
Wed.	18			Mercury at descending node		42013	121.38
Thur.	19			Venus at descending node	23 50	403**	133.52
		08		Spica 0.1° S. of Moon. Occ'n. ¹			
Fri.	20	08		Uranus 1° N. of Moon		41023	145.66
Sat.	21					24301	157.82
Sun.	22	03	16	Last Quarter	20 40	32104	169.98
		23		Neptune 0.9° S. of Moon. Occ'n.			
Mon.	23					30124	182.15
Tues.	24					31024	194.32 ⁴
Wed.	25				17 30	20134	206.50
Thur.	26					21034	218.69
Fri.	27	09		Venus 6° S. of Moon		10234	230.89
a .		19		Mercury 7° S. of Moon			
Sat.	28			Mercury at aphelion	14 20	20314	243.08
Sun.	29	18	25	New Moon		32104	255.28

¹Feb. 12, -7.07° ; Feb. 24, $+6.06^{\circ}$. ^bFeb. 1, -6.56° ; Feb. 15, $+6.53^{\circ}$; Feb. 28, -6.60° . ¹Visible in S. America.

THE SKY FOR MARCH 1976

This is the month of the vernal equinox. Notice that this happens on March 20 this year. Can you think of a reason why it is "early"? On this day the sun is on the equator and so we say that day and night are equal. Yet if you look at a table of sunrise and sunset for middle latitudes it will appear that the day is about 15 minutes longer than the night. This is partly because sunrise is defined as the moment when the sun's upper limb (rather than its centre) clears the horizon and partly because the refraction of the earth's atmosphere "lifts" the sun a bit. During March, the sunrise and sunset points move northward at the maximum rate. You can easily see this effect by observing the sunrise or sunset point over a period of two or three weeks.

The Sun—During March the sun's R.A. increases from 22 h 48 m to 0 h 42 m and its Decl. changes from 7° 37′ S. to 4° 30′ N. The equation of time changes from -12 m 18 s to -4 m 04 s. On the 20th at 6 h 50 m E.S.T. the sun crosses the equator on its way north, enters the sign of Aries and spring commences. This is the vernal equinox.

Mercury on the 1st is in R.A. 21 h 21 m, Decl. 17° 05' S., and on the 15th is in R.A. 22 h 46 m, Decl. 10° 17' S. It is too close to the sun for observation this month.

Venus on the 1st is in R.A. 21 h 01 m, Decl. 17° 29' S., and on the 15th it is in R.A. 22 h 10 m, Decl. 12° 27' S., mag. -3.3, and transits at 10 h 39 m. It rises barely an hour before the sun and is only about 9 degrees above the south-eastern horizon at sunrise.

Mars on the 15th is in R.A. 5 h 53 m, Decl. $25^{\circ} 49'$ N., mag. +0.8, and transits at 18 h 20 m. Moving from Taurus into Gemini, it is approximately on the meridian at sunset and sets a few hours after midnight. It now matches nearby Aldebaran in brightness.

Jupiter on the 15th is in R.A. 1 h 43 m, Decl. 9° 36' N., mag. -1.6, and transits at 14 h 10 m. Moving from Pisces into Aries, it is well down in the west at sunset and sets within three hours.

Saturn on the 15th is in R.A. 7 h 53 m, Decl. 21° 17' N., mag. +0.2, and transits at 20 h 18 m. Near the boundary between Cancer and Gemini, it is high in the eastern sky at sunset. On the 27th it is stationary in right ascension and resumes direct or eastward motion among the stars.

Uranus on the 15th is in R.A. 14 h 18 m, Decl. 13° 16' S., and transits at 2 h 46 m.

Neptune on the 15th is in R.A. 16 h 51 m, Decl. 20° 57' S., and transits at 5 h 19 m.

1 97 6				MARCH E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 20 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	4	L			h	1	
Mon			ш	June at apposition	n m	24021	267 40
Tues	2	00		Juno at opposition	11 10	42102	207.49
Wed	2	22		Moon at an area $(406,000 \text{ km})$	11 10	43102	219.09
Thur		15		Juniter ^{3°} S of Moon		42013	291.09
Fri	5	15		Jupiter 5 S. Of WIGON	8 00	42103	304.09
Sat	6]			0.00	40230	228 10
Sun	7					40130	240 68
Mon	8	23	38	> First Quarter	1 50	34021	252 86
Tues	9	14	50	Mars 6 N of Moon	4 50	31402	5 04
Wed	10	14				2014*	17 21
Thur	11	22		Saturn 5° N of Moon	1 30	21034	20 381
Fri	12	22		Saturn 5 11. Or Woon	1 50	01234	41 54
Sat	13				22 20	0234*	53 70
Sun	14				22 20	23104	65 85
Mon	15	20		Neptune stationary		3014*	78 00
	1.0	21	53	💬 Full Moon		5014	/0.00
Tues.	16	14	00	Moon at perigee (357.640 km)	19 10	31024	90 14
Wed.	17	18		Spica 0.3° S. of Moon. Occ'n	12 10	23041	102 29
Thur.	18	15		Uranus 1° N. of Moon, Occ'n.		24103	114 44
Fri.	19	10			16 00	40123	126 59
Sat.	20			Mercury at greatest hel, lat, S.		41023	138 75
		06	50	Equinox. Spring begins			
Sun.	21	05		Neptune 1° S. of Moon. Occ'n. ¹		423Od	150.92
Mon.	22	13	54	C Last Ouarter	12 50	4301*	163.09
Tues.	23					43102	175.27
Wed.	24					42301	187.46
Thur.	25			Venus at aphelion	9 40	42103	199.65
Fri.	26			•		O4 123	211.85
Sat.	27	13		Saturn stationary		10234	224.06
Sun.	28	19		Venus 6° S. of Moon	6 30	23014	236.27
Mon.	29					32014	248.48
Tues.	30	12	08	New Moon		31024	260.70
		17		Pluto at opposition			
Wed.	31	05		Moon at apogee (406,510 km)	3 20	32014	272.92

¹Mar. 11, -7.93° ; Mar. 23, $+7.09^{\circ}$. ^bMar. 13, $+6.61^{\circ}$; Mar. 26, -6.73° . ¹Visible in N. America.

THE SKY FOR APRIL 1976

Notice in the section on eclipses that a partial eclipse of the sun on the 29th will be visible in some parts of North America. April-May and October-November are the eclipse seasons now, but 1976 is a poor eclipse year for North America.

Notice that the greatest eastern elongation of Mercury on the 27th is called a very favourable one. This is because the ecliptic east of the setting sun at this season in middle latitudes is nearly perpendicular to the horizon, so the altitude of Mercury at sunset is nearly as great as its angular distance from the sun (21°) .

Mars, though a good deal fainter than before, should be identifiable. Watch it move rapidly eastward towards Saturn. Jupiter moves "into" the sun this month and we lose it for a while.

There is an occultation of ε Gem by Mars at almost exactly 20.00 E.S.T. on the 7th. This is visible over much of North America.

The Sun—During April the sun's R.A. increases from 0 h 42 m to 2 h 33 m and its Decl. changes from 4° 30' N. to 15° 03' N. The equation of time changes from -3 m 46 s to +2 m 51 s, being zero on the 15th. There is an annular eclipse of the sun on the 29th which is visible as a partial eclipse near the lower St. Lawrence, in the Atlantic provinces and the New England states.

Mercury on the 1st is in R.A. 0 h 41 m, Decl. 3° 08' N., and on the 15th is in R.A. 2 h 25 m, Decl. 15° 43' N. It is in superior conjunction on the 1st, and by the 27th it is in greatest eastern elongation, a very favourable one, at which time Mercury is nearly 19 degrees above the western horizon at sunset. During the last two weeks of the month it should be easy to see. On the evening of the 12th it is about 2° north of Jupiter which should facilitate locating it.

Venus on the 1st is in R.A. 23 h 29 m, Decl. 4° 55' S., and on the 15th it is in R.A. 0 h 32 m, Decl. 1° 51' N., mag. -3.3, and transits at 11 h 00 m. It rises in the east just about half an hour before the sun.

Mars on the 15th is in R.A. 6 h 58 m, Decl. $24^{\circ} 46'$ N., mag. + 1.3, and transits at 17 h 23 m. Moving rapidly through Gemini, it is now past the meridian at sunset and sets soon after midnight. It is no longer prominent among the nearby stars, but once located it will be interesting to watch it move eastward towards Saturn (see May).

Jupiter on the 15th is in R.A. 2 h 11 m, Decl. 12° 08' N., mag. -1.6, and transits at 12 h 35 m. It is visible low in the west just after sunset early in the month, but by the 27th it is in conjunction with the sun.

Saturn on the 15th is in R.A. 7 h 54 m, Decl. 21° 17' N., mag. +0.4, and transits at 18 h 17 m. In Cancer, it is close to the meridian at sunset and sets after midnight. See Mars.

Uranus on the 15th is in R.A. 14 h 14 m, Decl. 12° 53' S., and transits at 0 h 40 m.

Neptune on the 15th is in R.A. 16 h 50 m, Decl. 20° 54 'S., and transits at 3 h 16 m.

1976				APRIL E.S.T.	M C Al	in. of gol	Config. of Jupiter's Sat. 20 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	d	h	m		h	m		o
Thur.	1	09		Jupiter 2° S. of Moon			21034	285.14
		13		Mercury in superior conjunction				
Fri.	2						O1243	297.36
Sat.	3				0	10	10423	309.57
Sun.	4						24O1d	321.79
Mon.	5				21	00	432O*	334.00
Tues.	6	22		Mars 7° N. of Moon			431O2	346.20
Wed.	7	14	02	First Quarter			43O21	358.40
		19	56	Occultation of ε Gem by Mars				
Thur.	8	ļ		Mercury at ascending node	17	50	42103	10.59 ¹
		07		Saturn 6° N. of Moon				
Fri.	9							22.78
Sat.	10							34.96 ^b
Sun.	11				14	40		47.14
Mon.	12			Mercury at perihelion				59.30
		13		Mercury 1.9° N. of Jupiter				
Tues.	13							71.47
Wed.	14			Mars at greatest hel. lat. N.	11	30		83.63
		02		Moon at perigee (356,940 km)				
	Í	05		Spica 0.3° S. of Moon. Occ'n. ¹				
	ļ	06	49	😨 Full Moon				
Thur.	15	01		Uranus 1° N. of Moon. Occ'n.				95.80
		16		Juno stationary				
Fri.	16			Venus at greatest hel. lat. S.				107.96
Sat.	17	14		Neptune 1° S. of Moon. Occ'n.	8	10		120.13
Sun.	18							132.30
Mon.	19							144.48
Tues.	20				5	00		156.67
Wed.	21	02	14	Last Quarter				168.86
		21		Lyrid meteors				
Thur.	22							181.06
Fri.	23			Mercury at greatest hel. lat. N.	1	50		193.27
Sat.	24							205.49
Sun.	25	00		Uranus at opposition	22	40		217.71
Mon.	26							229.93
Tues.	27	07		Moon at apogee (406,400 km)				242.16
		15		Jupiter in conjunction with Sun				
	1	21		Mercury greatest elong. E. (21°)	1			
Wed.	28				19	30		254.39
Thur.	29	05	20	Mew Moon. Eclipse of ⊙, pg. 59				266.62
		06		Pallas in conjunction with Sun				
Fri.	30	23		Mercury 4° N. of Moon				278.86

¹Apr. 8, -8.01°; Apr. 20, +7.36°. ^bApr. 10, +6.75°; Apr. 22, -6.83°. ¹Visible in Central and S. America.

THE SKY FOR MAY 1976

If you missed seeing Mercury last month look low in the west just after sunset early this month.

An hour after sunset at mid-month Gemini is sinking into the west. Identify Castor and Pollux and watch Saturn and Mars move eastward across the southerly extension of the line joining them, Mars moving more rapidly than Saturn and passing it within a degree and a half on the 11th.

The Sun—During May the sun's R.A. increases from 2 h 33 m to 4 h 36 m and its Decl. changes from 15° 03' N., to 22° 02' N. The equation of time changes from +2 m 59 s to a maximum of +3 m 42 s on the 15th and then to +2 m 20 s at the end of the month.

The Moon—There is a partial eclipse of the moon on the 13th but this is not visible in any part of North America.

Mercury on the 1st is in R.A. 3 h 52 m, Decl. 22° 57' N., and on the 15th is in R.A. 4 h 01 m, Decl. 21° 03' N. For the first week of the month it should be seen quite easily low in the west just after sunset, but by the 20th it is in inferior conjunction.

Venus on the 1st is in R.A. 1 h 45 m, Decl. 9° 27' N., and on the 15th it is in R.A. 2 h 52 m, Decl. 15° 23' N., mag. -3.4, and transits at 11 h 21 m. It is difficult to observe, rising just minutes before the sun at mid-month.

Mars on the 15th is in R.A. 8 h 07 m, Decl. 21° 56' N., mag. + 1.6, and transits at 16 h 34 m. Moving into Cancer, it is now well past the meridian at sunset and sets at about midnight. It is no longer prominent in brightness. Late on the 11th (at 21 h E.S.T.) Mars and Saturn are in conjunction, Mars being 1°3 north.

Jupiter on the 15th is in R.A. 2 h 38 m, Decl. $14^{\circ} 27'$ N., mag -1.6, and transits at 11 h 05 m. In the latter part of the month it is easy to observe as a morning star rising to the north of east before the sun. On the morning of the 27th the proximity of Jupiter and the crescent moon will be a pretty sight.

Saturn on the 15th is in R.A. 8 h 01 m, Decl. 20° 58' N., mag. +0.5, and transits at 16 h 26 m. In Cancer, it is well past the meridian at sunset and sets at about midnight. See Mars.

Uranus on the 15th is in R.A. 14 h 09 m, Decl. 12° 29' S., and transits at 22 h 33 m.

Neptune on the 15th is in R.A. 16 h 48 m, Decl. 20° 49' S., and transits at 1 h 15 m.

1976				MAY E.S.T.	Min. of Algol	Config. of Jupiter's Sat.	Sun's Selen. Colong. 0h U.T.
	d	h	m		hm		0
Sat.	1				16 20		291.09
Sun	2						303.33
Mon	3						315.56
Tues.	4	23		η Aquarid meteors	13 10		327.79
		23		Mars 5° S. of Pollux			
Wed.	5	09		Mars 7° N. of Moon			340.01
		15		Saturn 6° N. of Moon			
Thur.	6						352.231
Fri.	7	00	17	First Quarter	10 00		4.44 ^b
Sat.	8						16.64
Sun.	9	11		Mercury stationary			28.84
Mon.	10				6 50		41.03
Tues.	11	09		Venus 0.2° S. of Jupiter			53.22
		15		Spica 0.3° S. of Moon. Occ'n.			
		21		Mars 1.3° N. of Saturn			
Wed.	12	10		Uranus 1° N. of Moon Occ'n.			65.40
		12		Moon at perigee (359,170 km)			
Thur.	13	15	04	Full Moon. Eclipse of <i>Q</i> , pg. 59.	3 40		77.58
Fri.	14	08		Vesta in conjunction with Sun			89.76
		23		Neptune 1° S. of Moon. Occ'n. ¹			1
Sat.	15						101.94
Sun.	16			Mercury at descending node	0 30		114.12
Mon.	17						126.31
Tues.	18				21 10		138.50
Wed.	19						150.69%
Thur.	20	07		Mercury in inferior conjunction			162.90
		16	22	C Last Quarter			
Fri.	21			Mars at aphelion	18 00		175.11
Sat.	22						187.33
Sun.	23						199.55
Mon.	24	19		Moon at apogee (405,600 km)	14 50		211.78
Tues.	25						224.01
Wed.	26			Mercury at aphelion			236.25
		23		Jupiter 0.8° S. of Moon. Occ'n.			
Thur.	27				11 40		248.49
Fri.	28	20	47	W New Moon			260.74
Sat.	29				0.00		272.99
Sun.	30				8 30		285.23
Mon.	31						297.48

¹May 6, -7.26° ; May 19, $+7.01^{\circ}$. ^bMay 7, $+6.82^{\circ}$; May 19, -6.79° . ¹Visible in N. America.

THE SKY FOR JUNE 1976

The summer solstice is on the 21st, the sun being then at its farthest north of the equator at declination $+23^{\circ} 26'$. So in middle northerly latitudes we have our longest day, the sun rising at about 60° east of north and setting about 60° west of north, and climbing to an altitude of about $68\frac{1}{2}^{\circ}$ at mid-day (for latitude 45° N). It is a combination of the most direct rays and the longest daily sunlight that gives us the maximum heating effect at this time. (But why, then, does our hottest weather usually come a month or more later? Accumulation is the answer.)

Realizing that the full moon is always exactly opposite to the sun in the sky, what will be noteworthy about the diurnal path of the full moon this month?

The Sun—During June the sun's R.A. increases from 4 h 36 m to 6 h 40 m and its Decl. changes from 22° 02' N. to 23° 07' N. The equation of time changes from +2 m 11 s to -3 m 40 s, being zero on the 13th. The summer solstice is on the 21st at 1 h 24 m E.S.T.

Mercury on the 1st is in R.A. 3 h 34 m, Decl. $15^{\circ} 27'$ N., and on the 15th is in R.A. 3 h 59 m, Decl. $16^{\circ} 53'$ N. On the 15th it is in greatest western elongation and may be glimpsed low in the east just before sunrise. However, this is an unfavourable elongation, Mercury being only 11 degrees above the horizon at sunrise.

Venus on the 1st is in R.A. 4 h 17 m, Decl. 20° 53' N., and on the 15th it is in R.A. 5 h 30 m, Decl. 23° 24' N., mag. -3.5, and transits at 11 h 57 m. It is too close to the sun all month for easy observation, superior conjunction being on the 17th.

Mars on the 15th is in R.A. 9 h 19 m, Decl. 17° 02' N., mag. + 1.8, and transits at 15 h 44 m. Moving from Cancer into Leo, it is well down in the west at sunset and sets about three hours later.

Jupiter on the 15th is in R.A. 3 h 06 m, Decl. 16° 30' N., mag. -1.6, and transits at 9 h 31 m. In Aries it rises to the north of east about two hours before the sun. It will be interesting to observe Jupiter relative to the crescent moon on the mornings of the 23rd and the 24th as the moon passes from a few degrees west of Jupiter to a few degrees east of him.

Saturn on the 15th is in R.A. 8 h 14 m, Decl. $20^{\circ} 22'$ N., mag. +0.5, and transits at 14 h 37 m. In Cancer, it is well down in the west at sunset and sets about two hours later.

Uranus on the 15th is in R.A. 14 h 05 m, Decl. 12° 09' S., and transits at 20 h 27 m.

Neptune on the 15th is in R.A. 16 h 44 m, Decl. 20° 43' S., and transits at 23 h 06 m.

1 97 6				JUNE E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 4 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	d	h	m		hm		0
Tues.	1	12		Mercury stationary			309.72
Wed.	2	01		Saturn 6° N. of Moon	5 20		321.96 ¹
		20		Neptune at opposition			
		21		Mars 7° N. of Moon			
Thur.	3					31024	334.20
Fri.	4					32O4d	346.43
Sat.	5	07	20	First Quarter	2 10	20134	358.65
Sun.	6					10234	10.87
Mon.	7				23 00	O 2 134	23.08
Tues.	8	00		Spica 0.4° S. of Moon. Occ'n. ¹		21034	35.29
		17		Uranus 1° N. of Moon			
Wed.	9	14		Moon at perigee (363,550 km)		3014*	47.48
Thur.	10				19 50	31042	59.68
Fri.	11			Venus at ascending node		3 42 O1	71.87
		08		Neptune 1° S. of Moon. Occ'n.			
		23	15	Full Moon			
Sat.	12					420**	84.06
Sun.	13				16 30	41023	96.24
Mon.	14					40213	108.43
Tues.	15	04		Mercury greatest elong. W. (23°)		42103	120.63
Wed.	16			Mercury at greatest hel. lat. S.	13 20	43201	132.83
Thur.	17	23		Venus in superior conjunction		43102	145.03
Fri.	18					34201	157.24
Sat.	19	08	15	C Last Quarter	10 10	204**	169.45
Sun.	20					O234d	181.67
Mon.	21	01	24	Solstice. Summer begins		01234	193.90
		12		Moon at apogee (404,590 km)			
Tues.	22	12		Mercury 3° N. of Aldebaran	7 00	21034	206.13
Wed.	23	18		Jupiter 0.1° S. of Moon. Occ'n.	r.	23014	218.37
Thur.	24					31024	230.61
Fri.	25	13		Vesta 0.8° S. of Moon. Occ'n.	3.50	3014d	242.86
		17		Mercury 1° N. of Moon. Occ'n.			
a .	2	23		Pluto stationary		22104	255 11
Sat.	26	-	50			23104	255.11
Sun.	21	09	50	W New Moon	0.40	40122	201.30
Mon.	28	1.2			0 40	40123	2/9.02
I ues.	29	13		Saturn o' N. OI Moon		42103	291.8/
wed.	30					42301	304.12

^{*i*}June 2, -6.01° ; June 16, $+6.22^{\circ}$; June 29, -5.14° . ^{*b*}June 3, $+6.74^{\circ}$; June 16, -6.70° ; June 30, $+6.60^{\circ}$. ^{*i*}Visible in Central and S. America.

THE SKY FOR JULY 1976

Early this month the earth is in aphelion; early in January it was at perihelion. The difference in distance from earth to sun between these two extremes is about 5,000,000 km or 3.3 per cent, which makes a difference in radiant heat received by the earth of nearly 7 per cent. Thus for the northern hemisphere the difference of large land masses in the northern hemisphere works the other way and tends to make our winters colder and summers hotter than those of the southern hemisphere.

The Sun—During July the sun's R.A. increases from 6 h 40 m to 8 h 45 m and its Decl. changes from 23° 07' N., to 18° 02' N. The equation of time changes from -3 m 51 s to a maximum of -6 m 27 s on the 26th and then to -6 m 18 s at the end of the month. The earth is in aphelion on the 3rd at a distance of 152,101,000 km (94,511,000 miles) from the sun.

Mercury on the 1st is in R.A. 5 h 30 m, Decl. $22^{\circ} 27'$ N., and on the 15th is in R.A. 7 h 36 m, Decl. $23^{\circ} 06'$ N. It is too close to the sun for observation, superior conjunction being on the 15th. However, a close conjunction with Venus on the 24th (at 9 h E.S.T., Mercury passing 0°4 north) offers an interesting opportunity to locate Mercury in the twilight sky by observing Venus on the evenings of the 23rd and 24th in binoculars or by wide-field telescope.

Venus on the 1st is in R.A. 6 h 56 m, Decl. $23^{\circ} 33'$ N., and on the 15th it is in R.A. 8 h 10 m, Decl. $21^{\circ} 16'$ N., mag. -3.4, and transits at 12 h 39 m. It is an evening star now, and at mid-month it stands about 6 degrees above the north-western horizon at sunset.

Mars on the 15th is in R.A. 10 h 28 m, Decl. 10° 42' N., mag. + 1.9, and transits at 14 h 55 m. In Leo, it is well down in the west at sunset and sets within about two hours. On the evening of the 5th it is less than a degree north of Regulus.

Jupiter on the 15th is in R.A. 3 h 30 m, Decl. 17° 59' N., mag. -1.8, and transits at 7 h 57 m. Moving into Taurus, it rises at about midnight. Between the mornings of the 21st and the 22nd the waning moon switches from west to east of Jupiter and in New Zealand an occultation will be visible.

Saturn on the 15th is in R.A. 8 h 29 m, Decl. $19^{\circ} 34'$ N., mag. +0.5, and transits at 12 h 54 m. It is too close to the sun for easy observation. On the 29th it is in conjunction with the sun.

Uranus on the 15th is in R.A. 14 h 04 m, Decl. 12° 05' S., and transits at 18 h 28 m.

Neptune on the 15th is in R.A. 16 h 41 m, Decl. 20° 38' S., and transits at 21 h 05 m.

1976				JULY E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 4 h E.S.T.	Sun's Selen. Colong. 0h U.T.
	d	h			h m		0
Thur	1	10	m	Mars 6° N of Moon	11 111	43102	316 36
Fri	2	23		Farth at aphelion		43021	328.61
Sat	3	25		Latinat appoint	18 10	42310	340.84
Sun	4	12	28	The First Quarter	10 10	40213	353.07
Mon.	5			Mercury at ascending node		4023*	5.30
	-	06		Spica 0.6° S. of Moon. Occ'n.			
		13		Mars 0.7° N. of Regulus			
		23		Uranus 1° N. of Moon. Occ'n.			
Tues.	6	21		Moon at perigee (368,370 km)	15 00	21043	17.51
Wed.	7					20314	29.72
Thur.	8	16		Neptune 1° S. of Moon. Occ'n.		31024	41.92
Fri.	9			Mercury at perihelion	11 50	3O214	54.12
Sat.	10					23 104	66.32
Sun.	11	07		Uranus stationary		O134*	78.51
		08	09	😨 Full Moon			
Mon.	12				8 40	O234*	90.70
Tues.	13					2O43d	102.89
Wed.	14					20413	115.08
Thur.	15			Venus at perihelion	5 30	34102	127.28
		10		Mercury in superior conjunction			
Fri.	16					43021	139.48
Sat.	17					43210	151.69
Sun.	18				2 20	42031	163.90
Mon.	19	01	29	Last quarter		41023	176.12
_		06		Moon at apogee (404,060 km)		40010	100 04
Tues.	20			Mercury at greatest hel. lat. N.	23 10	42013	188.34
Wed.	21	12		Jupiter 0.5° N. of Moon. Occ'n.		42013	200.57
Thur.	22	10			20.00	34102	212.81
Fri.	23	19		Vesta 0.5° N. of Moon. Occ n.	20 00	30412	225.05
Sat.	24	16	20	Annulas of Vests and SAO 77347		32104	237.29
Sun.	25	10	30	Appulse of vesta and SAO 77347		23014	249.34
Man	26	20	20	New Meen	16 40	10224	261 701
MOD.	20	20	39	W New Moon	10 40	01344	201.79
Tues.	21	10		8 A quarid meteors		2034*	286 20
Thur	20	19		Saturn in conjunction with Sun	13 30	31024	200.22
i nur.	27	21		Mars 5° N of Moon	15 50	1024	2,0.54
Fri	30	21				30124	310 78
Sat	31	23		Moon at perigee (369 100 km)		32104	323.02
<u> </u>	51	25				52104	

¹July 14, +5.37°; July 26, -5.25°. ^bJuly 13, -6.58°; July 27, +6.51°.

THE SKY FOR AUGUST 1976

This month we have, on the 26th, a greatest eastern elongation of Mercury which normally results in a good opportunity to see Mercury as an evening star. (See, for example April.) But, whereas in April the ecliptic east of the setting sun was steep relative to the horizon, in August it is very oblique and, in addition, Mercury is about 4 degrees south of the ecliptic. Thus while Mercury on the 26th is 27 degrees east of the sun along the ecliptic, it is only about 6 degrees above the western horizon at sunset and so very low indeed when it is dark enough to see the planet. This is therefore a very unfavourable elongation.

The Sun—During August the sun's R.A. increases from 8 h 45 m to 10 h 41 m and its Decl. changes from 18° 02' N. to 8° 19' N. The equation of time changes from -6 m 15 s to -0 m 09 s.

Mercury on the 1st is in R.A. 9 h 53 m, Decl. 14° 20' N., and on the 15th is in R.A. 11 h 12 m, Decl. 4° 44' N. On the 26th it is in greatest eastern elongation, but this is a very unfavourable one, Mercury being scarcely 6 degrees above the western horizon at sunset.

Venus on the 1st is in R.A. 9 h 35 m, Decl. 15° 53' N., and on the 15th it is in R.A. 10 h 41 m, Decl. 9° 51' N., mag. -3.3, and transits at 13 h 08 m. It is an evening star standing about 8 degrees above the western horizon at sunset and setting within an hour.

Mars on the 15th is in R.A. 11 h 40 m, Decl. 3° 01' N., mag. + 1.9, and transits at 14 h 04 m. Moving from Leo into Virgo, it is now so low in the west at sunset as to be difficult to observe.

Jupiter on the 15th is in R.A. 3 h 49 m, Decl. 18° 58' N., mag. -1.9, and transits at 6 h 13 m. In Taurus, it rises late in the evening. During the night of the 17th-18th it will be interesting to watch the moon overtake and pass Jupiter, being just 1° south of him at 04 hours E.S.T. In the southern part of South America this will be seen as an occultation.

Saturn on the 15th is in R.A. 8 h 45 m, Decl. $18^{\circ} 37'$ N., mag. +0.5, and transits at 11 h 09 m. It is now in the morning sky but too close to the sun for easy observation.

Uranus on the 15th is in R.A. 14 h 06 m, Decl. 12° 16' S., and transits at 16 h 28 m.

Neptune on the 15th is in R.A. 16 h 40 m, Decl. 20° 37' S., and transits at 19 h 01 m.

1 97 6				AUGUST E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 3 h E.S.T.	Sun's Selen. Colong. 0 hU.T.
	d	h	m		hm		o
Sun.	1	11		Spica 0.8° S. of Moon. Occ'n.	10 20	423O1	335.26
Mon.	2	05		Uranus 0.9° N. of Moon. Occ'n.		41023	347.49
		17	07	First Quarter			
Tues.	3	01		Mercury 0.7° N. of Regulus		40213	359.71
Wed.	4	21		Neptune 1° S. of Moon	7 10	42103	11.92
Thur.	5			-		43O*d	24.13
Fri.	6			Venus at greatest hel. lat. N.		43012	36.33
Sat.	7	11		Venus 1.1° N. of Regulus	4 00	34210	48.53
Sun.	8			_		23401	60.72
Mon.	9	18	44	Full Moon		10423	72.91
Tues.	10			-	0 50	O2143	85.09 ¹
Wed.	11	23		Perseid meteors		21034	97.28
Thur.	12			Mercury at descending node	21 40	O14*d	109.47
Fri.	13					3O24*	121.66
Sat.	14					31204	133.85
Sun.	15				18 20	23014	146.05
Mon.	16	01		Moon at apogee (404,400 km)		10324	158.25
Tues.	17	19	13	Last Quarter		40213	170.46
Wed.	18	04		Jupiter 1° N. of Moon. Occ'n. ¹	15 10	42103	182.67
Thur.	19			-		42O31	194.89
Fri.	20					43102	207.11
Sat.	21				12 00	43Odd	219.34
Sun.	22			Mercury at aphelion		43201	231.58
Mon.	23	03		Neptune stationary		41O32	243.81
		18		Saturn 6° N. of Moon			
Tues.	24				8 50	40123	256.05
Wed.	25	06	01	Mew Moon		214O3	268.30
Thur.	26	05		Mercury greatest elong. E. (27°)		20134	280.54
		19		Venus 5° N. of Moon	}		ļ.
Fri.	27	06		Mercury 0.5° N. of Moon. Occ'n.	5 40	31O24	292.78
		10		Mars 4° N. of Moon			
		21		Moon at perigee (364,540 km)			
Sat.	28	18		Spica 1° S. of Moon. Occ'n. ²		3O24d	305.02
Sun.	29	12		Uranus 0.6° N. of Moon. Occ'n.		32014	317.25
Mon.	30				2 30	10324	329.48
Tues.	31	22	35	First Quarter		01234	341.70

¹Aug. 10, $+5.06^{\circ}$; Aug. 22, -6.06° . ^bAug. 9, -6.55° ; Aug. 24, $+6.56^{\circ}$. ¹Visible in S. America.

²Visible in N. America.

THE SKY FOR SEPTEMBER 1976

The autumnal equinox occurs on the 22nd. On that day if you are at latitude 44° N what is the meridian altitude of the sun? Is the interval from sunrise to sunset on the 22nd equal to the interval from sunset on the 22nd to sunrise on the 23rd? What two factors account for the inequality? Forgotten? See March.

Refer back to March relative to the rapid rate of change of the sunrise and sunset points at the time of the equinoxes.

An hour after sunset at mid-month we have Arcturus in the west, Antares in the south-west, and near the zenith we have the summer triangle of Altair, Vega and Deneb. Of these five famous stars can you find out from the Handbook which is the brightest, which the reddest, which the closest, which the most luminous (intrinsically)?

The Sun—During September the sun's R.A. increases from 10 h 41 m to 12 h 29 m and its Decl. changes from 8° 19' N. to 3° 09' S. The equation of time changes from + 0 m 10 s to + 10 m 09 s. On the 22nd at 15 h 48 m E.S.T. the sun crosses the equator moving southward, enters the sign of Libra and autumn commences.

Mercury on the 1st is in R.A. 1 h 14 m, Decl. 4° 51' S., and on the 15th is in R.A. 12 h 15 m, Decl. 6° 05' S. It is too close to the sun for observation until the end of the month, inferior conjunction being on the 22nd. However, by month's end it may be seen just before sunrise low in the east.

Venus on the 1st is in R.A. 11 h 59 m, Decl. 1° 26' N., and on the 15th it is in R.A. 13 h 01 m, Decl. 5° 45' S., mag. -3.3, and transits at 13 h 25 m. It is an evening star about 9 degrees above the western horizon at sunset and setting within an hour. On the evening of the 25th Venus and the crescent moon will make a pretty sight in the west. Also see Mars on this page.

Mars on the 15th is in R.A. 12 h 53 m, Decl. 5° 09' S., mag. +1.9, and transits at 13 h 15 m. It is too close to the sun for easy observation, though on the evening of the 10th when it is in conjunction with Venus (passing 0°4 south at 17 h E.S.T.) it would be interesting to locate Mars by observing Venus by binocular in the twilight sky.

Jupiter on the 15th is in R.A. 3 h 57 m, Decl. 19° 19' N., mag. -2.1, and transits at 4 h 19 m. In Taurus, it rises about three hours after sunset. On the 19th it is stationary in right ascension and begins to retrograde, or move westward, among the stars.

Saturn on the 15th is in R.A. 9 h 00 m, Decl. 17° 40' N., mag. +0.6, and transits at 9 h 22 m. In Cancer, it rises about three hours before the sun.

Uranus on the 15th is in R.A. 14 h 11 m, Decl. 12° 42' S., and transits at 14 h 31 m.

Neptune on the 15th is in R.A. 16 h 40 m, Decl. 20° 39' S., and transits at 17 h 00 m.

1976				SEPTEMBER E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 2 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
Wed.	d 1	h 02	m	Neptune 2° S. of Moon	h m 23 20	12034	° 353.91
Fri.	3				20.00	314O2	18.32
Sun. Mon	5	23		Mercury 5° S. of Venus	20 00	432O* 4103*	42.70^{u} 54.88
Tues. Wed.	7	07	52	🗇 Full Moon. Harvest Moon.	16 50	40123 41203	67.06 79.24
Thur.	9	09	52	Mercury stationary		42013	91.41
Fri. Sat.	10 11	17		Venus 0.4° N. of Mars	13 40	413O2 3 O4 21	103.59 115.77
Sun.	12	18		Mercury at greatest hel. lat. S. Moon at apogee (405,330 km)		32104	127.95
Mon. Tues.	13 14	14		Jupiter 1° N. of Moon	10 30	3O4*d O1324	140.13 152.32
Wed. Thur.	15 16	12	20	6 Last Quarter	7 20	12O34 2O134	164.51 176.70
Fri. Sat.	17 18					13O24 3O124	188.91 201.11
Sun.	19	16 20		Jupiter stationary Venus 3° N. of Spica	4 10	32104	213.33 ¹
Mon. Tues.	20 21	10 20		Saturn 6° N. of Moon Mercury in inferior conjunction		342O1 4O32*	225.54 ^b 237.77
Wed. Thur.	22 23	16 14	48 55	Equinox. Autumn begins New Moon	1 00	41O3d 42O13	249.99 262.22
Fri. Sat.	24 25	22 00		Moon at perigee (359,840 km) Mars 2° N. of Moon	21 40	41O32 43O12	274.45 286.68
		02 13 22		Spica 1° S. of Moon. Occ'n. Venus 0.7° N. of Moon. Occ'n. Uranus 0.3° N. of Moon. Occ'n.			
Sun.	26	14		Mars 2° N. of Spice	18 20	43210	298.91
Tues. Wed	27 28 29	09		Neptune 2° S. of Moon	10 30	0432* 10243	323.34
Thur.	30	05 06 17	12	Mercury stationary First Quarter Venus 0.5° S. of Uranus	15 20	20134	347.75

¹Sept. 5, $+5.73^{\circ}$; Sept. 19, -6.94° . ^bSept. 5, -6.63° ; Sept. 20, $+6.73^{\circ}$.

THE SKY FOR OCTOBER 1976

I wonder how many astronomy enthusiasts ever rationalize the position and phase of the moon. Take this month, for example. We see from the preceding page that new moon was on Sept. 23 and first quarter on Sept. 30. Therefore on Oct. 1 the moon is 8 days old and about a day past first quarter. Where will we expect to see it that evening as the sun sets? Well, for one thing it will be near the ecliptic (the moon's orbit plane is tilted only about 5 degrees to the earth's orbit plane or ecliptic) and somewhat more than 90 degrees east of the sun, so about an hour east of the meridian. Will it be high or low in altitude? Well, since it is somewhat more than 90 degrees along (or nearly along) the ecliptic eastward of the sun, it must have declination about the same as the sun will have in three month's time, namely about Jan. 1, therefore nearly as far south as it ever is, i.e. about 23 degrees south. And this means its altitude as it approaches the meridian is very low, say in the neighbourhood of 20 degrees for those of us near latitude 44° N. This kind of mental gymnastics can deceive us by as much as the 5 degrees by which the moon's orbit is inclined to the earth's orbit, and in fact it did so in our example because the moon is almost exactly 5 degrees north of the ecliptic on the 1st so a more accurate altitude for the moon at sunset is 25 degrees.

The Sun—During October the sun's R.A. increases from 12 h 29 m to 14 h 25 m and its Decl. changes from 3° 09' S. to 14° 24' S. The equation of time changes from +10 m 28 s to +16 m 22 s. A total eclipse of the sun on the 23rd will not be visible in this hemisphere.

Mercury on the 1st is in R.A. 11 h 35 m, Decl. $2^{\circ} 42'$ N., and on the 15th is in R.A. 12 h 26 m, Decl. $0^{\circ} 41'$ S. Greatest western elongation is on the 7th, a favourable one, Mercury then standing about 17 degrees above the eastern horizon at sunrise. Thus until past mid-month Mercury will be easily found near the eastern horizon just before sunrise.

Venus on the 1st is in R.A. 14 h 14 m, Decl. $13^{\circ} 29'$ S., and on the 15th it is in R.A. 15 h 22 m, Decl. $19^{\circ} 11'$ S., mag. -3.4, and transits at 13 h 48 m. It is about 10 degrees above the south-western horizon at sunset and sets about an hour later.

Mars on the 15th is in R.A. 14 h 08 m, Decl. 12° 46' S., mag. + 1.8, and transits at 12 h 32 m. It is too close to the sun for easy observation.

Jupiter on the 15th is in R.A. 3 h 53 m, Decl. 19° 04' N., mag. -2.3, and transits at 2 h 17 m. In Taurus, it rises about two hours after sunset and on the 11th will be about 1° north of the moon.

Saturn on the 15th is in R.A. 9 h 11 m, Decl. 16° 55' N., mag. +0.6, and transits at 7 h 35 m. In Cancer, it rises about at midnight.

Uranus on the 15th is in R.A. 14 h 17 m, Decl. 13° 16' S., and transits at 12 h 40 m.

Neptune on the 15th is in R.A. 16 h 43 m, Decl. 20° 46' S., and transits at 15 h 05 m.

1976				OCTOBER E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 1 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	d	h	m		hm		0
Fri.	1	1		Mercury at ascending node		10234	359.95
		1		Venus at descending node			
Sat.	2	í		_		30124	12.1416
Sun.	3	1			12 10	31204	24.32
Mon.	4	12		Pluto in conjunction with Sun		32014	36.49
Tues.	5			Mercury at perihelion		10324	48.66
Wed.	6				9 00	O4123	60.83
Thur.	7	11		Mercury greatest elong. W. (18°)		24 O13	72.99
		23	55	😨 Full Moon			
Fri.	8					4103*	85.15
Sat.	9				5 50	43012	97.31
Sun.	10	07		Moon at apogee (406,140 km)		43120	109.47
		07	58	Possible occ'n of SAO 153844 by Pallas			
Mon.	11	20		Jupiter 1° N. of Moon		43201	121.64
Tues.	12				2 40	4102*	133.80
Wed.	13					40123	145.97
Thur.	14	18		Juno in conjunction with Sun	23 30	24O3*	158.14
Fri.	15			Mercury at greatest hel. lat. N.		1043*	170.32
Sat.	16	03	59	C Last Quarter		30124	182.50
Sun.	17	1			20 10	31204	194.691
Mon.	18	00		Saturn 6° N. of Moon		32014	206.88
)	17		Mars 0.4° S. of Uranus			
Tues.	19					1024*	219.08
Wed.	20				17 00	O1234	231.29
Thur.	21	02		Orionid meteors		21034	243.49
Fri.	22					2O43d	255.71
Sat.	23	00	10	New Moon. Eclipse of sun, ng. 59.	13 50	34012	267.92
		08		Moon at perigee (357,150 km)			
Sun.	24	1				341Od	280.13
Mon.	25	08		Venus 4° S. of Moon	}	43201	292.35
		19		Neptune 2° S. of Moon			
Tues.	26				10 40	413O2	304.55
Wed.	27	20		Venus 3° N. of Antares		40123	316.76
Thur.	28					42103	328.95
Fri.	29			Mars at descending node	7 30	42O3d	341.14
		17	05	First Quarter			
Sat.	30	14		Uranus in conjunction with Sun		43012	353.32"
Sun.	31	01		Venus 3° S. of Neptune		314O2	5.50
		1			h	····-	·

¹Oct. 2, $+6.96^{\circ}$; Oct. 17, -7.42° ; Oct. 30, $+7.80^{\circ}$. ^bOct. 2, -6.74° ; Oct. 17, $+6.83^{\circ}$; Oct. 30, -6.82° .

THE SKY FOR NOVEMBER 1976

Will the penumbral eclipse of the moon on the night of Nov. 6 be perceptible by eye? Textbooks say only if the moon passes within 700 miles of the umbra. Let us see if this eclipse meets this criterion. At the moon's distance the average diameter of the earth's umbral shadow is about 5700 miles, of the penumbral shadow about 10,000 miles; and the moon's diameter is 2160 miles. The Handbook's section on eclipses says that the magnitude of this penumbral eclipse is 0.86, meaning that at mideclipse this fraction of the moon's diameter is within the penumbra. This fraction of 2160 is 1952 miles. The difference between the umbral radius and the penumbral is $\frac{1}{2}(10,000 - 5700) = 2150$. Thus the edge of the moon comes within 2150 - 1952 = 198 miles of the umbra. So the darkening of about one quarter of the moon's diameter should be perceptible. Draw a diagram of the event to scale and see if our conclusion looks reasonable.

The Sun—During November the sun's R.A. increases from 14 h 25 m to 16 h 29 m and its Decl. changes from 14° 24' S. to 21° 47' S. The equation of time changes from + 16 m 23 s to a maximum of + 16 m 24 s on the 3rd and then to + 11 m 09 s at the end of the month.

The Moon—A penumbral eclipse of the moon will be visible in most of North America on the night of the 6th.

Mercury on the 1st is in R.A. 14 h 11 m, Decl. 12° 18' S., and on the 15th is in R.A. 15 h 39 m, Decl. 20° 14' S. Superior conjunction is on the 7th and the planet remains too close to the sun for observation all month.

Venus on the 1st is in R.A. 16 h 49 m, Decl. 23° 54' S., and on the 15th it is in R.A. 18 h 03 m, Decl. 25° 25' S., mag. -3.5, and transits at 14 h 27 m. It is quite prominent low in the south-western sky for about two hours after sunset.

Mars on the 15th is in R.A. 15 h 33 m, Decl. $19^{\circ} 21'$ S., mag. + 1.6, and transits at 11 h 56 m. It is too close to the sun for observation, conjunction being on the 25th.

Jupiter on the 15th is in R.A. 3 h 38 m, Decl. $18^{\circ} 17'$ N., mag. -2.4, and transits at 0 h 00 m and 23 h 56 m (two transits on the same date!) In Taurus, it rises at about sunset and is visible all night. At about 20 hours on the 7th Jupiter passes 1° north of the moon. On the 18th Jupiter is at opposition.

Saturn on the 15th is in R.A. 9 h 18 m, Decl. $16^{\circ} 32'$ N., mag. +0.6, and transits at 5 h 39 m. Near the boundary between Cancer and Leo, it rises before midnight. On the 28th it is stationary in right ascension and begins to retrograde, or move westward, among the stars.

Uranus on the 15th is in R.A. 14 h 25 m, Decl. 13° 54' S., and transits at 10 h 46 m.

Neptune on the 15th is in R.A. 16 h 47 m, Decl. 20° 54' S., and transits at 13 h 07 m.

197 6				NOVEMBER E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 0 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	d	h	m		h m		0
Mon.	1	-			4 20	32014	17.67
Tues.	2					31024	29.83
Wed.	3					01324	41.99
Thur.	4			Venus at aphelion	1 10	21034	54.14
				Taurid meteors			
Fri.	5					20134	66.29
Sat.	6	10		Moon at apogee (406, 320 km)	22 00	O24*d	78.44
		18	15	Full Moon. Penumbral Eclipse of C pp 54 and 59			
Sun	7	04		Mercury in superior conjunction		31024	90.58
Sum	·	20		Jupiter 1° N. of Moon. Occ'n.		01021	20100
Mon.	8			Mercury at descending node		32041	102.73
Tues.	9				18 40	3410*	114.87
Wed.	10					40312	127.02
Thur.	11					412O3	139.17
Fri.	12				15 30	42013	151.32
Sat.	13					41O32	163.48*
Sun.	14	10		Saturn 6° N. of Moon		43O2d	175.64
		17	39	Last Quarter			
Mon.	15				12 20	43 2O1	187.81 ¹
Tues.	16	19		Leonid meteors*		34120	199.99
Wed.	17					O4312	212.17
Thur.	18			Mercury at aphelion	9 10	12043	224.35
		03		Jupiter at opposition			
Fri.	19	00		Spica 1° S. of Moon. Occ'n.		20134	236.54
Sat.	20	01		Uranus 0.05° S. of Moon. Occ'n.		10324	248.74
		20		Moon at perigee (357,490 km)			
Sun.	21	10	11	New Moon	6 00	30124	260.94
Mon.	22	21		Vesta stationary		32014	273.14
Tues.	23					31204	285.34
Wed.	24	08		Venus 7° S. of Moon	2 50	O31 42	297.53
		20		Mars in conjunction with Sun			
Thur.	25	10		Mercury 3° S. of Neptune		12403	309.72
Fri.	26				23 40	42013	321.91°
Sat.	27			Venus at greatest hel. lat. S.		41032	334.09
Sun.	28	02		Saturn stationary		43012	346.26
		07	59	First Quarter		4220*	250 42
Mon.	29				20 30	4320*	358.43
Tues.	30					43120	10.59

¹Nov. 15, -7.21°; Nov. 27, +7.81°.

^bNov. 13, +6.78°; Nov. 26, -6.75°.

*The behaviour of this shower is rather unpredictable, but it is not expected to be conspicuous.

THE SKY FOR DECEMBER 1976

The winter solstice is on the 21st at 12.36 E.S.T. Let us also write down the dates and times of the vernal equinox, summer solstice and autumnal equinox, and let us also toss in the date and time of last year's winter solstice from the 1975 Handbook (Dec. 22 6.40 E.S.T.). Now let us calculate the intervals WS to VE, VE to SS, SS to AE and AE to WS. They are not equal, are they? Why not? Historical note: this is precisely how Hipparchus in the 2nd century B.C. calculated fairly accurately the time of the closest approach of the sun to the earth, although he believed that the sun moved around the earth on an eccentric circle, whereas we know that it is the earth which moves about the sun on an ellipse.

Look back to the note about the diurnal path of June's full moon and answer the same question about December's full moon.

The Sun—During December the sun's R.A. increases from 16 h 29 m to 18 h 46 m and its Decl. changes from 21° 47' S. to 23° 02' S. The equation of time changes from +10 m 46 s to -3 m 15 s, being zero on the 25th. The winter solstice occurs on the 21st at 12 h 36 m E.S.T.

Mercury on the 1st is in R.A. 17 h 24 m, Decl. 25° 16' S., and on the 15th is in R.A. 18 h 54 m, Decl. 25° 02' S. Greatest eastern elongation is on the 20th, but this is an unfavourable one, Mercury being only about 10 degrees above the south-western horizon at sunset.

Venus on the 1st is in R.A. 19 h 27 m, Decl. 24° 13' S., and on the 15th it is in R.A. 20 h 38 m, Decl. 20° 48' S., mag. -3.7, and transits at 15 h 03 m. It is quite prominent in the south-western sky for about three hours after sunset

Mars on the 15th is in R.A. 17 h 05 m, Decl. 23° 18' S., mag. + 1.6, and transits at 11 h 29 m. Though now in the morning sky it is too close to the sun for observation.

Jupiter on the 15th is in R.A. 3 h 23 m, Decl. 17° 28' N., mag. -2.3, and transits at 21 h 43 m. Moving from Taurus westward into Aries, it is now well up in the east at sunset and is visible until nearly dawn. At about 19 hours on the 4th Jupiter is less than a degree north of the moon.

Saturn on the 15th is in R.A. 9 h 17 m, Decl. 16° 39' N., mag. +0.4, and transits at 3 h 41 m. Near the boundary between Cancer and Leo, it rises in the late evening.

Uranus on the 15th is in R.A. 14 h 31 m, Decl. 14° 26' S., and transits at 8 h 54 m.

Neptune on the 15th is in R.A. 16 h 51 m, Decl. 21° 03' S., and transits at 11 h 14 m.

1 97 6				DECEMBER E.S.T.	Min. of Algol	Config. of Jupiter's Sat. 23 h E.S.T.	Sun's Selen. Colong. 0 h U.T.
	d	h	m		h m		o
Wed.	1					41023	22.74
Thur.	2	1			17 20	24013	34.89
Fri.	3	13		Moon at apogee (405,930 km)		10234	47.04
Sat.	4	19		Jupiter 0.8° N. of Moon. Occ'n. ¹		30124	59.18
Sun.	5	12		Neptune in conjunction with Sun	14 10	32104	71.31
Mon.	6	13	15	😨 Full Moon		32O4d	83.45
Tues.	7					30124	95.58
Wed.	8				11 00	10234	107.71
Thur.	9			Mercury at greatest hel. lat. S.		20143	119.84
Fri.	10					1043*	131.98
Sat.	11	16		Saturn 6° N. of Moon	740	43012	144.12 ^b
Sun.	12					43210	156.26
Mon.	13	16		Geminid meteors		432O1	168.41 ¹
Tues.	14	05	14	Last Quarter	4 30	43O2*	180.56
Wed.	15					41023	192.72
Thur.	16	09		Spica 1° S. of Moon. Occ'n.		42013	204.89
Fri.	17	14		Uranus 0.3° S. of Moon. Occ'n.	1 20	4103*	217.07
Sat.	18					43012	229.25
Sun.	19	07		Moon at perigee (360, 950 km)	22 10	31204	241.43
Mon.	20	05		Mercury greatest elong. E. (20°)		32014	253.62
		21	08	🕲 New Moon			l.
Tues.	21	12	36	Solstice. Winter begins		31024	265.81
Wed.	22	06		Ursid meteors	19 00	10234	278.00
	1	10		Mercury 6° S. of Moon			
Thur.	23					20134	290.19 ^b
Fri.	24	10		Venus 7° S. of Moon		12O34	302.38
Sat.	25	00		Pallas stationary	15 50) O124d	314.56 ⁱ
Sun.	26					31204	326.73
Mon.	27	17		Mercury stationary		32041	338.90
Tues.	28			Mercury at ascending node	12 40	43102	351.07
		02	48	First Quarter			
Wed.	29	l				40 32 d	3.23
Thur.	30	1				42013	15.38
Fri.	31	04		Moon at apogee (405,100 km)	9 30	412O3	27.53

¹Dec. 13, -6.23° ; Dec. 25, $+7.01^{\circ}$. ^bDec. 11, $+6.65^{\circ}$; Dec. 23, -6.59° . ¹Visible in S. America.

Date	Р	B ₀	Lo	Date	Р	B ₀	Lo
	0	o	0		0	0	0
Jan. 1 6 11 16 21	$\begin{array}{r} + 2.44 \\ + 0.01 \\ - 2.40 \\ - 4.77 \\ - 7.08 \end{array}$	-2.97 -3.55 -4.09 -4.61 -5.09	148.62 82.77 16.93 311.09 245.25	July 4 9 14 19 24	$ \begin{array}{r} -1.27 \\ +1.00 \\ +3.24 \\ +5.45 \\ +7.60 \end{array} $	+3.25 +3.78 +4.28 +4.75 +5.19	226.94 160.77 94.60 28.44 322.29
26 31 Feb. 5 10 15 20	$ \begin{array}{r} -9.32 \\ -11.45 \\ -13.48 \\ -15.38 \\ -17.16 \\ -18.79 \end{array} $	$ \begin{array}{r} -5.52 \\ -5.92 \\ -6.27 \\ -6.56 \\ -6.81 \\ -7.00 \end{array} $	179.42 113.59 47.76 341.92 276.09 210.24	Aug. 3 8 13 18 23	$\begin{array}{r} + 9.68 \\ + 11.68 \\ + 13.58 \\ + 15.38 \\ + 17.07 \\ + 18.64 \end{array}$	+5.59 +5.96 +6.28 +6.56 +6.80 +6.99	256.15 190.02 123.90 57.80 351.71 285.63
25 Mar. 1 6 11 16 21	-20.28 -21.61 -22.79 -23.81 -24.66 25.23	-7.14 -7.22 -7.25 -7.22 -7.14 7.00	144.39 78.53 12.67 306.78 240.88	28 Sept. 2 7 12 17	+20.08 +21.39 +22.56 +23.58 +24.45	+7.13 +7.21 +7.25 +7.23 +7.16	219.57 153.52 87.48 21.46 315.44
21 26 31 Apr. 5 10 15 20 25	-25.33 -25.84 -26.17 -26.32 -26.29 -26.08 -25.68	-6.81 -6.57 -6.28 -5.95 -5.57 -5.15	174.97 109.04 43.09 337.12 271.13 205.12 139.08	227 Oct. 2 7 12 17 22	+25.16 +25.71 +26.09 +26.29 +26.32 +26.16 +25.80	+7.04 +6.87 +6.65 +6.37 +6.05 +5.68 +5.27	249.44 183.46 117.48 51.50 345.54 279.59 213.64
25 30 May 5 10 15 20	$ \begin{array}{r} -25.10 \\ -24.33 \\ -23.38 \\ -22.26 \\ -20.97 \\ -19.51 \\ \end{array} $	$ \begin{array}{r} -4.70 \\ -4.22 \\ -3.71 \\ -3.17 \\ -2.61 \\ -2.03 \end{array} $	73.03 6.96 300.88 234.77 168.65 102.51	27 Nov. 1 6 11 16 21	+25.26 +24.52 +23.58 +22.44 +21.12 +19.61	+4.82 +4.33 +3.81 +3.25 +2.67 +2.07	147.70 81.77 15.84 309.92 244.00 178.09
25 30 June 4 9 14 19 24 29	$ \begin{array}{r} -17.90 \\ -16.15 \\ -14.26 \\ -12.27 \\ -10.18 \\ -8.02 \\ -5.80 \\ -3.54 \end{array} $	$\begin{vmatrix} -1.44 \\ -0.85 \\ -0.24 \\ +0.36 \\ +1.55 \\ +2.14 \\ +2.70 \end{vmatrix}$	36.36 330.20 264.03 197.86 131.67 65.49 359.31 293.12	26 Dec. 1 6 11 16 21 26 31	+17.92 +16.06 +14.06 +11.93 + 9.69 + 7.37 + 4.98 + 2.56	$ \begin{array}{r} +1.45 \\ +0.82 \\ +0.18 \\ -0.46 \\ -1.10 \\ -1.73 \\ -2.34 \\ -2.94 \end{array} $	112.19 46.30 340.41 274.52 208.64 142.78 76.91 11.06

SUN—EPHEMERIS FOR PHYSICAL OBSERVATIONS, 1976 For 0 h U.T.

P—is the position angle of the axis of rotation, measured eastward from the north point on the disk, B_0 is the heliographic latitude of the centre of the disk, and L_0 is the heliographic longitude of the centre of the disk, from Carrington's solar meridian, measured in the direction of rotation.

Carrington's Rotation Numbers—Greenwich Date of Commencement of Synodic Rotation, 1976

No.	Commences	No.	Commences	No.	Commences
1637	Jan. 12.29	1642	May 27.75	1647	Oct. 10.90
1638	Feb. 8.63	1643	June 23.95	1648	Nov. 7.20
1639	Mar. 6.96	1644	July 21.15	1649	Dec. 4.51
1640	Apr. 3.27	1645	Aug. 17.37	1650	Dec. 31.84
1641	Apr. 30.53	1646	Sept. 13.63		

SUN-SPOTS

The diagram compares the sun-spot activity for cycles 19, 20 (immediately past) and 21 (beginning in 1975), with the mean of that for cycles 8 to 19. Sun-spot activity was expected to reach a minimum by the end of 1975. The first sun-spot of cycle 21 was observed by Waldmeier on Nov. 15, 1974 at solar latitude 37° N.



ECLIPSES DURING 1976

In 1976 there will be four eclipses, two of the sun and two of the moon.

1. An annular eclipse of the sun on April 29, beginning in mid Atlantic, tracking across north-west Africa, southern Asia and ending in China. It is visible as a partial eclipse along the shores of the St. Lawrence from about Montreal eastward and throughout the Atlantic provinces and the New England states. Mid-eclipse will be at about 5:30 a.m. A.S.T. in these regions and the duration will be about 40 minutes.

2. A partial eclipse of the moon on May 13, not visible in North America.

3. A total eclipse of the sun on October 23, beginning in east Africa, tracking across the Indian Ocean and the extreme southern part of Australia and ending in the south Pacific.

4. A penumbral eclipse of the moon on the night of Nov. 6, visible in part in North America except in the north-western section.

Moon enters penumbra	15.46 E.S.T.
Middle of eclipse	18.01 E.S.T.
Moon leaves penumbra	20.17 E.S.T.
Penumbral magnitude of eclipse 0.86	

PLANETARY APPULSES AND OCCULTATIONS

A planetary appulse is a close approach of a star and a solar system object, as seen from the earth. According to Gordon E. Taylor, of H.M. Nautical Almanac Office, the following appulses will occur in 1976, and may be of interest to observers. The geocentric separation, in declination, is given in the sense planet *minus* star. The horizontal parallax is the angle subtended at the planet by the earth's equatorial radius. Times are given in U.T.; to get E.S.T., subtract 5 hours.

Planet	Date	UT of conjunction	Star Name or SAO No.	Vis. Mag.	Geocentric Separation	Horizontal Parallax
Venus	Jan. 21 Feb. 4 Nov. 25 Mar. 10	h m 22 39 14 01 13 05 22 04	185584 187342 187562 77550	6.7 6.2 6.4	-13.1 - 9.6 -11.4	7.2 6.8 7.9
Jupiter Neptune Ceres Pallas	Apr. 8 Mar. 25 Nov. 15 Nov. 3 July 29 Aug. 9	00 56 23 50 15 31 16 05 17 06 23 18	ε Gem 92688 184653 99608 132592 132996	8.5 3.2 7.0 8.7 9.0 8.5 9.0	$ \begin{array}{r} -9.8 \\ + 2.0 \\ -16.0 \\ - 1.3 \\ + 1.8 \\ + 2.2 \\ + 0.5 \\ \end{array} $	5.9 1.5 0.3 2.9 3.1 3.1
Juno Vesta	Feb. 27 Mar. 5 Mar. 11 Aug. 8 Mar. 10 July 17	01 47 01 50 00 18 21 03 20 24 08 53	133644 118514 118449 118410 119163 110087 94517	0.9 9.2 6.3 7.6 8.3 8.8 8.4	+ 1.6 - 2.1 + 6.8 + 5.6 - 4.2 - 2.6 - 0.2	5.3 5.2 5.2 2.4 2.7 2.6

Three of the appulses give rise to observable occultations: that of SAO 77347 by Vesta is visible from western Australia, that of SAO 153844 by Pallas is possibly visible from part of western North America. Detailed predictions will be issued by Mr. Taylor at a later date. The occultation of ε Gem by Mars is widely visible, as follows:

	Disappe	arance	Reappe	arance	Altitude of		
Place	UT	Р	UT	Р	Star	Sun	
Hawaii Palomar, Calif. Rothney, Alta. McDonald Obs., Texas Toronto, Ont. Washington, D.C. Bermuda	0 ^h 52 ^m 2 0 ^h 54 ^m 5 0 ^h 54 ^m 6 0 ^h 55 ^m 5 0 ^h 56 ^m 6 0 ^h 56 ^m 9 0 ^h 57 ^m 8	134° 115° 84° 119° 87° 94° 98°	0 ^h 56 ^m 3 0 ^h 59 ^m 4 0 ^h 59 ^m 6 1 ^h 00 ^m 3 1 ^h 01 ^m 7 1 ^h 02 ^m 0 1 ^h 02 ^m 9	237° 256° 288° 253° 284° 278° 273°	47° 80° 64° 82° 59° 59° 59°	$ \begin{array}{c c} +52^{\circ} \\ +14^{\circ} \\ +13^{\circ} \\ +3^{\circ} \\ -12^{\circ} \\ -28^{\circ} \end{array} $	

OCCULTATIONS BY THE MOON

The moon often passes between the earth and a star; the phenomenon is called an occultation. During an occultation a star suddenly disappears as the east limb of the moon crosses the line between the star and observer. This is referred to as immersion (I). The reappearance from behind the west limb of the moon is called emersion (E). Because the moon moves through an angle about equal to its own diameter every hour, the longest time for an occultation is about an hour. The time can be shorter if the occultation is not central. Occultations are equivalent to total solar eclipses, except that they are total eclipses of stars other than the sun.

The elongation of the moon is its angular distance from the sun, in degrees, counted eastward around the sky. Thus, elongations of 0° , 90° , 180° and 270° correspond to new, first quarter, full and last quarter moon. When elongation is less than 180° , a star will disappear at the dark limb and reappear at the bright limb. If the elongation is greater than 180° the reverse is true.

As in the case of eclipses, the times of immersion and emersion and the duration of the occultation are different for different places on the earth's surface. The tables given below, are adapted from data supplied by the British Nautical Almanac Office and give the times of immersion or emersion or both for occultations visible from six stations distributed across Canada. Stars of magnitude 7.5 or brighter are included as well as daytime occultations of very bright stars and planets. Since an occultation at the bright limb of the moon is difficult to observe the predictions are limited to phenomena occurring at the dark limb.

The terms *a* and *b* are for determining corrections to the times of the phenomena for stations within 300 miles of the standard stations. Thus if λ_0 , ϕ_0 , be the longitude and latitude of the standard station and λ , ϕ , the longitude and latitude of the neighbouring station then for the neighbouring station we have: Standard Time of phenomenon = Standard Time of phenomenon at the standard station + $a(\lambda - \lambda_0)$ + $b(\phi - \phi_0)$ where $\lambda - \lambda_0$ and $\phi - \phi_0$ are expressed in degrees. This formula must be evaluated with due regard for the algebraic signs of the terms. The quantity *P* is the position angle of the point of contact on the moon's disk reckoned from the north point towards the east.

Since observing occultations is rather easy, provided the weather is good and the equipment is available, timing occultations should be part of any amateur's observing program. The method of timing is as follows: Using as large a telescope as is available, with a medium power eyepiece, the observer starts a stopwatch at the time of immersion or emersion. The watch is stopped again on a time signal from a WWV or CHU station. The elapsed time is read from the stopwatch and is then subtracted from the standard time signal to obtain the time of occultation. All times should be recorded to 0.1 second and all timing errors should be held to within 0.5 second if possible. The position angle P of the point of contact on the moon's disk reckoned from the north point towards the east may also be estimated.

The following information should be included: (1) Description of the star (catalogue number), (2) Date, (3) Derived time of the occultation, (4) Longitude and latitude to nearest second of arc, height above sea level to the nearest 100 feet, (5) Seeing conditions, (6) Stellar magnitude, (7) Immersion or emersion, (8) At dark or light limb; Presence or absence of earthshine, (9) Method used, (10) Estimate of accuracy, (11) Anomalous appearance: gradual disappearance, pausing on the limb. All occultation data should be sent to the world clearing house for occultation data: H.M. Nautical Almanac Office, Royal Greenwich Observatory, Herstmonceux Castle, Hailsham, Sussex, England.

The co-ordinates of the standard stations are given in the tables.

			I	Elong.	w.	HALIF 63°6, N	AX 1. 44:6		W.	10NTR 73°6, N	EAL 1. 45:5	
Date	Z.C. No.	Mag.	or E	of Moon	A.S.T.	a	b	Р	E.S.T.	a	b	P
Jan. 5 7 14 14 14/5	3290 3512 760 888 895	7.3 5.8 6.5 6.0 5.9	I I I I I	。 51 72 143 154 154	h m Low 17 26.3 4 00.4 22 51.6 0 25.6	m -1.7 +0.2 -1.8 -1.4	m +0.9 -1.6 -0.6 -1.5	° 57 111 100 109	h m 19 35.4 Sun 3 00.4 21 32.8 23 09.0	m - 1.0 + 0.1 - 1.8 - 1.6	m -2.8 -1.9 -0.1 -1.4	。 113 119 99 113
15 15 21 Feb. 6 8	913 913 1688 299 517	5.2 5.2 6.3 6.3 6.4	I E I I I	156 156 235 76 97	Graze 5 05.2 No occ. Low 18 58.5		_	2 144	3 46.4 No occ. 2 15.6 22 13.5 Sun	-0.9 -0.5	+0.3 +0.2	40 9 40
10 11/2 17 18 19	798 985 1623 1759 1888	6.4 6.9 5.4 6.5 6.2	I I E E E	120 134 202 217 231	18 31.4 0 11.0 1 26.6 3 42.0 2 42.8	-1.9 -0.4 -0.7 -0.6 -0.4	$ \begin{array}{c} -0.3 \\ -3.3 \\ -2.2 \\ -2.6 \\ -2.2 \end{array} $	115 151 342 350 349	Sun 23 04.0 0 16.6 2 32.7 1 35.5	-0.9 -0.9 -0.7	-1.4 -1.8 -1.2	163 329 336 335
22 22 22 Mar. 5 8	2302 2302 2303 374 760	2.9 2.9 5.1 6.1 6.5	I E I I I	271 271 271 56 90	2 59.2 3 19.4 Graze 21 26.9 21 51.0	 	 	30 357 50 29	Low Low 20 21.4 20 32.7	-0.7 -2.0	$^{-0.3}_{+1.5}$	53 40
10 12 12 16 21	1057 1309 1318 1815 Nept.	6.9 5.7 5.7 4.8 7.7	I I E E	114 138 139 196 254	22 20.9 18 50.7 20 58.3 23 58.2 07 04.3	-1.0 -1.5 -1.7 -1.7 -1.7	$\begin{array}{c} -2.2 \\ +3.5 \\ -0.5 \\ +0.5 \\ +0.9 \end{array}$	128 52 111 271 213	21 08.3 Sun 19 42.3 22 42.5 05 51.9	-1.1 -1.5 -1.6 -1.5	-2.4 -0.2 +1.3 +0.7	136 115 260 221
Apr. 3 5 5 7 8	593 862 863 1147 1256	5.8 7.5 6.7 5.1 7.1	I I I I I	48 71 71 95 106	Low 21 43.0 21 52.0 23 19.7 19 22.5		- <u>-</u> - <u>2.6</u> -0.5	176 169 147 103	21 18.3 No occ. No occ. 22 17.6 Sun	+0.3	-1.9 -3.1	116 158
9 17 17 20 20	1384 2343 2353 2826 2826	7.4 6.4 4.6 4.0 4.0	I E E I E	120 220 221 260 260	No occ. 2 17.0 4 36.7 4 19.9 Sun	- <u>1.8</u> 	+0.3	266 339 16	20 00.2 1 00.3 3 15.5 3 07.6 3 36.6	- <u>1.6</u>	+0.8	50 261 346 11 329
20 24 May 4 5 9	2828 3320 1106 1237 1587	6.0 5.3 3.6 6.4 6.0	E E I I I	260 306 65 77 117	Sun 4 09.3 22 28.3 23 05.4 Low	-0.7 + 0.1 - 0.3	+0.7 -1.7 -0.9	301 116 66	3 37.9 Low 21 27.0 21 59.9 0 35.2	-1.6 -0.1 -0.6	+2.5 -2.0 -1.1	204 123 74 187
9 June 5 7 10	1705 1815 1662 1914 2192	7.5 4.8 6.3 6.8 6.2	I I I I I I	130 142 98 125 153	23 26.2 20 41.2 21 59.2 23 21.8 0 34.7	-0.7 -1.4 -1.0 -1.2	-2.4 -0.4 -1.5 -1.3	156 120 101 101 28	22 17.1 19 28.8 20 46.3 22 07.6 Graze	-0.6 -1.0 -1.3 -1.4	-2.6 -0.3 -1.4 -1.1	167 130 107 104
13 July 2 5 6 7	2828 1623 2002 2136 2275	6.0 5.4 6.8 6.8 5.9	E I I I I	206 67 108 121 134	23 07.6 21 06.3 22 25.0 22 44.8 20 57.7	$-1.3 \\ -0.9 \\ -1.4 \\ -1.9$	$+1.9 \\ -0.6 \\ -1.6 \\ +0.3$	231 58 179 122 79	Low Sun 21 11.8 21 29.1 Sun	-1.5	-1.3	180 121
8 9 14 14 Aug. 4	2296 2456 3185 3187 2394	7.1 6.2 5.3 6.2 6.5	I I E E I	136 149 212 213 117	Low Low 2 16.5 2 37.2 22 12.0	$-2.1 \\ -0.7 \\ -1.7$	$^{+0.1}_{+2.3}_{-1.7}$	270 194 126	0 15.8 0 38.6 0 55.4 1 29.2 20 53.7	-1.3 -0.8 -2.1 -1.2 -1.7	$-2.1 \\ -0.2 \\ +0.2 \\ +1.8 \\ -1.2$	132 52 284 211 120
6/7 13 28 28 Sept. 3	2715 3515 1925 1925 2826	6.5 6.2 1.2 1.2 4.0	I E I E I	144 217 46 46 125	0 34.1 Sun 19 27.5 Low 23 02.6	-1.0 -0.7 -1.2	-0.4 -2.0 -0.5	64 131 72	23 23.3 2 46.0 18 17.7 19 16.9 21 49.5	-1.1 -2.2 -0.9 -0.6 -1.3	$^{+0.1}_{-0.4}_{-1.9}_{-1.2}_{0.0}$	52 273 129 264 59

LUNAR OCCULTATIONS VISIBLE AT HALIFAX AND MONTREAL, 1976

			I	Elong.	w.	HALIF. 63:6, N	AX . 44:6		W.	10NTR 73°6, N	EAL . 45°5	
Date	Z.C. No.	Mag.	Or E	Moon	A.S.T.	a	b	Р	E.S.T.	a	b	Р
Sept. 5 5 11 29 Oct. 3	2968 2969 272 2611 3184	6.2 3.2 5.9 6.8 7.1	I I E I I	。 139 139 218 82 132	h m Low 23 03.4 20 15.2 23 26.8	m -0.6 -1.2 -1.5	m +2.3 -0.6 -0.9	。 229 71 84	h m 0 53.8 1 00.4 22 00.2 19 02.4 22 11.1	m -0.6 -0.4 -1.3 -1.5		。 55 59 240 61 69
3 11 15 16 28	3185 590 1029 1147 2871	5.3 6.3 5.1 5.1 7.1	I E E I I	132 220 255 267 76	23 30.2 22 11.8 3 13.3 2 32.2 18 29.0	$ \begin{array}{r} -1.7 \\ 0.0 \\ -1.5 \\ -0.8 \\ -1.5 \end{array} $	-1.4 + 2.7 + 2.8 + 3.0 + 0.1	96 218 234 233 62	22 12.3 21 14.3 2 02.9 1 28.3 Sun	-1.7 +0.1 -1.1 -0.5	$ \begin{array}{c} -0.4 \\ +2.2 \\ +2.4 \\ +2.4 \end{array} $	81 230 242 242
30 Nov. 2 9/0 11/2	3154 3515 730 832 1106	7.4 6.2 5.1 4.7 3.6	I E E I	102 136 204 213 236	23 12.4 23 11.6 No occ. 0 12.1 1 08.0	-0.4 -1.4 -1.4	+0.1 + 0.4	45 55 195 132	22 08.6 21 58.7 5 46.5 23 11.4 23 55.2	-0.4 -1.3 -0.6 -1.1	+0.7 +1.2 +3.8 +0.2	31 42 346 214 124
12 13 14 26 Dec. 2	1106 1237 1359 3093 272	3.6 6.4 5.1 4.5 5.9	E E I I I	236 249 262 70 137	2 14.9 4 04.2 Sun 20 12.0 17 12.2	-1.7 -2.1 $-\overline{0.7}$	+2.2 + 1.2 +1.2	244 255 1 72	1 01.9 2 46.0 5 16.6 No occ. Sun		+2.1 +1.7 -1.2	249 253 305
2 4 10 11 11	284 422 1309 1318 1332	7.4 5.5 5.7 5.7 5.7	I E E E	138 151 229 230 232	19 37.1 2 30.7 22 55.4 1 32.2 Sun	-1.6 -1.0 -0.9 -1.7	+1.3 +0.4 -1.9 +1.5	76 42 340 259	18 24.9 1 20.2 21 43.0 0 19.0 5 57.1	$ \begin{array}{r} -1.1 \\ -1.2 \\ -1.2 \\ -0.4 \end{array} $	+1.9 +0.6	62 43 359 259 338
26 29/0 31 31	3420 264 380 469	7.1 7.0 7.4 7.3	I I I I	73 109 120 129	17 34.5 0 40.6 1 42.0 18 47.8	$-0.5 \\ -0.3 \\ 0.0 \\ -2.5$	+3.0 -1.3 -2.4 -0.8	8 84 118 124	No occ. 23 34.7 0 38.3 17 28.1	$-0.6 \\ -0.3 \\ -1.5$	$-1.3 \\ -2.8 \\ +0.7$	84 121 106

LUNAR OCCULTATIONS VISIBLE AT TORONTO AND WINNIPEG, 1976

	7.0	Z.C. I Elong or of			TORONTO W. 79°4, N. 43°7				WINNIPEG W. 97?2, N. 49?9			
Date	Z.C. No.	Mag.	E	Moon	E.S.T.	a	b	Р	C.S.T.	a	b	Р
Jan. 5 14 14 14 15	3290 760 888 895 913	7.3 6.5 6.0 5.9 5.2	I I I I I	° 51 143 154 154 154 156	h m 19 33.7 3 04.7 21 22.2 23 01.8 3 41.6		m -3.0 -2.3 0.0 -1.6 -0.3	° 116 129 104 122 56	h m 18 00.2 1 46.7 20 01.4 21 26.8 2 20.7			。 74 129 77 102 59
21 27 Feb. 3 5 6	1688 2509 3494 186 299	6.3 6.0 4.6 7.1 6.3	E E I I I	235 315 42 65 76	2 20.7 6 20.5 19 58.6 22 02.4 22 10.1	$ \begin{array}{c} -0.5 \\ -1.8 \\ -0.5 \\ 0.0 \\ -0.6 \end{array} $	-2.1 + 2.5 - 1.9 - 3.5 - 0.1	346 225 98 130 50	1 02.1 No occ. 18 37.2 20 37.1 20 57.4	-0.3 -0.9 -0.8 -1.0	$ \begin{array}{c} -1.6 \\ -0.7 \\ -2.2 \\ +1.0 \end{array} $	346 69 105 31
7 8 11 16/7 18	423 434 985 1623 1759	6.4 6.9 6.9 5.4 6.5	I I E E	88 88 134 202 217	23 45.3 No occ. No occ. 0 12.6 2 29.6	-1.1 -1.1	-0.8 -1.3	7 317 323	No occ. 0 15.5 21 26.2 22 53.4 1 06.0	$ \begin{array}{c} 0.0 \\ -0.7 \\ -0.9 \end{array} $	-1.6 -0.3 -0.4	93 167 319 316
19 Mar. 5 7 8 10	1888 374 639 760 1057	6.2 6.1 6.0 6.5 6.9	E I I I I	231 56 80 90 114	1 32.6 20 17.8 No occ. 20 19.4 21 06.6	-0.9 -0.8 -2.0 -1.1	$-0.6 \\ -0.5 \\ +0.8 \\ -3.2$	321 62 53 149	0 19.0 18 58.8 23 08.3 18 56.0 19 27.1	$-0.4 \\ -1.2 \\ +0.1 \\ -1.5$	$\begin{array}{c} 0.0 \\ +0.5 \\ -2.6 \\ -\overline{1.7} \end{array}$	318 45 130 32 137
10/1 12 16 17 21	1072 1318 1815 1853 Nept.	6.2 5.7 4.8 4.9 7.7	I E E E	115 139 196 200 254	0 28.5 19 34.5 22 29.7 No occ. 5 40.3	-1.5 -1.4 -1.8 -1.9	$+0.5 \\ -0.4 \\ +2.3 \\ +1.3$	44 124 244 218	23 01.5 No occ. 21 21.6 5 30.2 4 14.4	-1.8 -0.9 -0.7 -1.9	$ +0.4 \\ +2.3 \\ -1.3 \\ +1.5$	51 248 269 225

	70		I	Elong.	w.	T ORO N 79°4, N.	TTO 43:7		w.	VINNII 97°2, N.	PEG 49:9	
Date	2.C. No.	Mag.	E	Moon	E.S.T.	a	b	Р	C.S.T.	a	b	P
Apr. 2 3 7 9 11	466 593 1147 1384 1528	7.3 5.8 5.1 7.4 6.6	I I I I I	° 38 48 95 120 136	h m No occ. 21 23.8 22 26.4 19 41.4 2 26.4	$ \begin{array}{c} m \\ +0.3 \\ -2.3 \\ -0.6 \end{array} $	m = -2.3 = -1.4 = -0.4	° 127 177 69 55	h m 20 54.6 20 08.9 No occ. No occ. 1 06.0	m + 0.1 - 0.3	m -1.9 -2.6 t	° 108 123 57
11 15 16 17 17	1623 2182 2192 2343 2353	5.4 6.3 6.2 6.4 4.6	I E E E E	148 206 207 220 221	21 35.8 23 18.4 2 36.6 0 48.8 3 11.4	$-\overline{0.1}$ $-\overline{1.7}$	$-\overline{1.5}$ $+\overline{1.3}$	58 344 357 253 337	20 09.2 No occ. 1 10.6 No occ. 1 42.5	-1.7 -0.6	+2.4	63 352 339
20 20 20 May 4 5	2826 2826 2828 1106 1237	4.0 4.0 6.0 3.6 6.4	I E I I I	260 260 260 65 77	$\begin{array}{c} 2 & 51.5 \\ 3 & 29.5 \\ 3 & 23.3 \\ 21 & 30.3 \\ 21 & 58.2 \end{array}$	 	 	21 322 198 132 84	No occ. No occ. No occ. 20 10.8 20 33.6	-0.5 -1.1	$-2.5 \\ -1.2$	137 90
9 16 June 5 7 9	1705 2591 1662 1914 2192	7.5 6.5 6.3 6.8 6.2	I E I I I	130 215 98 125 153	22 20.7 No occ. 20 41.2 22 01.0 23 04.2	-1.4 -1.5	-1.4 -1.1	184 116 111 38	No occ. 2 10.4 No occ. No occ. 21 32.9	-1.5	-0.2	301 42
July 6 6 7/8 7 7 7	2136 2147 2296 2302 2303	6.8 7.0 7.1 2.9 5.1	I I I I I	121 122 136 136 136	21 22.3 No occ. 0 11.7 No occ. No occ.	-1.6 -1.5	-1.2 -2.0	126 132	No occ. 22 37.2 22 36.4 23 46.7 23 46.8	-1.3 -1.5 -1.2 -1.2	-1.4 -0.9 -1.1 -1.1	120 116 97 96
8/9 13/4 14 Aug. 4 6	2456 3185 3187 2394 2715	6.2 5.3 6.2 6.5 6.5	I E I I I	149 212 213 117 144	0 33.3 0 42.7 1 18.8 20 45.2 23 15.7	-1.1 -2.0 -1.3 -1.8 -1.3	0.0+0.4+1.9-1.0+0.3	51 287 214 123 50	23 16.2 23 16.3 0 07.5 No occ. 21 58.6	- <u>1.1</u>	+1.7	25 314 237 22
13 18 19 28 28	3515 577 718 1925 1925	6.2 6.0 6.1 1.2 1.2	E E I E	217 272 284 46 46	2 33.1 No occ. No occ. 18 15.4 19 15.2	-2.4 -1.1 -0.8	-0.2 -1.9 -1.1	276 133 261	No occ. 2 28.3 3 53.8 16 44.7 17 50.1	+0.1 -0.8 -1.2 -1.2	+3.2 +1.7 -1.3 -1.0	205 257 133 268
Sept. 3 4/5 4/5 5 11	2826 2968 2969 2969 2969 272	4.0 6.2 3.2 3.2 5.9	I I E E	125 139 139 139 218	21 40.8 0 50.5 0 57.4 No occ. 21 54.5	-1.5 -0.7 -0.7 -0.3	$+0.3 \\ -0.2 \\ -0.3 \\ +2.0$	57 54 58 241	20 20.5 23 44.0 23 48.4 0 33.0 No occ.	-1.3 -0.1 -0.3 -1.8	$^{+1.4}_{+1.7}_{+1.3}_{-2.5}$	33 11 17 302
15 17 20 29 29	658 943 1332 2611 2629	4.2 6.2 5.7 6.8 6.3	I E I I	252 276 313 82 83	No occ. No occ. No occ. 18 54.0 No occ.	-1.6	+0.1	61	Graze 4 09.1 4 52.8 No occ. 20 30.1	-0.6 -1.3	-0.3 -1.2	332 322 99
Oct. 2 3 3 11	3051 3184 3185 3187 590	7.0 7.1 5.3 6.2 6.3	I I I E	119 132 132 132 220	No occ. 22 01.5 22 02.4 No occ. 21 11.1	$-1.7 \\ -1.9 \\ +0.2$	$^{+0.3}_{-0.1}$ +2.1	66 78 232	18 42.5 20 42.0 20 38.2 21 29.0 No occ.	-1.6 -1.1 -1.3	+0.6 +1.5 +1.3 —	109 34 47 126
15 16 18 27 30	1029 1147 1397 2731 3154	5.1 5.1 5.5 6.5 7.4	E E I I I	255 267 294 64 102	1 52.9 1 21.7 No occ. No occ. 22 04.4	$-0.8 \\ -0.3 \\ -0.6$	+2.5 +2.4 +0.9	239 240 30	0 53.7 0 30.5 5 35.5 19 45.5 No occ.	-0.5 -0.1 -1.2 -0.1	+1.5 +1.4 +0.5 +0.9	269 271 288 20
31 Nov. 2 4 9 9	3281 3515 98 730 832	7.5 6.2 6.2 5.1 4.7	I I E E	114 136 149 204 213	No occ. 21 48.4 3 30.2 5 56.5 23 01.8	$-1.4 \\ -0.3 \\ -0.1 \\ -0.3$	$^{+1.6}_{-0.5}_{-3.7}_{+3.8}$	39 60 325 213	21 03.5 No occ. 2 20.1 4 21.7 22 10.6	-2.0 -0.7 -0.4	-1.0 +0.3 +2.0	103 38 337 250

Z.C. I Elong				Elong.	TORONTO W. 79°4, N. 43°7				WINNIPEG W. 97:2, N. 49:9			
Date	No.	Mag.	E	Moon	E.S.T.	a	b	Р	C.S.T.	a	b	Р
Nov. 9 11 11/2 13 14	836 1106 1106 1237 1359	5.5 3.6 3.6 6.4 5.1	E I E E	° 213 236 236 249 262	h m No occ. 23 49.2 0 51.1 2 32.5 5 08.6	m -0.9 -1.0 -1.6 -1.8	m +0.1 +2.4 +2.4 -0.6	° 127 244 245 294	h m 22 47.3 22 45.1 23 48.9 1 23.7 3 40.2	m -0.4 -0.2 -0.5 -0.8 -1.2	m + 2.7 + 1.3 + 1.4 + 1.5 - 0.1	。 96 273 268 302
24 25 25 30 Dec. 2	2826 2968 2969 64 284	4.0 6.2 3.2 6.6 7.4	I I I I I	45 58 58 118 138	No occ. No occ. No occ. No occ. 18 15.4	-0.9	+2.0	60	18 16.6 19 34.3 19 41.1 23 07.1 17 21.3	$-1.1 \\ -0.7 \\ -0.7 \\ -0.1$	$-1.2 \\ -0.5 \\ -0.6 \\ +2.6$	94 61 64 129 30
3/4 9 9 10/1 11	422 1106 1106 1318 1332	5.5 3.6 3.6 5.7 5.7	I E E E	151 210 210 230 232	1 11.7 No occ. No occ. 0 09.0 5 58.4	-1.4 -1.0 -0.8	+0.4 +2.1 -2.6	51 252 325	23 55.3 6 32.7 7 36.3 23 06.8 4 25.8	$-1.3 \\ -0.6 \\ -0.1 \\ -0.4 \\ -1.2$	+3.0 -1.6 -1.9 +1.3 -1.9	21 97 293 277 320
24 24 24 29 29	3184 3185 3187 252 264	7.1 5.3 6.2 7.4 7.0	I I I I I	51 51 51 108 109	No occ. No occ. No occ. No occ. 23 32.9	-0.8	-1.5	91	19 40.6 19 36.3 20 07.3 18 43.8 22 07.0	$ \begin{array}{r} -0.2 \\ -0.4 \\ -0.9 \\ -2.1 \\ -1.3 \end{array} $	$ \begin{array}{c} +0.7 \\ +0.1 \\ -2.3 \\ -0.1 \\ -0.4 \end{array} $	25 39 109 104 68
30 30/1	272 380	5.9 7.4	I I	110 120	No occ. 0 42.1	-0.4	-3.7	133	0 25.3 23 08.0	$-0.5 \\ -1.3$		82 110

LUNAR OCCULTATIONS VISIBLE AT EDMONTON AND VANCOUVER, 1976

-	7.0	I Elong. or of			EDMONTON W. 113?4, N. 53?6				VANCOUVER W. 123°1, N. 49°2			
Date	Z.C. No.	Mag.	or E	of Moon	M.S.T.	a	b	Р	P.S.T.	a	b	Р
Jan. 3 9 13/4 14 14	3051 241 760 888 895	7.0 6.9 6.5 6.0 5.9	I I I I I	° 28 97 143 154 154	h m 18 00.3 No occ. 0 26.6 18 57.1 20 09.2	m -1.5 -0.9 -0.4 -1.0	m - 3.2 -2.5 +2.4 +1.3	° 125 130 55 85	h m Sun 22 35.2 23 33.6 17 44.2 18 54.3	m -0.9 -0.2 -0.8	m +1.9 +2.2 +1.3	° 19 159 60 91
14/5 20 20 22 25	913 1587 1688 1815 2217	5.2 6.0 6.3 4.8 5.5	I E E E E	156 225 235 250 291	0 57.9 7 05.7 23 53.5 No occ. 5 49.9	-1.4 -1.0 -0.1	0.0 -1.1 -1.4 	63 247 348 210	23 43.8 5 53.0 Low 2 07.7 No occ.	-1.6 -0.4	-0.3 -1.0	84 222 337
Feb. 2 5 6 7 7	3370 186 299 423 434	6.2 7.1 6.3 6.4 6.9	I I I I I	32 65 76 88 88	18 11.4 19 13.3 19 48.5 No occ. 23 06.5	-0.8 -1.3 -0.4	-1.2 -1.2 -1.7	79 88 9 93	Sun 18 04.0 18 26.2 20 03.0 22 10.6	-1.8 -1.2 -0.6	$\begin{vmatrix} -1.2 \\ +2.4 \\ -\overline{2.3} \end{vmatrix}$	97 25 8 111
11 16 17 17 Mar. 3	985 1623 1662 1759 143	6.9 5.4 6.3 6.5 6.8	I E E I	134 202 206 217 35	$\begin{array}{c} 19 & 51.0 \\ 21 & 44.7 \\ 5 & 24.4 \\ 23 & 53.5 \\ 19 & 56.4 \end{array}$	$ \begin{array}{r} -1.4 \\ -0.4 \\ -0.5 \\ -0.6 \\ -0.5 \end{array} $	-1.8 0.0 -2.1 +0.2 +1.5	147 322 334 311 16	Graze Low 4 26.9 22 45.5 18 47.8	-0.8 -0.6 -0.6	-1.9 + 0.7 + 0.4	320 291 37
7 7/8 9 10 10	639 654 796 943 1072	6.0 6.0 6.8 6.2 6.2	I I I I	80 81 93 104 115	21 57.2 0 07.1 Low 0 58.4 21 34.7	$ \begin{array}{c} -0.3 \\ -0.2 \\ +0.1 \\ -1.7 \end{array} $	-3.0 -0.9 -2.1 +0.7	133 58 124 58	Graze 23 09.8 0 29.5 0 09.4 20 16.0	$ \begin{array}{c} -0.2 \\ -0.3 \\ +0.1 \\ -1.8 \end{array} $	$-1.2 \\ -0.5 \\ -2.6 \\ +0.2$	76 50 143 81
11 11 12 17 21	1091 1212 1332 1853 Nept.	6.7 7.1 5.7 4.9 7.7	I I E E	117 129 141 200 254	2 16.5 No occ. No occ. 4 10.7 2 52.9	0.0 -1.1 -1.7	-1.7 -1.1 +2.6	105 269 216	1 24.2 22 47.3 20 45.5 3 01.5 No occ.	$\begin{array}{c} 0.0 \\ -2.0 \\ -1.9 \\ -1.6 \end{array}$	$ \begin{vmatrix} -1.9 \\ +0.5 \\ +1.7 \\ -0.5 \end{vmatrix} $	118 58 66 257

	7.0		I	Elong.	w.	DMON 113:4, N	TON N. 53°6		W.	ANCOU 123:1, N	VER N. 49°2	
Date	No.	Mag.	E	Moon	M.S.T.	a	b	Р	P.S.T.	a	b	Р
Apr. 10 12 15/6 21 21	1528 1662 2192 2969 2969	6.6 6.3 6.2 3.2 3.2 3.2	I I E I E	。 136 152 207 274 274	h m 23 40.7 2 43.3 0 01.1 No occ. No occ.	m -1.5 -0.8 -0.3	m -0.5 -0.9 -0.6	° 70 57 338	h m 22 28.0 1 38.5 22 57.3 3 01.8 3 26.3	m -1.6 -1.1 -0.6 	m -0.6 -1.0 +0.1 	。 91 71 315 6 327
May 7 9 10 10 11	1482 1605 1726 1727 1853	6.3 6.2 6.9 7.1 4.9	I I I I I	105 119 132 132 146	No occ. Low 1 01.0 No occ. 1 03.2	-0.5 -0.8	-2.2 -1.6	160 127	22 53.5 0 59.2 0 07.8 0 18.8 0 01.2	-1.4 -0.3 -1.3 -1.0	$ \begin{array}{r} -0.3 \\ -\overline{3.0} \\ -0.3 \\ -1.7 \end{array} $	53 182 175 51 137
14 15/6 17 31 July 5	2302 2591 2774 1091 2021	2.9 6.5 6.3 6.7 6.7	I E I I	190 215 229 37 109	Sun 0 50.4 Sun Low 22 16.6	-1.0	+0.5 -0.4	303 49	3 45.9 23 38.3 1 22.0 20 58.3 21 04.3	$-0.9 \\ -0.9 \\ -0.0 \\ -1.6$	$-0.9 \\ +0.8 \\ -1.2 \\ -0.2$	76 291 199 77 61
7 7 7 8 18	2302 2302 2303 2456 146	2.9 2.9 5.1 6.2 4.4	I E I I I	136 136 136 149 260	22 22.4 23 34.0 22 22.6 Graze Sun	-1.4 -1.2 -1.4	-0.4 -1.0 -0.4	91 281 91	21 09.2 22 24.8 21 09.3 Sun 2 26.4	-1.6 -1.5 -1.6 -1.2	-0.1 -0.7 -0.1 +1.3	99 277 98 93
18 Aug.6/7 18 19 28	146 2733 577 718 1925	4.4 6.4 6.0 6.1 1.2	E I E I I	260 145 272 284 46	Sun 0 07.0 1 37.7 2 49.0 15 23.1	$ \begin{array}{r} -0.9 \\ 0.0 \\ -0.5 \\ -1.0 \end{array} $	-0.2 + 2.3 + 1.4 - 0.8	49 231 279 141	3 34.2 22 56.5 0 29.0 1 39.2 14 19.7	$-0.9 \\ -1.2 \\ +0.2 \\ -0.3 \\ -0.6$	+2.3 +0.3 +2.1 +1.3 -1.5	216 49 231 278 162
28 Sept. 13 14 14/5 17	1925 422 658 658 934	1.2 5.5 4.2 4.2 6.4	E E E E	46 232 252 252 252 275	$\begin{array}{c} 16 & 25.2 \\ 3 & 53.3 \\ 23 & 51.5 \\ 0 & 39.5 \\ 1 & 41.6 \end{array}$	-1.5 -2.2 -0.4 0.0 +0.5	$-0.4 \\ -2.3 \\ +1.0 \\ +2.7 \\ +4.1$	263 309 120 219 204	15 07.9 2 38.7 22 44.6 23 29.4 0 30.7	$ \begin{array}{c} -2.1 \\ -2.4 \\ -0.2 \\ +0.3 \\ \end{array} $	+0.7 -1.2 +0.8 +2.5	245 301 121 218 197
20 27 28 29 Oct. 3	1332 2322 2465 2629 3184	5.7 4.3 7.4 6.3 7.1	E I I I I	313 57 70 83 132	3 41.5 Low 19 04.7 19 05.5 19 34.9	-0.5 -1.6 -1.4 -0.6	-1.8 -1.9 -0.4 +2.5	348 144 84 10	No occ. 18 30.3 Sun Sun Sun	-1.3	-1.6	120
3 3 13 14 18	3185 3187 764 1029 1397	5.3 6.2 5.0 5.1 5.5	I I E E E	132 132 235 255 294	19 26.4 19 55.5 4 42.8 23 53.4 4 23.0	$-0.9 \\ -1.5 \\ -1.6 \\ -0.2 \\ -0.7$	+1.9 +0.7 +0.2 +1.1 +0.7	28 99 256 289 295	18 07.9 18 37.1 3 23.0 Low 3 13.2	-0.9 -1.4 -1.7 -0.5	+2.3 +1.0 +1.4 +1.1	29 99 243 282
27 27 31 Nov.1/2 3/4	2733 2745 3281 3420 98	6.4 6.9 7.5 7.1 6.2	I I I I	64 65 114 127 149	18 43.0 Low 19 34.6 0 55.0 1 12.2	-1.5 -0.8 -0.7	+0.6 -2.1 +1.9	149 78 103 16	Sun 19 14.8 18 16.0 23 54.9 23 56.7	-0.9 -1.5 -1.4 -1.0	-0.6 + 1.1 - 2.6 + 1.6	65 76 112 27
9 9 9 11 12/3	730 832 836 1106 1237	5.1 4.7 5.5 3.6 6.4	EEEE	204 213 213 236 249	No occ. 21 13.1 21 50.8 22 47.3 0 19.2	-0.1 + 0.2 - 0.2 - 0.4	$^{+1.6}_{+2.0}_{+1.1}_{+1.2}$	270 252 292 284	1 48.3 20 06.4 20 41.7 Low 23 11.0	+0.1 +0.1 -0.1	+1.4 +1.8 +1.3	324 269 250 277
14 24 25 25 25	1359 2826 2968 2969 2969	5.1 4.0 6.2 3.2 3.2	E I I E	262 45 58 58 58	2 24.6 16 54.5 18 21.9 18 28.0 19 27.7	$-0.8 \\ -1.2 \\ -0.7 \\ -0.7 \\ -1.1$	$+0.1 \\ -0.4 \\ +0.3 \\ +0.1 \\ -1.6$	311 73 36 40 281	1 16.2 Sun 17 11.6 17 17.9 18 21.4	-0.6 -1.0 -1.0 -1.5	+0.6 +0.8 +0.6 -1.3	297 35 39 281
28 30 Dec. 9 9 10	3366 64 1106 1106 1318	6.6 6.6 3.6 3.6 5.7	I I E E	95 118 210 210 230	$\begin{array}{c} 21 & 08.9 \\ 21 & 30.2 \\ 5 & 13.3 \\ 6 & 23.4 \\ 22 & 06.6 \end{array}$	$-1.5 \\ -1.7 \\ -1.0 \\ -0.6 \\ -0.1$	$-1.7 \\ -0.9 \\ -1.6 \\ -1.8 \\ +1.0$	104 97 104 286 294	19 58.8 20 14.9 4 10.5 5 22.4 Low	$ \begin{array}{c} -2.1 \\ -2.1 \\ -1.1 \\ -1.1 \end{array} $	-1.5 -0.5 -1.9 -1.3	106 99 122 269

Date	Z.C. No.		I or E	Elong. of Moon	W.	DMON' 113°4, N	TON N. 53°6	VANCOUVER W. 123°1, N. 49°2					
		Mag.			M.S.T.	a	b	Р	P.S.T.	a	b	Р	
Dec. 11 23 24 24 28	1332 3051 3185 3187 132	5.7 7.0 5.3 6.2 6.9	E I I I I	° 232 38 51 51 98	h m 3 01.1 18 19.7 18 34.8 18 45.9 19 01.4		$ \begin{matrix} m \\ -1.1 \\ -1.5 \\ +2.0 \\ -1.0 \\ -1.1 \end{matrix} $	° 318 97 5 83 113	h m 1 51.4 17 14.3 17 25.3 17 38.2 17 42.7	m -1.4 -1.4 -0.1 -1.4 -2.4		。 299 98 7 84 113	
29 29 29 30 30	252 264 272 284 380	7.4 7.0 5.9 7.4 7.4	I I I I I	108 109 110 111 120	17 21.1 20 46.7 23 09.4 1 24.4 21 38.4	-1.2 -1.3 -0.9 -0.3 -1.6	+1.5 +0.9 -0.9 0.0 -0.8	79 48 71 38 93	Sun 19 28.5 22 02.8 0 21.5 20 24.1	-1.5 -1.3 -0.5 -2.0	+1.3 -1.0 -0.5 -0.6	53 83 56 100	

NAMES OF OCCULTED STARS

The stars which are occulted by the moon are stars which lie along the zodiac; hence they are known by their number in the "Zodiacal Catalogue" (ZC), compiled by James Robertson and published in the Astronomical Papers Prepared for the Use of the American Ephemeris and Nautical Almanac, vol. 10, pt. 2 (U.S. Govt. Printing Office; Washington, 1940). The other names listed in the table are either (1) Bayer names, in which small Greek letters are used for the brighter stars in a constellation and Roman letters, if necessary, for the fainter stars (2) Flamsteed names, in which the stars are numbered consecutively from west to east across the constellation (3) numbers in the catalogues of Bode (B.), Heis (H¹), Gould (G.) and Hevelius (H.) or (4) numbers in the Bonner Durchmusterung or BD catalogue (e.g. + 18° 325).

						_	
Z.C. No.	Name	Z.C. No.	Name	Z.C. No.	Name	Z.C. No.	Name
64 98 132 144 143 146 241 252 252 252 252 252 252 252 252 252 25	116 B. Psc 60 Psc 169 B. Psc + $8^{\circ}158$ ϵ Psc 222 B. Psc 231 B. Psc 231 B. Psc 231 B. Psc 230 B. Psc 54 Cet + $10^{\circ}257$ 12 H! Ari 29 Ari + $13^{\circ}411$ σ Ari 124 B. Ari + $15^{\circ}447$ + $15^{\circ}427$ + $12^{\circ}427$ + $12^{\circ}47$ + $12^$	796 798 832 836 863 888 913 943 943 943 943 943 943 943 943 1029 1057 1071 106D 1106D 1106D 1106D 1106D 11029 1091 11057 1091 11057 1091 11057 1091 11050 1318 13359 13359	352 B. Tau 353 B. Tau 119 Tau 120 Tau +18° 920 127 Tau +9°1110 57 Ori 64 Ori 64 Ori 19 B. Gem +18°1214 26 Gem 98 B. Gem +17°1518 λ Gem +17°1518 λ Gem +15° 1734 +14° 1879 45 Cnc 50 Chc & Cnc +10° 1972 +10° 1972	1605 1623 1662 1726 1726 1727 1759D 1815 1815 1815 1883 1914 1925 2002 2021 2022 2021 2022 2022 2022 20	62 Leo 69 Leo 388 B. Leo -2° 3411 24 B. Vir. 18 G. Vir. 78 B. Vir 24 Vir 50 Vir 62 Vir 62 Vir 607 B. Vir 607 B. Vir 607 B. Vir 607 B. Vir 607 B. Vir 617 B. Vir 618 B. Lib 64 B. Lib 28 Lib 28 Lib 21 H. Lib 24 Lib 28 Lib	2509D 2591 2611 2629D 2715 2733 2745 2742 2828 2871D 2968D 2969 3051 3093 3154 3184 3185 3184 3185 3187 3281 3290 3320 3320	190 B. Oph 16 G. Sgr -19° 4832 39 G. Sgr 89 G. Sgr 18° 5182 -18° 5183 173 B. Sgr ρ Sgr 45 Sgr -17° 5699 337 B. Sgr 16 B. Cap β Cap 87 B. Cap γ Aqr -10° 5696 117 G. Cap 47 Cap 47 Cap 47 Cap 47 Cap 47 Cap 5699 162 B. Aqr -5° 5790 κ Aqr 255 B. Aqr 6 G. Sgr ρ Sgr 4 G. Sgr γ Sgr
730 760D 764D	97 Tau 333 B. Tau 104 Tau	1397D 1482 1528 1587D	ω Leo 14 Sex 84 B. Sex 55 Leo	2353 2394 2456 2465	ψ Opn 123 B. Sco 109 B. Oph -20° 4661	3420 3494 3512 3515	2° 5914 λ Psc 22 Psc 25 Psc
		1					

BY L. V. MORRISON

The maps show the tracks of stars brighter than $7^{n}5$ which will graze the limb of the Moon when it is at a favourable elongation from the Sun and at least 10° above the observer's horizon (5° in the case of stars brighter than $5^{n}5$ and 2° for those brighter than $3^{n}5$). Each track starts in the West at some arbitrary time given in the tables and ends beyond the area of interest, except where the letters A, B or S are given. A denotes that the Moon is at a low altitude, B that the bright limb interferes, and S that daylight interferes. The tick marks along the tracks denote 10 minute intervals of time which, when added to the time at the beginning of the track, give the approximate time of the graze at places along the tracks.

Observers positioned on, or very near, one of these tracks will probably see the star disappear and reappear several times at the edge of features on the limb of the Moon. The recorded times of these events (to a precision of a second, if possible) are very valuable in the study of the shape and motion of the Moon currently being investigated at the Royal Greenwich Observatory and the U.S. Naval Observatory. Observers situated near to any of these tracks who are interested should write to Dr. David W. Dunham, Cincinnati Observatory, Observatory Place, Cincinnati, Ohio, 45208, U.S.A., at least two months before the event, giving their approximate latitude and longitude, and details of the event will be supplied.*

The following table gives, for each track, the date, the name, Zodiacal Catalogue number and magnitude of the star, the time (U.T.) at the beginning of the track in the West, the percent of the Moon sunlit and whether the track is the northern (N) or southern (S) limit of the occultation. An asterisk after the track number refers the reader to the notes following the table; a dagger indicates that the star is a spectroscopic binary.

The numbering of the graze tracks differs slightly from that in previous years; there is no longer a continuous sequence. This arises from the method of preparing and editing the maps. It is easier and safer to preserve the original computer sequential numbering, even when certain tracks are later eliminated.

No.	Date	Name	Z.C.	Mag.	U.T.	%	L	No.	Date	Name	Z.C.	Mag.	U.T.	%	I
2 3 4 5 6	Jan. 6 9 10 10 10	$\begin{array}{r} -5^{\circ} 5790 \\ +9^{\circ} 158 \\ \pi \operatorname{Psc} \\ 281 \text{ B. Psc} \\ +13^{\circ} 351 \end{array}$	3290 201 240 241 325	7.3 7.5 5.6 6.9 7.4	h m 0 57 21 22 6 40 6 50 22 37	19 54 57 57 64	SSZZS	21 23 *24 25 26	Mar. 9 20 20 23 23	107 Tau 41 Lib 47 Lib 92 G. Sgr – 19° 5182	769 2233 2275 2718 2733	6.6 5.5 5.9 6.7 6.4	h m 2 40 3 43 10 46 8 25 10 46	51 76 73 42 41	N S N N N
7 8 9 10 11	13 22 24 25 Feb. 7	43 Tau -7° 3443 2 G. Lib. 11 H. Lib 12 H ! Ari	614 1809 2063 2217 299	5.7 6.9 6.7 5.5 6.3	7 06 9 07 7 10 12 36 2 51	83 67 45 32 38	ZSSSZ	27 *28 29 30 *31	Apr. 5 10 20 20 21	312 B. Tau ω Leo ρ Sgr 45 Sgr 16 B. Cap	736 1397 2826 2828 2968	6.2 5.5 4.0 6.0 6.2	3 36 5 21 8 14 7 51 11 06	26 77 58 58 47	NNNSN
11a 12 13 14 15	7 8 8 10 19	12 H! Ari 124 B. Ari 26 B. Tau 282 B. Tau 49 Vir	299 423 517 691 1884	6.3 6.4 6.4 6.6 5.3	3 23 4 14 22 51 3 27 5 46	38 48 56 68 81	ZZSZS	†32 33 34 †35 36	21 24 May 4 7 8	β Cap κ Aqr +18° 1214 κ Cnc 14 Sex	2969 3320 985 1359 1482	3.2 5.3 6.9 5.1 6.3	11 16 8 03 3 08 3 42 6 59	46 20 20 51 63	N N N N N N
16 17 18 19 *20	21 22 25 Mar. 6 9	-17° 4273 58 G. Sco 187 B. Sgr 29 Ari 333 B. Tau	2173 2343 2787 374 760	7.0 6.4 6.4 6.1 6.5	9 55 13 23 12 55 1 40 1 11	59 46 17 22 50	ZZZZZ	38 39 40 41 43	19 21 June 17 18 23	87 B. Cap -5° 5790 -7° 5727 6 G. Psc +14° 469	3051 3290 3259 3370 413	7.0 7.3 7.4 6.2 6.8	9 33 8 31 10 03 9 34 8 33	64 44 70 61 16	SZZZZ

*Editor's Note: A nominal fee is now charged for this service.

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No.	Date	Name	Z.C.	Mag.	U.T.	%	L	No.	Date	Name	Z.C.	Mag.	U.T.	%	L
44 45 46 †48 49	June 23 24 July 6 Aug. 1 17	σ Ari +16° 497 607 B. Vir. α Vir 145 B. Ari	422 532 2002 1925 450	5.5 7.2 6.8 1.2 6.6	h m 10 08 10 48 2 08 15 27 8 33	16 9 65 35 57	SSSSZ	67 68 69 70 72	Oct. 27 27 28 Nov. 1 18	89 G. Sgr 92 G. Sgr. - 19° 5182 162 B. Aqr. 28 Vir	2715 2718 2733 3281 1822	6.5 6.7 6.4 7.5 7.2	h m 23 03 23 28 1 43 3 26 10 26	28 28 29 70 15	S S S S S S
50 51 52 †53 54	18 20 20 28 Sept. 1	163 B. Tau 119 Tau +18° 950 α Vir ψ Oph	593 832 871 1925 2353	5.8 4.7 6.9 1.2 4.6	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	46 29 27 15 47	Zoooo	73 75 *76 *77 *78	19 25 28 29 Dec. 11	86 Vir ρ Sgr 231 B. Aqr. 255 B. Aqr. 45 Cnc	1971 2826 3344 3366 1309	5.8 4.0 6.8 6.6 5.7	13 10 1 04 22 37 4 23 2 36	7 15 53 55 82	SSSS Z
55 56 †57 58 60	4 13 15 15 17	45 Sgr σ Ari +17° 703 +18° 629 19 B. Gem	2828 422 629 643 943	6.0 5.5 7.5 6.7 6.2	2 23 10 15 3 40 5 16 9 27	79 81 66 65 44	いズズズス	*79 80 82 83 84	13 14 24 24 25	36 Sex 359 B. Leo 87 B. Cap 10° 5714 117 G. Cap	1566 1649 3051 3163 3184	6.6 6.3 7.0 7.3 7.1	14 25 6 17 1 52 21 31 1 54	59 52 11 17 19	SZSSZ
61 62 63 65	17 19 29 Oct. 14	$^{+18^{\circ}1141}_{+15^{\circ}1734}_{-20^{\circ}4661}_{+18^{\circ}1040}$	951 1212 2465 904	6.8 7.1 7.4 7.1	$\begin{array}{cccc} 10 & 53 \\ 10 & 35 \\ 2 & 21 \\ 10 & 39 \end{array}$	44 24 33 70	S Z S Z	85 86 87 88	25 29 29 30	47 Cap 169 B. Psc ε Psc +9° 206	3187 132 146 252	6.2 6.9 4.4 7.4	2 28 1 57 6 08 0 58	19 57 58 66	S S N S

NOTES ON DOUBLE STARS

- Track 20: ZC 760 is the mean of the double star ADS 3672. The components are 7^m0 and 7^m6; separation 1.12 in p.a. 306°.
- Track 24: ZC 2275 is the mean of the double star ADS 9834. The components are 6^m0 and 8^m1; separation 0'.'5 in p.a. 121°.
- Track 28: ZC 1397 is the mean of the double star ADS 7390. The components are 5^{m9} and 6^{m5}; separation 0'.'5 in p.a. 8°.
- Track 31: ZC 2968 is the brighter component of the double star ADS 13717. The companion is 10th magnitude; separation 0'.'8 in p.a. 84°.
- Track 76: ZC 3344 is the mean of the two brightest components of the system ADS 16270. These components are 7^m3 and 7^m8; separation 2''4 in p.a. 278°. A third component 8^m0, has a separation of 49'' in p.a. 98°.
- Track 77: ZC 3366 is the brighter component of the double star ADS 16392. The companion is 10th magnitude; separation 10'.'4 in p.a. 117°.
- Track 78: ZC 1309 is the mean of a close double star. The components are both estimated to be 6^m4 with separation 0'.'05.
- Track 79: ZC 1566 is the mean of a possible close double star. The components are both estimated to be 7^m3 with separation 0.403.








MAP OF THE MOON



South appears at the top.

MARS-LONGITUDE OF THE CENTRAL MERIDIAN

The following table lists the longitude of the central meridian of the geometric disk of Mars for each date at 0 hours U.T. (19 hours E.S.T. on the preceding date). To obtain the longitude of the central meridian for other times, add 14.6° for each hour elapsed since 0 hours U.T.

Date	Jan.	Feb.	Mar.	Apr.	May	June	July	Aug.
1	63.37	142.36	231.15	296.49	8.74	69.37	137.75	194.60
2	54.48	133.12	221.70	286.93	359.12	59.68	127.99	184.79
3	45.54	123.87	212.25	277.37	349.49	49.98	118.24	174.99
4	36.67	114.61	202.80	267.81	339.86	40.29	108.48	165.18
5	27.75	105.34	193.34	258.25	330.23	30.59	98.72	155.38
6	18.82	96.06	183.87	248.68	320.59	20.89	88.96	145.57
7	9.87	86.77	174.40	239.11	310.95	11.19	79.20	135.76
8	0.91	77.47	164.93	229.54	301.32	1.49	69.44	125.95
9	351.94	68.17	155.45	219.96	291.68	351.78	59.67	116.14
10	342.96	58.85	145.97	210.38	282.03	342.08	49.90	106.32
11	333.96	49.53	136.48	200.80	272.39	332.37	40.14	96.51
12	324.96	40.20	126.99	191.22	262.74	322.66	30.36	86.70
13	315.94	30.87	117.50	181.64	253.09	312.94	20.59	76.88
14	306.91	21.52	108.00	172.05	243.44	303.23	10.82	67.07
15	297.86	12.17	98.50	162.46	233.79	293.51	1.04	57.25
16	288.81	2.81	89.00	152.87	224.13	283.79	351.26	47.43
17	279.74	353.45	79.49	143.28	214.48	274.07	341.48	37.62
18	270.66	344.08	69.98	133.68	204.82	264.34	331.70	27.80
19	261.57	334.70	60.46	124.09	195.16	254.62	321.92	17.98
20	252.46	325.31	50.94	114.49	185.49	244.89	312.13	8.16
21	243.35	315.92	41.42	104.88	175.83	235.16	302.35	358.34
22	234.22	306.53	31.90	95.28	166.16	225.43	292.56	348.52
23	225.08	297.12	22.37	85.67	156.49	215.69	282.77	338.70
24	215.94	287.71	12.84	76.07	146.82	205.96	272.98	328.87
25	206.78	278.30	3.31	66.45	137.14	196.22	263.19	319.05
26 27 28 29 30 31	197.61 188.42 179.23 170.03 160.82 151.59	268.88 259.46 250.03 240.59	353.77 344.23 334.69 325.14 315.59 306.04	56.84 47.23 37.61 27.99 18.37	127.47 117.79 108.11 98.43 88.74 79.06	186.48 176.74 166.99 157.24 147.50	253.39 243.60 233.80 224.00 214.20 204.40	309.23 299.41 289.58 279.76 269.94 260.11





Latitude is plotted on the vertical axis (south at the top); longitude is plotted on the horizontal axis

ASTEROIDS—EPHEMERIDES AT OPPOSITION, 1976

Only one of the four major asteroids—Juno—comes to opposition in 1976. Nevertheless, all four will be prominent at some time during the year. Early in 1976, Ceres will be bright, and moves through the region between the Hyades and the Pleiades (see map). Juno comes to opposition on March 1, mag. 8.7 in Sextans; by April it is near Regulus (see map). Pallas approaches opposition late in 1976, but is rather far south (in Pyxis) for northern observers (see map). Vesta also approaches opposition late in 1976, but is much brighter and more conspicuous than Pallas, as it moves in retrograde fashion through Gemini (see map).

The following table lists the 1950 co-ordinates (for convenience in plotting on the *Atlas Coeli*) and the visual magnitudes of the four major asteroids on selected dates, at 0 h U.T.



74



JUPITER-LONGITUDE OF CENTRAL MERIDIAN

is favourably placed. Longitude increases hourly by 36.58° in System I (which applies to regions between the middle of the North Equatorial Belt and the middle of the South Equatorial Belt) and by 36.26° in System II (which applies to the rest of the planet). Detailed ancillary The table lists the longitude of the central meridian of the illuminated disk of Jupiter at 0th U.T. daily during the period when the planet tables may be found on pages 274 and 275 of The Planet Jupiter by B. M. Peek (Faber and Faber, 1958).

		Dec.	•	345.4 135.8 286.1 76.5 226.8	$17.2\\167.5\\317.9\\108.2\\258.6$	48.9 199.2 349.6 139.9 290.2	80.5 230.9 21.2 321.8	$\begin{array}{c} 112.1\\ 262.4\\ 52.6\\ 202.9\\ 353.2\end{array}$	143.5 293.7 84.0 234.3 24.5	174.8
		Nov.	0	$153.4 \\ 303.8 \\ 94.2 \\ 244.6 \\ 35.0 \\$	185.4 335.8 126.2 276.7 67.1	217.5 7.9 158.3 308.7 99.1	249.5 39.9 190.3 340.7 131.1	281.5 71.9 222.3 12.7 163.1	313.5 103.9 254.3 44.6 195.0	
		Oct.	۰	171.3 321.7 112.1 262.4 52.8	203.2 353.5 143.9 294.3 84.7	235.0 25.4 175.8 326.2 116.6	267.0 57.3 207.7 358.1 148.5	298.9 89.3 30.1 180.5	330.9 121.3 271.7 62.1 62.1	3.0
		Sept.	•	$\begin{array}{c} 341.7\\ 132.0\\ 282.3\\ 72.6\\ 222.9\end{array}$	13.2 163.5 313.8 104.1 254.4	44.7 195.0 345.3 135.6 285.9	76.2 226.6 16.9 167.2 317.5	$\begin{array}{c} 107.9 \\ 258.2 \\ 48.6 \\ 198.9 \\ 349.2 \\ 349.2 \end{array}$	139.6 289.9 80.3 21.0	
	EM II	Aug.	0	4.2 154.4 304.6 94.8 245.0	35.2 185.5 335.7 125.9 276.1	66.4 216.6 6.8 157.1 307.3	97.5 247.8 38.0 188.3 338.5	128.8 279.0 69.3 219.5 9.8	160.1 310.3 100.6 250.9 41.2	191.4
	ISYST	July	٥	28.9 179.1 329.2 119.3 269.5	59.6 209.8 359.9 150.1 300.2	90.4 240.6 30.7 180.9 331.0	121.2 271.4 61.5 211.7 1.9	152.1 302.3 92.4 32.8 32.8	183.0 333.2 123.4 123.4 63.8	214.0
		June	0	205.7 355.8 145.9 295.9 86.0	236.1 26.2 176.3 326.4 116.5	266.6 56.7 56.7 206.8 356.9 147.0	297.1 87.2 87.3 237.3 27.4 177.5	327.7 117.8 267.9 58.0 208.1	358.3 148.4 298.5 88.7 238.8	
		Mar.	0	83.0 233.0 23.0 173.0 323.1	$\begin{array}{c} 113.1\\ 263.1\\ 53.1\\ 53.1\\ 203.1\\ 353.1\\ 353.1\end{array}$	143.1 293.1 83.2 233.2 233.2	173.2 323.2 113.2 53.2 53.2	203.3 353.3 143.3 293.3 83.3	233.3 23.3 173.3 323.3 323.3 113.4	263.4
,		Feb.	-	52.2 202.2 352.3 142.3 292.3	82.4 232.4 172.5 322.5	112.6 262.6 52.6 202.7 352.7	142.7 292.7 82.8 232.8 22.8	172.8 322.9 112.9 262.9 52.9	202.9 352.9 143.0 293.0	
		Jan.	•	79.8 229.9 20.0 170.1 320.2	110.3 260.4 50.5 350.7 350.7	140.8 290.9 81.0 231.1 21.1	171.2 321.3 111.4 261.4 51.5	201.6 351.6 141.7 291.8 81.8	231.9 21.9 172.0 322.0 112.1	262.1
•		Dec.		116.6 274.6 72.6 230.6 28.6	186.6 344.6 142.5 300.5 98.5	256.4 54.4 212.4 10.3 168.3	326.2 124.2 80.0 238.0	35.9 193.8 351.7 149.6 307.5	105.4 263.3 61.2 17.0	174.9
•		Nov.	•	$\begin{array}{c} 55.7\\ 213.7\\ 11.8\\ 169.8\\ 327.9\end{array}$	125.9 283.9 82.0 240.0 38.1	196.1 354.1 152.2 310.2 108.3	266.3 64.3 222.4 20.4 178.4	336.5 134.5 292.5 90.5 248.6	46.6 204.6 2.6 160.6 318.6	
		Oct.	۰	197.1 355.1 153.1 311.1 109.1	267.1 65.1 65.1 223.1 21.1 179.1	337.1 135.1 293.2 91.2 249.2	47.2 205.2 3.2 161.3 319.3	117.3 275.4 73.4 231.4 29.4	187.5 345.5 143.5 301.6 99.6	257.7
		Sept.	0	138.6 296.5 94.4 252.3 50.3	208.2 6.1 164.0 322.0 119.9	277.8 75.8 233.7 31.7 189.6	347.6 145.5 303.5 101.4 259.4	57.4 215.3 13.3 171.3 329.2	127.2 285.2 83.2 83.2 241.2 39.1	
	EM I	Aug.	•	284.5 82.3 82.3 38.0 195.9	353.7 151.6 309.4 107.3 265.1	63.0 220.8 18.7 176.6 334.4	$\begin{array}{c} 132.3\\ 290.2\\ 88.1\\ 88.1\\ 245.9\\ 43.8\end{array}$	201.7 359.6 157.5 315.4 113.3	271.2 69.1 227.0 24.9 182.8	340.7
	SYST	July	0	72.7 230.5 28.2 186.0 343.8	141.6 299.3 97.1 254.9 52.7	210.5 8.3 166.0 323.8 121.6	279.4 77.2 235.0 32.8 32.8 190.6	348.5 146.3 304.1 101.9 259.7	57.5 215.4 13.2 171.0 328.8	126.7
		June	•	20.6 178.3 336.0 133.7 291.4	89.1 246.9 44.6 202.3 0.0	157.7 315.5 113.2 113.2 270.9 68.7	226.4 24.1 181.9 339.6 137.4	295.1 92.9 250.6 48.4 206.1	3.9 161.6 319.4 117.2 274.9	
		Mar.	0	275.9 73.6 231.2 28.8 186.5	344.1 141.8 299.4 97.1 254.7	22.3 210.0 7.6 322.9 322.9	$\begin{array}{c} 120.5\\ 278.2\\ 75.8\\ 233.5\\ 31.1\\ 31.1 \end{array}$	188.8 346.4 144.0 301.7 99.3	257.0 54.6 212.2 9.9 167.5	325.2
		Feb.	0	23.8 181.5 339.2 136.9 294.5	92.2 249.9 47.5 205.2 2.9	160.5 318.2 115.8 273.5 71.1	228.8 26.5 184.1 341.8 139.4	297.1 94.7 252.4 50.0 207.7	5.3 163.0 320.6 118.3	
		Jan.	0	174.9 332.7 130.4 288.2 85.9	243.6 41.4 199.1 356.8 154.5	312.2 110.0 267.7 65.4 223.1	20.8 178.5 336.2 133.9 291.6	89.3 247.0 44.7 202.4 0.1	157.8 315.4 113.1 270.8 68.5	226.2
		Day (0 ^h U.T.)		-9640	96876	122242	20 18 19 20 20	222222	30,58,7,8	31

JUPITER-PHENOMENA OF THE BRIGHTEST SATELLITES 1976

Times and dates given are E.S.T. The phenomena are given for latitude 44° N., for Jupiter at least one hour above the horizon, and the sun at least one hour below the horizon, as seen from Central North America. The symbols are as follows: E—eclipse, O—occultation, T—transit, S—shadow, D—disappearance, R—reappearance, I—ingress, e—egress. Satellites move from east to west across the face of the planet, and from west to east behind it. Before opposition, shadows fall to the west, and after opposition to the east. Thus eclipse phenomena occurs on the west side until November 18, and on the east thereafter phenomena occur on the west side until November 18, and on the east thereafter.

	JANU	JARY		d	h m	Sat.	Phen.	d	h m	Sat.	Phen.	d	h m	Sat.	Phen.
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JUPITER'S BELTS AND ZONES

Viewed through a telescope of 6-inch aperture or greater, Jupiter exhibits a variety of changing detail and colour in its cloudy atmosphere. Some features are of long duration, others are short-lived. The standard nomenclature of the belts and zones is given in the figure.



COMETS IN 1976

BY BRIAN G. MARSDEN

Comet	Perihelion Date	Perihelion Distance	Period
Westphal Wolf Gunn Churyumov-Gerasimenko Harrington-Abell Tempel-Swift Neujmin 2 d'Arrest Klemola Schaumasse Pons-Winnecke	Jan. 3 Jan. 25 Feb. 10 Apr. 6 Apr. 21 May 25 June 18 Aug. 12 Aug. 20 Sept. 5 Nov. 28	1.26 A.U. 2.50 2.44 1.30 1.77 1.60 1.28 1.16 1.28 1.16 1.77 1.21 1.25	63.0 yr 8.4 6.8 6.6 7.6 6.4 5.4 6.2 11.0 8.2 6.4

The following periodic comets are expected to be at perihelion during 1976:

Comet Westphal faded out on its way in to perihelion in 1913, and it is doubtful whether it can be observed any more. Comet Wolf was recovered in May 1975 but will remain very faint. Comet Gunn, discovered in 1970, has been followed all the way around its orbit. Comets Churyumov-Gerasimenko and Klemola have been observed at only one previous perihelion passage, and thus the predictions, particularly in the case of Comet Klemola, are more uncertain than for the other comets listed here. Comet Harrington-Abell, which passed only 0.04 A.U. from Jupiter in 1974, will be faint. Comets Tempel-Swift and Neujmin 2 have been missing since 1908 and 1927, respectively, and it is extremely improbable that they will be observed in 1976.

Comet d'Arrest is making its most favourable return since its discovery in 1851 and will pass only 0.15 A.U. from the earth in August 1976. A brief ephemeris is:

Date	R.A. (1950)	Dec. (1950)	Mag.
July 1	19 ^h 00 ^m 0 19 ^h 16 ^m 9	$+22^{\circ}04'$ +21^{1}2'	8.9
21	19 ^h 42 ^m 5	+21 12 +17°25′ + 8°43′	7.5
Aug. 10	20°22°0 21 ^h 19 ^m 4 22 ^h 28 ^m 6	$-6^{\circ}25'$ $-23^{\circ}26'$	6.4
30 Sent 9	23 ^h 31 ^m 6	- 34°34′ - 39°24′	7.1
19	0 ^h 42 ^m 6	-40°40′	8.5

Comets Schaumasse and Pons-Winnecke are badly placed for observation and it is doubtful that the former will be detectable, even with very large telescopes.

Any bright comets, other than d'Arrest, that may appear during 1976 will be completely unexpected.

METEORS, FIREBALLS AND METEORITES

By Peter M. Millman

Meteoroids are small solid particles moving in orbits about the sun. On entering the earth's atmosphere at velocities ranging from 15 to 75 kilometres per second they become luminous and appear as meteors or fireballs and in rare cases, if large enough to avoid complete vaporization, they may fall to the earth as meteorites.

Meteors are visible on any night of the year. At certain times of the year the earth encounters large numbers of meteors all moving together along the same orbit. Such a group is known as a meteor shower and the accompanying list gives the more important showers visible in 1976.

An observer located away from city lights and with perfect sky conditions will see an overall average of 7 sporadic meteors per hour apart from the shower meteors. These have been included in the hourly rates listed in the table. Slight haze or nearby lighting will greatly reduce the number of meteors seen. More meteors appear in the early morning hours than in the evening, and more during the last half of the year than during the first half.

The radiant is the position among the stars from which the meteors of a given shower seem to radiate. The appearance of any very bright fireball should be reported immediately to the nearest astronomical group or other organization concerned with the collection of such information. Where no local organization exists, reports should be sent to Meteor Centre, National Research Council, Ottawa, Ontario, K1A 0R8. Free fireball report forms and instructions for their use, printed in either French or English, may be secured at the above address. If sounds are heard accompanying a bright fireball there is a possibility that a meteorite may have fallen. Astronomers must rely on observations made by the general public to track down such an object.

	Show	er Maxin	num		Ra	diant		Circula		Normal Duration
Shower	Date	E.S.T.	Moon	Posit at M R.A.	tion Iax. Dec.	Da Mo R.A.	aily otion Dec.	Observer Hourly Rate	Velocity	to 1/4 strength of Max.
Quadrantids Lyrids η Aquarids δ Aquarids Perseids Orionids Taurids Geminids Ursids	Jan. 4 Apr. 22 May 4 July 28 Aug. 12 Oct. 21 Nov. 4 Dec. 13 Dec. 22	h 02 21 23 19 00 02 	N.M. L.Q. F.Q. N.M. F.M. F.M. L.Q. N.M.	h m 15 28 18 16 22 24 22 36 03 04 06 20 03 32 07 32 14 28	$^{\circ}$ +50 +34 00 -17 +58 +15 +14 +32 +76	$ m \\ +4.4 \\ +3.6 \\ +3.4 \\ +5.4 \\ +4.9 \\ +2.7 \\ +4.2 \\ $	$ \overset{\circ}{\begin{array}{c} 0.0 \\ +0.4 \\ +0.17 \\ +0.12 \\ +0.13 \\ +0.13 \\ -0.07 \end{array} } $	40 15 20 20 50 25 15 50 15	km/sec 41 48 64 40 60 66 28 35 34	days 1.1 2 3 4.6 2 2.6 2

METEOR SHOWERS FOR 1976

CANADIAN METEORITE IMPACT SITES

BY P. BLYTH ROBERTSON

The search for ancient terrestrial meteorite craters, and investigations in the related fields of shock metamorphism and cratering mechanics, have been carried out since 1951 at the Earth Physics Branch (formerly Dominion Observatory) Department of Energy, Mines and Resources. Approximately 40 percent of the craters recognized in the world have been discovered in Canada. At large impact sites (greater than approximately 1500 m diameter) original meteoritic material is not recognizable. Extreme shock pressures and temperatures at impact vapourize or melt the meteorite and it becomes intimately mixed and disseminated in the melted target rocks. Hypervelocity impact craters are therefore identified by the presence of shock metamorphic effects, the characteristic suite of deformation in the target rocks produced by shock pressures exceeding approximately 75 kilobars. The twenty-three "confirmed" structures in the Table contain definitive evidence of shock metamorphism, and are listed in order of their discovery. The latter three of these features were recognized during 1974. The "possible" sites represent only a few of those under consideration but where definitive shock metamorphic effects have not been found. Craters where data have been obtained through diamond-drilling or geophysical surveys are marked "D" and "G", respectively, and "A" signifies those sites accessible by road. "Float" includes boulders and pebbles in glacial deposits.

				* · · · · · · · · · · · · · · · · · · ·		
Name	Lat.	Long.	Diam. (km)	($\times 10^{6}$ years)	Surface Expression	Visible Geologic Featues
A Confirmed sites Brent, Ont. Brent, Ont. Manicougan, Que. Clarwater Lake West, Que. Clarwater Lake Bast, Que. Holleford, Ont. Deep Bay, Sask. Carswell, Sask.	61°17' 46°05' 51°23' 56°13' 56°24' 58°27'	73°40° 78°29° 68°42° 74°30° 74°30° 76°38° 102°59° 109°30°	3092.5 3092.5 3092.5	<pre>^ 1 451 210±4 210±4 285±30 585±30 585±30 100±50 100±50 100±50 100±50</pre>	rimmed circular lake sediment-filled shallow depression circumferal lake, central elevation island ring in circular lake circular lake sediment-filled shallow depression discontinuous circular ridge	raised rim fracturing impact melt sedimentary fill sedimentary fill b D sedimentary fill b D Sedimentary foat b D
Lac Couture, Que. West Hawk Lake, Man. Pilot Lake, N.W.T. Nicholson Lake, N.W.T. Steen River, Alta. Sudbury, Ont.	60°08′ 49°46′ 60°17′ 59°31′ 46°36′	75°18' 95°11' 111°01' 117°38' 117°38' 81°11'	10 3 12.5 13.5 100	300 ± 50 150 ± 50 300 ± 150 300 ± 150 95 ± 7 1700 ± 200	circular lake circular lake circular lake irregular lake with islands none, buried to 200 metres elliptical basin	breccia breccia float none fracturing, breccia float A D G breccia none breccia, impact melt, A D G
Charlevoix, Que.	47°32′	70°18′	35	350±25	semi-circular trough, central elevation	breccia float, shatter cones, impact melt A G
take Mistatun, Labr. Lake Six Martin, Man. Lake Wanapitei, Ont. Lake Wanapitei, Ont. Lac La Moinerie, Que. Haughton Dome, N.W.T. Slate Islands, Ont.	55°53 51°47 56°24 57°26 75°22 75°22 75°22	63°18 98°33' 80°24' 66°36' 87°00' 87°00'	20 8.5 13 13 13 13 13 13 13 13 13 13 13 13 13	225 ± 25 225 ± 25 37 ± 2 > 150 < 150 < 1100	eliptical lake and central island none, buried and eroded lake-filled, partly circular lake and central island lake filled, partly circular islands are central uplift of	breccia, impact melt impact melt breccia float breccia float breccia float breccia float shatter cones, breccia shatter cones, breccia dikes
Ile Rouleau (L. Mistassini) Que.	50°41′	73°53′	4	< 1000	submerged structure island is central uplift of submerged structure.	shatter cones, breccia dikes
B Possible sites Skeleton Lake, Ont. Kakiatukallak Lake, Que. Darron Lake, N.W.T. Charron Lake, Man. Eagle Butte, Alta. Eagle Butte, Alta. Poplar Bay, Ont. Poplar Bay, Cut. Viewfield, Sask.	45°15′ 57°42′ 57°42′ 52°44′ 52°35′ 50°23′ 49°33′	79°26' 71°40' 87°41' 95°15' 110°305' 94°05' 95°48' 103°04'	4040000 2300 230		lake-filled partly circular circular lake circular lake circular lake slight, buried and eroded circular lake circular lake circular lake circular depression	breccia, sedimentary float A G breccia float preccia float disturbed beds A D disturbed beds A D ? A D G

SATURN AND ITS SATELLITES

BY TERENCE DICKINSON

Saturn, with its system of rings, is a unique sight through a telescope. There are three rings. The outer ring A has an outer diameter 169,000 miles. It is separated from the middle ring B by Cassini's gap, which has an outer diameter 149,000 miles, and an inner diameter 145,000 miles. The inner ring C, also known as the dusky or crape ring, has an outer diameter 112,000 miles and an inner diameter 93,000 miles. Evidence for a fourth, innermost ring has been found; this ring is very faint.

Saturn exhibits a system of belts and zones with names and appearances similar to those of Jupiter (see diagram pg. 79).

Titan, the largest and brightest of Saturn's moons is seen easily in a 2-inch or larger telescope. At elongation Titan appears about 5 ring-diameters from Saturn. The satellite orbits Saturn in about 16 days and at magnitude 8.4^* dominates the field around the ringed planet.

Rhea is considerably fainter than Titan at magnitude 9.8 and a good quality 3-inch telescope may be required to detect it. At elongation Rhea is about 2 ring-diameters from the centre of Saturn.

lapetus is unique among the satellites of the solar system in that it is five times brighter at western elongation (mag. 10.1) than at eastern elongation (mag. 11.9). When brightest, Iapetus is located about 12 ring-diameters west of its parent planet.

Of the remaining moons only Dione and Tethys are seen in "amateur"-sized telescopes.

*Magnitudes given are at mean opposition.

ELONGATIONS OF SATURN'S SATELLITES, 1976 (E.S.T.)

d 3 8	JANUARY h Sat. H 19.2 Rh 07.5 Rh 14 4 Ti	Elong. E E W	d 25 28 29	h 01.3 03.3 13.8	Sat. Rh Ti Rh	Elong. E W E	d 8 9 14	h 08.8 22.0 10.6	Sat. Ti Rh Rh	Elong. E E E	d 14 19 22 23	h 14.1 02.6 06.8	Sat. Rh Rh Ti Rh	Elong. E W F
12 16 17	19.8 Rh 19.7 Ti 08.1 Rh	E E F	d	\mathbf{A}	PRIL Sat. Rh	Elong. F	Sat sun	urn be , eloi	eing n ngatio	ear the ns are	28 30	03.6 13.6	Rh Ti	Ë E
21	20.4 Rh	Ĕ	5	09 .1	Ti	Ĕ	Jun	e 14 ar	id Au	gust 25.		NOV	ЕМВ	ER
24	11.7 Ti	W	.7	14.7	Rh	E			aug	-	d	h	Sat.	Elong.
26	08.7 Rh	E	12	03.1	Rh T:	E	4	AU	GUS.	Flore		16.0	Rh	E
30	21.0 Ki	Е	16	15 6	Rh	Ē	25	20 0	Rh	Elong. E		04.3	Ti	w
	FEBRUARY	Y	21	04.1	Rh	Ē	26	15.9	Ia	ŵ	10	17.0	Ŕ'n	Ë
d	h Sat. I	Elong.	21	08.4	Ti	Е	27	12.8	Ti	E	15	05.4	Rh	E
1	17.0 Ti	E	25	16.6	Rh	E	30	08.6	\mathbf{Rh}	E	15	07.9	la	w
4	09.3 Kh	E	27	12 1	la Ti	E		CEDT	TEMD	FD	15	12.8		E
8	21 7 Rh	Ē	30	05 1	Rh	F	Ы	h	Sat	Elong	23	05 4	Ti	Ŵ
ğ	09.1 Ti	w	50	05.1	I CII	Ľ	3	21.2	Rh	E E	24	06.3	Ŕ'n	Ë
13	10.0 Rh	E		N	1AY		4	06.2	Ti	w	28	18.7	Rh	Ē
17	14.4 Ti	E	d	h	Sat.	Elong.	8	09.8	Rh	E				
17	22.3 Rh	E	4	17.6	Rh	E	12	13.4	Ti Dh	E		DEC	EMB	ER
25	10.7 Kn	W		08.2	Ph	E	17	10 9	Rh	F	1	11 6	Sat.	Elong.
26	23.0 Rh	Ë	13	18.7	Rh	Ĕ	20	06.7	Ťi	ŵ	3	07.1	Rh	Ē
		-	15	01.6	Ti	ŵ	21	23.4	Rh	E	7	19.5	Rh	Ē
	MARCH		18	07.2	Rh	Е	26	12.0	Rh	E	9	03.9	Ti	W
d	h Sat. I	Elong.	22	19.8	Rh	Ē	28	13.8	Ti	E	12	07.9	Rh	E
2	11.4 Kh	E	23	08.3	11	E		OCTO	סססר		16	20.2	Rh	E
6	23.7 Rh	F	31	00.3	Ti	W	a	h	Sat	Flong	21	09.9	Ph	Ē
11	12.1 Rh	Ĕ	31	20.9	Rh	Ë	1	00.5	Rh	E E	25	02.0	Ti	ŵ
12	04.8 Ti	w				-	5	13.0	Rh	E	25	19.1	Ĩa	Ê
16	00.5 Rh	E		្សា	UNE		6	06.9	Ti	w	25	21.0	Rh	Е
18	03.8 la	w	d	h	Sat.	Elong.	7	02.4	la	E	30	09.3	Rh	Ē
20 20	10.4 Ti 12.9 Rh	E E	5 6	09.4 07.5	Rh Ia	W	10	13.9	Rh Ti	Е Е	33	07.8	Ti	Е

TABLE OF PRECESSION FOR 50 YEARS If Declination is positive, use inner R.A. scale; if declination is negative, use outer R.A. scale, and reverse the sign of the precession in declination

R.A.	Dec. –	h H 24 00 23 30 23 00	22 30 22 00 21 30	21 00 20 30 20 00	19 30 19 00 18 30 18 00	12 00 11 30 11 00	10 30 10 00 9 30	9 8 90 90 90 90	00 20 00 20 00 20
R.A.	Dec.+	h m 12 00 11 30 11 00	10 30 10 00 9 30	9 00 8 30 00	7 30 6 30 6 00	24 00 23 30 23 00	22 30 21 30 30	21 00 20 30 20 00	19 30 19 00 18 30 18 00
Prec.	Dec.	-16.7 -16.7 -16.6 -16.1	-15.4 -14.5 -13.2	$^{-11.8}_{-10.2}$	6.4 - 2.2 0.0	$^{+16.7}_{+16.6}$	+15 4 +14.5 +13 2	$^{+11.8}_{+10.2}$	+++ + 4.3 0.0
	00	+2.56 2.56 2.56	2.56 2.56 2.56	2.56 2.56 2.56	2.56 2.56 2.56	2.56 2.56 2.56	2.56 2.56 2.56	2.56 2.56 2.56	2.56 2.56 2.56 2.56
1	10°	н +2.56 2.59 2.61	2.64 2.68 2.68	2.70 2.72 2.73	2.75 2.75 2.75 2.76	2.56 2.53 2.51	2.49 2.46 2.44	2.42 2.40 2.39	2.38 2.37 2.37 2.36
	20°	+2.56 2.61 2.67	2.72 2.76 2.81	2.85 2.88 2.91	2.93 2.95 2.96	2.56 2.51 2.45	2.40 2.36 2.31	2.27 2.24 2.21	2.19 2.16 2.16 2.16
	30°	m +2.56 2.64 2.73	2.81 2.88 2.95	3.02 3.07 3.12	3.16 3.18 3.20 3.20	2.56 2.48 2.39	2.31 2.24 2.17	2.11 2.05 2.00	1.97 1.94 1.92 1.92
cension	40°	+2.56 2.68 2.80	2.92 3.03 3.13	3.22 3.30 3.37	3.42 3.46 3.50	2.56 2.44 2.32	2.20 2.09 1.99	$1.90 \\ 1.81 \\ 1.75 \\ 1.75$	1.70 1.66 1.63 1.63
in right as	50°	+2.56 2.73 2.90	3.07 3.22 3.37	3.50 3.61 3.71	3.79 3.84 3.88 3.88	2.56 2.39 2.22	2.05 1.90 1.75	1.62 1.51 1.41	$1.33 \\ 1.28 \\ 1.25 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ 1.23 \\ $
Precession	60°	+2.56 2.81 3.06	3.30 3.52 3.73	3.92 4.09 4.23	4.44 4.42 4.47 49 49	2.56 2.31 2.06	1.82 1.60 1.39	$ \begin{array}{c} 1.20\\ 1.03\\ 0.89\end{array} $	0.78 0.70 0.65 0.63
	70°	+2.56 2.96 3.36	3.73 4.09 4.42	4.73 4.99 5.21	5.33 5.60 5.60 5.60	2.56 2.16 1.77	1.39 1.03 0.70	$^{0.40}_{-0.09}$	$\begin{array}{c} -0.27 \\ -0.40 \\ -0.47 \\ -0.50 \end{array}$
	75°	+2.56 3.10 3.64	4.15 4.64 5.09	5.50 5.86 6.16	6.40 6.58 6.72 6.72	2.56 2.02 1.48	${0.97 \\ 0.46 \\ +0.03 }$	$\begin{array}{c} -0.38 \\ -0.74 \\ -1.04 \end{array}$	-1.28 -1.45 -1.56 -1.60
	80°	+2.56 3.38 4.19	4.98 5.72 6.40	7.02 7.57 8.03	8.40 8.66 8.82 8.88	2.56 1.82 0.93	$^{+0.14}_{-0.60}$	-1.90 -2.45 -2.91	-3.27 -3.54 -3.70 -3.70 -3.75
	δ=85°	+ 2.56 4.22 5.85	7.43 8.92 10.31	11.56 12.66 13.58	14.32 14.85 15.18 15.29	$^{+}_{-0.73}$	- 2.31 - 3.80 - 5.19	- 6.44 - 7.54 - 8.46	$\begin{array}{c} - & 9.20 \\ - & 9.73 \\ -10.06 \\ -10.17 \end{array}$
Prec.	Dec.	, +16.7 +16.6 +16.1	+15.4 +14.5 +13.2	$^{+11.8}_{+8.3}$	+++ 0.02334	-16.7 -16.6 -16.1	-15.4 -14.5 -13.2	-11 8 -10.2 - 8.3	- 6.4 - 4.3 - 2.2 0.0
Ŗ.A.	tor Dec.+	$^{ m h}_{ m 0} {}^{ m m}_{ m 1} {}^{ m m}_{ m 0} {}^{ m m}_{ m 1} {}^{ m m}_{ m 0} {}^{ m m}_{ m 1} {}^{ m m}_{ m 0} {}^{ m m}_{ m 1} {}^{ m m}_{ m 0} {}^{ m m}_{ m 1} {}^{ m m}_{ m$	2 30 2 30 30	4 3 3 00 4 00 0 00	4 30 5 30 6 00	12 00 13 00 13 00	13 30 14 00 14 30	15 00 15 30 16 00	16 30 17 00 17 30 18 00
R.A.	Dec. –	h m 12 00 13 00	13 30 14 00 14 30	15 00 15 30 16 00	16 30 17 00 18 00	0 00 0 300 1 000	1 30 2 30	3 00 4 00	6 20 6 20 00 00 00 00

FINDING LIST OF NAMED STARS

Name	Con.	R.A.	Name	Con.	R.A.
Acamar, ā'ka-mär Achernar, ā'kēr-när Acrux, ā'krŭks Adhara, a-dā'ra Al Na'ir, ăl-nâr'	 θ Eri α Eri α Cru ε CMa α Gru 	02 01 12 06 22	Gienah, jē'n <i>a</i> Hadar, hăd'är Hamal, hăm'ăl Kaus Australis, kôs ôs-trā'lis	γ Crv β Cen α Ari ε Sgr	12 14 02 18
Albireo, ăl-bir'ē-ō Alcyone, ăl-sī'ō-nē Aldebaran, ăl-dëb'a-ran Alderamin, ăl-dër'a-mĭn Algenib, ăl-jē'nĭb	β Cyg η Tau α Tau α Cep γ Peg	19 03 04 21 00	Kochab, kō'kăb Markab, mär'kăb Megrez, më'grĕz Menkar, měn'kär Menkent, měn'kěnt	$ \begin{array}{c} \beta \ \text{UMi} \\ \alpha \ \text{Peg} \\ \delta \ \text{UMa} \\ \alpha \ \text{Cet} \\ \theta \ \text{Cen} \end{array} $	14 23 12 03 14
Algol, ăl'gŏl Alioth, ăl'ĭ-ŏth Alkaid, ăl-kād' Almach, ăl'măk Alnilam, ăl-nī'lăm	β Per ε UMa η UMa γ And ε Ori	03 12 13 02 05	Merak, mē'răk Miaplacidus, mī'a-plăs'i-dus Mira, mī'ra Mirach, mī'rāk	$\begin{array}{c} \beta \ \text{UMa} \\ \beta \ \text{Car} \\ o \ \text{Cet} \\ \beta \ \text{And} \end{array}$	10 09 02 01
Alphard, ăl'färd Alphecca, ăl-fĕk'a Alpheratz, ăl-fē'răts Altair, ăl-târ' Ankaa	α Hya α CrB α And α Aql α Phe	09 15 00 19 00	Mirfak, mĭr'făk Mizar, mī'zär Nunki, nŭn'kē Peacock Phecda, fĕk'da	α Per ζ UMa σ Sgr α Pav γ UMa	03 13 18 20 11
Antares, ăn-tā'rēs Arcturus, ärk-tū'r <i>ă</i> s Atria, ā'trī-a Avior, ă-vĭ-ôr' Bellatrix, bě-lā'trĭks	α Sco α Boo α TrA ε Car γ Ori	16 14 16 08 05	Polaris Pollux, pŏl'ŭks Procyon, prō'sī-ŏn Ras-Algethi, ràs'ŭl-jē'the Rasalhague, ràs'ŭl-hā'gwē	α UMi β Gem α CMi α Her α Oph	01 07 07 17 17
Betelgeuse, bět'el-juz Canopus, ka-nō'păs Capella, ka-pēl'a Caph, kăf Castor, kàs'tẽr	α Ori α Car α Aur β Cas α Gem	05 06 05 00 07	Regulus, rěg'u-l <i>ŭ</i> s Rigel, ri'jel Rigil Kentaurus ri'jīl kěn-tô'r <i>ŭ</i> s Sabik, sā'bĭk	α Leo β Ori α Cen η Oph	10 05 14 17
Deneb, děn'ěb Denebola, dě-něb'ō-la Diphda, dĭf'da Dubhe, dŭb'ê Elnath, ěl'năth	$\begin{array}{c} \alpha \ Cyg \\ \beta \ Leo \\ \beta \ Cet \\ \alpha \ UMa \\ \beta \ Tau \end{array}$	20 11 00 11 05	Scheat, shē'ăt Schedar, shēd'ar Shaula, shô'la Sirius, sir'i-ŭs Spica, spī'ka	β Peg α Cas λ Sco α CMa α Vir	23 00 17 06 13
Eltanin, ĕl-tā'nĭn Enif, ĕn'īf Fomalhaut, fō'm <i>ā</i> l-ôt Gacrux, gä'krŭks	γ Dra ε Peg α PsA γ Cru	17 21 22 12	Suhail, sŭ-hāl' Vega, vē'ga Zubenelgenubi, zōō-běn'ěl-jĕ-nū'bē	λ Vel α Lyr α Lib	09 18 14

Pronunciations are generally as given by G. A. Davis, *Popular Astronomy*, **52**, 8 (1944). Key to pronunciation on p. 5.

THE BRIGHTEST STARS

BY DONALD A. MACRAE

The 286 stars brighter than apparent magnitude 3.55.

Star. If the star is a visual double the letter A indicates that the data are for the brighter component. The brightness and separation of the second component B are given in the last column. Sometimes the double is too close to be conveniently resolved and the data refer to the combined light, AB; in interpreting such data the magnitudes of the two components must be considered.

Visual Magnitude (V). These magnitudes are based on photoelectric observations, with a few exceptions, which have been adjusted to match the yellow coloursensitivity of the eye. The photometric system is that of Johnson and Morgan in Ap, J, vol. 117, p. 313, 1953. It is as likely as not that the true magnitude is within 0.03 mag. of the quoted figure, on the average. Variable stars are indicated with a 'v''. The type of variability, range, R, in magnitudes, and period in days are given.

Colour index (B-V). The blue magnitude, B, is the brightness of a star as observed photoelectrically through a blue filter. The difference B-V is therefore a measure of the colour of a star. The table reveals a close relation between B-V and spectral type. Some of the stars are slightly reddened by interstellar dust. The probable error of a value of B-V is only 0.01 or 0.02 mag.

Type. The customary spectral (temperature) classification is given first. The Roman numerals are indicators of *luminosity class*. They are to be interpreted as follows: Ia—most luminous supergiants; Ib—less luminous supergiants; II—bright giants; III—normal giants; IV—subgiants; V—main sequence stars. Intermediate classes are sometimes used, e.g. Iab. Approximate absolute magnitudes can be assigned to the various spectral and luminosity class combinations. Other symbols used in this column are: p—a peculiarity; e—emission lines; v—the spectrum is variable; m—lines due to metallic elements are abnormally strong; f—the O-type spectrum has several broad emission lines; n or nn—unusually wide or diffuse lines. A composite spectrum, e.g. M1 Ib+B, shows up when a star is composed of two nearly equal but unresolved components. The table now includes accurate spectral and luminosity classes for most stars in the southern sky. These were provided by Dr. Robert Garrison of the Dunlap Observatory. A few types in italics and parentheses remain poorly defined. Types in parentheses are less accurately defined (g—giant, d—dwarf, c—exceptionally high luminosity). All other types were very kindly provided especially for this table by Dr. W. W. Morgan, Yerkes Observatory.

Parallax (π). From "General Catalogue of Trigonometric Stellar Parallaxes" by Louise F. Jenkins, Yale Univ. Obs., 1952.

Absolute visual magnitude (M_v) , and distance in light-years (D). If π is greater than 0.030'' the distance corresponds to this trigonometric parallax and the absolute magnitude was computed from the formula $M_v = V + 5 + 5 \log \pi$. Otherwise a generally more accurate absolute magnitude was obtained from the luminosity class. In this case the formula was used to *compute* π and the distance corresponds to this "spectroscopic" parallax. The formula is an expression of the inverse square law for decrease in light intensity with increasing distance. The effect of absorption of light by interstellar dust was neglected, except for three stars, ζ Per, σ Sco and ζ Oph, which are significantly reddened and would therefore be about a magnitude brighter if they were in the clear.

Annual proper motion (μ) , and radial velocity (R). From "General Catalogue of Stellar Radial Velocities" by R. E. Wilson, Carnegie Inst. Pub. 601, 1953. The information on radial velocities was brought up-to-date in 1975 by Dr. C. T. Bolton of the Dunlap Observatory. Italics indicate an average value of a variable radial velocity.

The star names are given for all the officially designated navigation stars and a few others. Throughout the table, a *colon* (:) indicates an uncertainty.

		Sun	Alpheratz Caph Y Peg = Algenib Ankaa Schedar Diphda Ruchbah Achernar
			Manganese star Var. <i>R</i> 0,98, 0,10 ^d β CMa type, <i>R</i> in <i>V</i> 2.8 <i>B</i> 12 ^m 28'' Var. <i>B</i> 8.18 ^m 2'' <i>A</i> 4.1 ^m <i>B</i> 4.1 ^m 1'' Ecl. ? <i>R</i> 0.08. ^m 759 ^d
Radial Velocity	R	km/sec	-11.7 +111.8 +201.1 +22.8 +22.8 +11.8 +12.2 +22.8 +12.8 +13.1 +03.3 +13.1 +03.3 +13.1 +03.3 +13.1 +13.1 +13.1 +11.5 +11.5 +12.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +11.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +12.8 +
Proper Motion	д		0.209 0.555 0.010 0.555 0.255 0.161 0.058 0.234 1.221 0.058 0.234 0.025 0.025 0.250 0.035 0.250 0.035 0.250 0.098 1.921
Distance light-years	D	l.y.	96 570 570 150 150 150 150 150 150 150 150 150 15
Absolute Magnitude	M	+4.84	$\begin{array}{c} -0.\\ -1.\\ -1.\\ -1.\\ -1.\\ -1.\\ -1.\\ -1.\\ -1$
Parallax	ĸ	:	$\begin{array}{c} 0.024\\ 0.072\\ 0.072\\ 0.035\\ 0.035\\ 0.037\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.023\\ 0.$
Spectral Classification	Type	^	VSV H H H H H H H H H H H H H
		G2	BREESE BOEREESE
Colour Index	B-V	+0.63	-0.08 -0.08 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.23 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.62 -0.
Visual Magnitude	7	-26.73	80.222.23 80.222.23 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222 80.222
Declination	80 Dec.	0	$\begin{array}{c} ++59\\ ++59\\ +155\\ 04\\ +155\\ 04\\ +56\\ 255\\ +57\\ +56\\ 255\\ 255\\ 255\\ 255\\ 255\\ 255\\ 255\\ 2$
Right Ascension	R.A. 19	h m	00 07.3 08.11 12.22 12.22 224.65 339.42 337.0 07.65 07.65 07.65 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 241.55 251.55 241.55 251.55 241.55 251.55 251.55 251.55 251.55 251.55 251.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 271.55 2
	Star	SUN	a And β Cas β Cas β Hyi β Hyi β Hyi α Cet β Cet β Cet β Phe <i>AB</i> β Phe <i>AB</i> β And β And β And β And α Eri τ Cet

	Sheratan	.5'' Almach Hamal	Polaris Mira Acamar	Menkar	Algol Mirfak	Alcyone			Idebaran
		$B5.4^{\rm m}C6.2^{\rm m}A-BC10''B-C0$ γ And =	Cep., R0.11 ^m 4.0 ^d , B ⁸ .9 ^m 18'' LP, R 2.0–10.1, 332 ^d , B 10 ^m 1'' A 3.57 ^m B 6.2 ^{3m} 3'' A 3.25 ^m B 4.36 ^m 8''		Ecl. R 2.06–3.28, 2.87 ^d Ecl. R 2.06–3.28, 2.87 ^d	in Pleiades	B9.36m 13'' B7.99m 9''	B 12 ^m 49′′	Silicon star Irr.? R0.78–0.93, B13 ^m 31 '' A
R	km/sec -12.6 -08.1 +07	-11.7 -14.3	+15.2 -17.4 +63.8 -05.1 +11.9	-25.9 +02.5	+28.2 +06.0 -02.4	+10.1 +10.1 +16.0	$^{+20.6}_{-01}$ +61.7	+35.6+38.6	+25.6 +54.1 +24.3 +17.5
ц	,, 0.230 0.038 0.147 0.265	0.068	0.156 0.046 0.232 0.203 0.061	0.075	0.035	0.050	0.015 0.036 0.126	0.064 0.118 0.108	0.051 0.202 0.468 0.021
D	1.y. 65 520 52 31	260 76	680 680 680 680 680	130	260 570 570 570	300 300	1000 680 160	390 160	260 68 330
$M_{\boldsymbol{\nu}}$	$^{+2.0}_{-2.7}$ +1.7 +2.9	-2.4 +0.2	-0.1 -4.6 +2.0 +1.7	-0.5 + 0.3	-1.0 -4.4	- 1.5	-6.1 -3.7 -0.5	-2.1 +0.1	-1.2 -0.7 +3.65 -2.4
π	" 0.050 0.007 0.063	0.005 0.043	0.012 0.003 0.013 0.028 0.028	0.003	0.008	0.002	0.007 001 0.003	0.008 0.018	0.011 0.048 0.125 0.015
Type	VI VI V	II	III Ib V III	III: + <u>A</u> 3:			.5 V III		dIII N V II
	F6 B3 A5 F0	K2 K3	<u> </u>	88 18 10 10 10 10 10 10 10 10 10 10 10 10 10	F188	<u>a ma</u>	MBB	60 ₹	KS K3 K3 K3
B^-V	+0.50 -0.15 -0.15 +0.14 +0.28	+1.16 +1.15	+0.13 +0.60 +0.11 +0.11	$^{+1.63}_{+0.72}$	-0.07+0.48	-0.14 -0.09 +1.61	+0.13 -0.17 +1.58	+0.91 +1.02	-0.08 +1.52 +0.45 +1.49
л	3.42 3.37 2.65 2.84	2.14: 2.00	$ \frac{3.00}{2.92} $	2.54 2.91:	3.5v 2.06v 1.80	2.86 3.30	2.83 2.96 2.96	3.33 3.54 7.7	3.28 3.28 3.17 2.68:
80 Dec.	。 + 29 29 + 63 34 + 20 43 - 61 40	+42 14 +23 22	$\begin{array}{c} +34 \\ +34 \\ +89 \\ -03 \\ 04 \\ +03 \\ 10 \\ -40 \\ 23 \end{array}$	+04 00 +53 25	+ 40 52 + 49 47	+4/ 44 +24 03 -74 18	+31 50 +39 57 -13 34	-62 32 + 19 08 + 15 49	-55 05 +16 28 +06 56 +33 08
R.A. 19	h m 01 52.0 52.9 53.6 58.1	02 02.7 06.1	08.4 12.5 42.2 57.5	03 01.2 03.3	03.7 06.6 22.9	46.3 47.5	52.7 56.5 57.1	04 14.1 27.5 27.5	33.5 34.8 48.3 55.7
Star	α Tri ε Cas β Ari α Hyi	γ And <i>A</i> α Ari	β Tri α UMi <i>A</i> ο Cet <i>A</i> θ Eri <i>AB</i>	α Cet γ Per	β Per β Per α Per	η Tau γ Hyi	ζ Per <i>A</i> ε Per <i>A</i> γ Eri	α Ret A ε Tau θ ² Tau	α Dor α Tau A π ³ Ori ι Aur

		Rigel Capella B4,98m1 // Bellatrix Elnath	53'' 29'' 11-i1 ₀₋₁₁	Atniam Phact Alnitak	Betelgeuse Ienkalinan ', var., 1.4 ^d	Canopus Alhena
	Ecl. R 0.81 ^m 9886 ^d	Manganese star Irr.? R 0.08–0.20, B 6.65 ^m 9'' Ecl. R 3.32–3.50, 8.0 ^d , A 3.59 ^m B 9.4 ^m 3''	Ecl. R 2.20–2.35 5.7 ^d , B 6.74 ^m A 3.56 ^m B 5.54 ^m 4'' C 10.92 ^m A 2.78 ^m B 7.31 ^m 11''	Shell star B 12 ^m 12'' A 1.91 ^m B4.05 ^m 3''	Irr.? R 0.06:-0.75: ^m Silicon star A 2.67 ^m B 7.14 ^m 3' R 0.77 ^m R 6.70 ^m 1''	R0.14 ^m BCMa type variable, 0.25 ^d
×	km/sec -01.4 +01.0 +07.4	+27.7 +20.7 +30.2 +19.8 +18.2 +08.0 -13.5	+22.0 +24.7 +27.6	+20.1 + 35 + 35 + 18.1 + 20.6 + 89.4	+21.0 -18.2 +29.3 +19.0	+32.2 + 54.8 + 33.7 + 20.5 - 12.5
=	,, 0.008 0.077 0.077 0.122	0.049 0.001 0.435 0.008 0.015 0.178 0.178	0.000 0.000 0.000 0.000 0.000	0.026 0.026 0.004 0.004	0.028 0.051 0.097	0.004 0.129 0.004 0.005 0.005
٩	1.y. 3400 370 370	390 900 940 300 113	1500 2000 2000	940 140 2100 140	520 88 108	390 160 98 105
Mr	-7.1 -0.4 +0.9	+ 0.22 + 0.21 + 0.21 + 0.22 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21 + 0.21	- 6.1 - 5.1 - 6.1 - 6.1	+	-5.6 +0.3 -0.4	-2.4 -0.6 -4.8 -3.1 -0.6
н	,, 0.004 0.013 0.042	$\begin{array}{c} 0.018\\ -0.03\\ 0.073\\ 0.004\\ 0.018\\ 0.018\\ 0.014\end{array}$	0.004 0.002 0.006 0.021	$^{-1.005}_{-0.002}$	0.005 0.037 0.018	$\begin{array}{c}003\\ 0.021\\ 0.014\\ 0.018\\ 0.031\end{array}$
Type	F0 Iap 33 V 43 III	89 11 38 11 58 11:+F 30.5 V 82 11 35 111 35 111 35 111 35 111 35 111 35 111 35 111 36 12 37 111 38 12 38 12 48 1	05.05 05 05 05 05 05 05 05 05 05 05 05 05 0	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	M2 Iab A2 V 39.5pv M3 III	32.5 V M3 III B1 II-III F0 Ib-II A0 IV
B-V	+0.50:] +1.46 +0.18 +0.13	+ 0.08 + - 0.18 + - 0.13 - 0.13 + - 0.13 - 0.13 - 0.13 - 0.13 - 0.13 - 0.13 - 0.13 - 0.13 - 0.13 - 0.04 - 0.05 - 0.04 - 0.05 - 0.05	-0.20 +0.22 -0.18 -0.24	-0.13: -0.13 : -0.13 : -0.11 -0.22 -0.22 $+1.16$ -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 -0.17 $-$	+1.87: +0.06 -0.07 +1.58	$\begin{array}{c} -0.18\\ +1.63\\ -0.24\\ +0.16\\ 0.00 \end{array}$
7	3.21 3.21 3.17 2.79	$\begin{array}{c} 3.29\\ 0.14v\\ 0.05\\ 3.32v\\ 1.66\\ 2.81\end{array}$	2.20v 3.40 2.76	3.07: 2.64 3.12 3.12	0.41v 1.86 2.65v 3 33v	3.04 2.92v 1.96v 1.93
80 Dec.	。 + 43 48 - 22 24 + 41 13 - 05 06	$\begin{array}{c} -16 \ 13 \\ +45 \ 59 \\ -20 \ 27 \\ -20 \ 47 \\ \end{array}$	-00 19 -17 51 +09 55 -05 56	-34 05 -34 05 -01 57 -09 41 -35 47	+07 24 +44 57 +37 13 +22 31	-30 03 +22 32 -17 56 +16 25 +16 25
R.A. 15	h m 05 00.5 04.6 05.1 06.9	13.6 13.5 23.5 23.5 23.5 23.5 23.5 23.5 23.5 2	31.0 31.8 34.5 34.5	39.5 39.7 39.7 50.2 50.2	54.0 58.0 58.4 06 13.7	21.7 21.7 21.8 23.5 36.6
Star	ε Aur ε Lep η Aur β Eri	μ Lep β Ori <i>A</i> α Aur γ Ori <i>AB</i> β Tau β Lep <i>A</i>	δ Ori A α Lep λ Ori AB ι Ori AB	ζ Tau ζ Oci <i>AB</i> κ Oci <i>AB</i> β Col	α Ori β Aur θ Aur <i>AB</i> η Gem <i>A</i>	ζ CMa μ Gem β CMa α Car γ Gem

	Sirius Adhara	'' Castor Procyon Pollux	Avior 12 ^m 20''
	B 8.66ª 1976: 11′′, p.a. 57° B 7.5ª 8′′	LP, R3.4-6.2, 141 ^d B9.4 ^m 22'' $B 10.7^m 4''$	Var. <i>R</i> 2.72–2.87, 0.14 ^d <i>B</i> 4.31 ^m 41'' <i>B</i> 15 ^m 7'' <i>A</i> 2.0 ^m <i>B</i> 5.1 ^m 3'' <i>CD</i> 10 ^m 69'' <i>A</i> 3.7 ^m <i>B</i> 5.2 ^m 0.2'' 15'', C6.8 ^m 3'' D <i>BC</i> 10.8 ^m 4''
R	km/sec +28.2 +09.9 +25.3 +25.3 -07.6 +36.4 +27.4	$\begin{array}{c} + + \\ + 33.0 \\ + 53.0 \\ + 53.0 \\ + 15.8 \\ + 23.0 \\ + 15.8 \\ + 23.0 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\ - 00.2 \\$	-24 +46.6 +35 +111.5 +192.8 +222.8 +222.8 +122.2
н	,, 0.010 0.224 1.324 0.272 0.079 0.004	$\begin{array}{c} 0.000\\ 0.342\\ 0.008\\ 0.008\\ 0.008\\ 0.195\\ 0.199\\ 0.199\\ 0.199\\ 0.095\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.005\\ 0.$	0.033 0.098 0.011 0.030 0.171 0.086 0.198 0.101 0.505
D	1.y. 620 64 64 8.7 57 124 680	3400 21000 650 140 210 180 180 180 180 11.3 35.3 11.3 35.3 430	2400 105: 520 340 150 140 220 49
M_{ν}	- 3.2 - 4.6 +1.9 +2.1 +2.1 -5.1	$\begin{array}{c} - & - & - & - & - & - & - & - & - & - $	+
я	" 0.009 0.375 0.375	$\begin{array}{c} - & 018 \\ 0 & 016 \\ 0 & 023 \\ 0 & 022 \\ 0 & 072 \\ 0 & 072 \\ 0 & 072 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\ 0 & 073 \\$	0.031 0.004 0.010 0.013 0.029 0.066
Type	[∃] ¹ 3× ² ∃∃	Ia Ia (gM5e) (gK4) (gK4) Ia V III II IV IV IV	III 111 111 111 111 111 111 111 1111 1
	B7 B2 B2 B2 B2 B2 B2 B2 B2 B2 B2 B2 B2 B2	B3 B3 B3 B3 B3 B3 B3 B3 B3 B3 B3 B3 B3 B	A7602KWC
B-V	-0.10 +1.39 +0.43 +0.21 +1.21 +1.21 -0.18:	$\begin{array}{c} + & - & - & - & - & - & - & - & - & - &$	-0.26 +0.42 +1.30: +1.30: +0.05 +10.05 +10.05 +10.05
Л	$\begin{array}{c} 3.19\\ 3.00\\ 3.38\\ -1.47\\ 2.92\\ 1.48\\ 1.48\end{array}$	$\begin{array}{c} 3.02\\ 1.85\\ 1.85\\ 2.91\\ 2.97\\ 3.34\\ 3.34\\ 3.34\\ 3.34\\ 3.34\\ 3.34\\ 3.48\\ 3.34\\ 3.48\\ 3.34\\ 3.48\\ 3.34\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\ 3.48\\$	2.23 2.80v 1.83 3.37 3.33 3.11 3.12 3.12
80 Dec.	。 43 11 +25 09 +12 55 -16 42 -61 55 -50 36 -28 37	-23 48 -26 225 -26 225 -37 04 -37 05 -37 05 -37 04 -37 05 -37 05	$\begin{array}{c} -39.57\\ -24157\\ -4718\\ -5926\\ +6047\\ +6632\\ +0630\\ +0630\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ +4807\\ $
R.A. 19	h m 66 37.1 42.7 44.2 44.2 48.2 57.8	07 07 07 07 07 07 07 07 07 07 07 07 07 0	08 02.9 06.7 08.9 08.9 22.1 28.6 54.3 54.3 54.3 57.9
Star	v Pup ε Gem ξ Gem α CMa A α Pic τ Pup ε CMa A	o ² CMa δ CMa L ₂ Pup Π Pup β CMa σ Pup A σ Pup A β Gem A β Gem A β Gem A β CMi A	ζ Pup ρ Pup γ Vel A ε Car ε UMA A δ Vel AB ε Hya ABC ζ Hya 1 UMa A

	Suhail	olacidus	Alphard		Regulus	Merak Dubhe
		Mia		B 14 ^m 5'' Cep. max. 3.4 ^m min. 4.8 ^m , 35.52 ^d A 3.02 ^m B 6.03 ^m 5''	B 8. 1 ^m 177' [,] Var. R 3.38–3.44 A 2.29 ^m B 3.54 ^m 4' [,] Var. R 3.22–3.39	A 2.7 ^m B 7.2 ^m 1'' A 1.88 ^m B 4.82 ^m 1''
R	km/sec + 18.4	+ 13 - 1	+37.6 +21.9 -04.3	-13.9 +15.4 +05.0 +04.0 +13.6	$\begin{array}{c} + \\ + \\ - \\ 15.0 \\ + \\ - \\ 18.3 \\ - \\ - \\ - \\ - \\ - \\ - \\ - \\ - \\ - \\ $	+0.06 -01.00 -012.00 -03.80 +07.80 -01
ㅋ	,, 0.026	0.183	0.034	0.004 1.094 0.016 0.016	0.248 0.029 0.170 0.023 0.023 0.023 0.023 0.021 0.021	0.085 0.221 0.221 0.087 0.087 0.087 0.087 0.033 0.033
۵	1.y. 750 800	286 86	180 9470 9470	2700 340 340 340	84 300 130 130 130 105 105 105 710	108 150 105 130 130 130 130 130 130 130 130 130 130
Μ _ν	-4.6	 		+1.8 +1.8 -5.5 -2.1 -2.1	-1.1 - 1.5	+ + + + + - + - + - + +
ц	<i></i> 0.015	0.038	0.021 0.007 0.017 0.017	0.022 0.019 0.020 0.020	0.039 0.009 0.018 0.019 0.031	0.022 0.042 0.031 0.040 0.019
Type	4 Ib-IIa IV V				V III III III III V III V III V III V III	
	54: 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1		A 8 4 7		2-232228 84 22 22 22 22 22 22 22 22 22 22 22 22 22	88824KKA KG
B-1	+1.6	-0. + +	+ + + + + + + + + + + + + + + + + + + +	+0.4		++ $++$ $++$ $++$ $++$ $++$ $++$ $++$
7	2.24	1.67	3.17 2.49 1.98	3.12 2.99 2.95 2.95	1.36 3.45 3.45 3.45 1.99 1.99 3.30v 2.74	2.67 3.12 3.12 3.15 3.34 3.15 3.15
80 Dec.	。 - 43 21 - 58 57	- 69 38	+ 34 29 - 54 56 - 08 35	+51 46 +23 51 -64 59	+12 +66 +23 +43 -61 +43 -61 14 -61 35 -61 35 -64 17	$\begin{array}{c} -49 \\ -16 \\ 05 \\ ++56 \\ 30 \\ ++61 \\ 52 \\ +120 \\ 33 \\ +120 \\ 33 \\ +14 \\ 24 \\ +120 \\ 33 \\ +14 \\ 24 \\ +14 \\ 24 \\ +14 \\ +14 \\ 24 \\ +14 \\ +14 \\ +16 \\ +14 \\ +16 \\ +14 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 \\ +16 $
R.A. 19	h m 09 07.3	13.0	21.5 26.6 26.6	31.5 44.7 46.6	10 07.3 15.7 15.7 15.9 15.9 16.4 16.4 21.1 31.4 42.2	45.9 48.6 02.5 03.6 33.2 48.0 88.0 48.0
Star	λ Vel a Car	B Car Car	α Lyn κ Vel α Hya N Vel	b UMa A E Leo I Car v Car AB	C C C C C C C C C C C C C C C C C C C	μ Vel AB v Hya β UMa α UMa AB v UMa AB δ Leo δ Leo δ Leo δ Leo

	Phecda		Megrez Gienah Acrux	Gacrux		r Crucis Alioth 10''		Mizar t., S pica	Alkaid	
					,	Beto um star star. $B5.61^{m}$	3	r, 708′′) .0 ^d ,βCMa vai		
		ır. R 2.56–2.62 ır R 2.78–2.84	5'', C 4,90 ^m 89''	8.26 ^m 24′′	rr. R 2.66–2.73 2.9m B 2.9m 2'' 3.50m B 3.52m 4'	3.7 ^m B 4.0 ^m 1'' CMa var., 0.25 ^d : 1romium-europi licon-europium		3.94" 14′′ (Alco :R 0.91–1.01,4	CMa var., 0.17 ^d	ır. <i>R</i> 3.08–3.17
	2 6.	<u>, 9, 4, 0</u> 2, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2	<u>بنەنى</u> ~	<u>o nr</u>	<u>77 × × × × × × × × × × × × × × × × × × </u>	<u>نەنە</u> 29 <u>7</u> 84	0.4.1.	<u>10.05</u> 10.050	<u></u>	.5 V ⁶ .5
R	km/s - 12	+++	1918	1++	- 101-	++42003	+ 105	-05 + 01 - 05 - 13 - 05 - 13 - 05 - 13 - 05 - 13 - 05 - 05 - 05 - 05 - 05 - 05 - 05 - 0	+ +	+12+01
н	,, 0.094	0.042	0.163	0.255 0.274 0.274 0.274	0.197	0.041 0.049 0.113 0.238	0.274 0.086 0.351	0.127 0.054 0.287	0.033 0.123 0.037	0.032 0.370 0.076
D	1.y. 90	370 570	370	220 220	$^{+160}_{-120}$	470 490 118	90 113 71	520 88 93	570 210 750	470 32 520
$M_{I\!\!P}$	+0.2	-2.7 -0.2 -3.4	+1.9	- +	+3.5	-2.1 -4.6 +0.2 +0.1	$^{+0.6}_{+0.3}$	$+0.1 \\ -3.3 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ +1.1 \\ $	3.9 - 2.1 - 3.4	-2.7 +2.7 -3.4
н	,, 0.020		700.0	0.018	0.006	0.008 0.023	0.036 0.021 0.046	0.037 0.021 0.035	0.004	0.102
Cype	٧	IVne III IV	>==>;	×× u: UII	N^2 21/2	> II >	ш-ш Л	>>>	∃>2	V:pne IV IV
	A 0	BZ K3	B0.5 B0.5	B9.5	58 <i>9</i> 8	B2 B0.5 A0pv B9.5p	42 69 78	B1 B3	B1 B3 B2	B2 G0 B2.5
B-V	0.00	-0.11: +1.33 -0.23	+0.0 -0.10 -0.25	-0.04 +1.55	+0.20++0.00++0.34	-0.17: -0.25 -0.03 -0.10	$^{+0.93}_{+0.05}$	+0.02 + 0.10	-0.23 -0.20 -0.22	-0.13: +0.59 -0.23:
V	2.44	2.59v 3.00 2.81v	2.59 2.59	1.80 2.97 1.69	2.70v 2.17 2.17	3.06 1.28v 1.79v 2.90v	2.83 2.98 2.76	2.26 0.91v 3.37	2.33v 1.87 3.42	3.12v 2.69 2.56
0 Dec.	° , +53 49	- 50 36 - 22 30 - 58 38	- 17 25 - 62 59	- 57 00	-69 01 -68 51 -01 20	-68 00 -59 35 +56 04 +38 26	+11 05 -23 04 -36 36	+55 02 -11 03 -00 30	- 53 22 + 49 25 - 41 35	-42 23 +18 30 -47 12
R.A. 198	h m 11 52.7	12 07.3 09.1 14.1	14.4 14.8 25.4	28.87 30.1	36.0 40.5 40.6	45.0 46.6 53.2 55.1	13 01.2 17.8 19.5	23.1 24.1 33.7	38.6 46.8 48.3	48.4 53.8 54.3
Star	UMa	SCOC	Crv Cru A	E C C S S S S S S S S S S S S S S S S S	Cen AB Vir AB	Mus AB Cru UMa CVn A	, Vir Hya Cen	UMa <i>A</i> Vir Vir	Cen Cen Cen	Cen

	var. Hadar Menkent Arcturus igil Kentaurus 8.61 ^m 16'' Zubenelgenubi Kochab	Alphecca Dschubba
	A 0.7 ^m B 3.9 ^m 1'', β CMa [*] Var, R 2.33-2.45 β 18'' R β CMa var., 0.26 ^d Strontum star. A 3.19 ^m B A 2.47 ^m B 5.04 ^m 3'', B 5.15 ^m 231''	B7.8 ^m 71'' B7.84 ^m 105'' Europium star β CMa var., 0.165 ^d A 3.5 ^m B3.7 ^m 1'' Ecl. R 0.11 ^m , 17.4 ^d A 3.47 ^m B7.70 ^m 15''
×	km/sec -12 +27.2 +27.2 -05.2 -05.2 -05.2 -05.2 -20.7 +07.3 +07.3 +07.3 +07.3 +07.3 +07.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +00.3 +	$\begin{array}{c} -19.9\\ -09.7\\ -04.3\\ -09.7\\ -00.3\\ -00.3\\ -00.3\\ -00.3\\ -00.3\\ -14\\ -01.7\\ -14\\ -11.0\\ -00.3\\ -11.0\\ -00.3\\ -11.0\\ -00.3\\ -11.0\\ -00.3\\ -00.3\\ -00.3\\ -00.3\\ -00.3\\ -00.3\\ -0.0\\ -0.3\\ -0.0\\ -0.3\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0.0\\ -0$
=	 0.035 0.156 0.156 0.738 2.284 0.738 2.284 0.186 0.033 0.051 0.033 0.066 0.033 	$\begin{array}{c} 0.059\\ 0.089\\ 0.135\\ 0.148\\ 0.101\\ 0.067\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.032\\ 0.$
D	$\begin{array}{c} \begin{array}{c} 1.9.\\ 4.90\\ 84\\ 84\\ 84\\ 84\\ 84\\ 84\\ 390\\ 86\\ 44.3\\ 44.3\\ 86\\ 103\\ 66\\ 103\\ 540\\ 770\\ 770\\ 770\\ 770\\ 770\\ 770\\ 770\\ 7$	140 58: 58: 570 570 570 570 570 570 570 570 570 570
$\mathbf{M}_{\mathbf{V}}$		+++++-+++++++++++++++++++++++++++++++
ĸ	$\begin{array}{c} & 0.016 \\ 0.039 \\ 0.039 \\ 0.090 \\ 0.016 \\ \end{array} \\ \begin{array}{c} 0.039 \\ 0.016 \\ 0.013 \\ 0.031 \end{array}$	$\begin{array}{c} 0.022\\ 0.056\\ 0.036\\ 0.036\\ 0.012\\ 0.005\\ 0.032\\ 0.043\\ 0.043\\ 0.078\\ 0.078\\ 0.078\end{array}$
Type	III III III III III V V:ne V:ne V:ne V:ne V III III III III III III III III III	
	B2 K4 B2 K4 CA B2 K4 CA CA CA CA CA CA CA CA CA CA CA CA CA	BB2
B-V	-0.23	$\begin{array}{c} + + \\ + \\ - \\ - \\ - \\ - \\ - \\ - \\ - \\ -$
7		346222228382842 3462228288888892 346228853288888888
0 Dec.	$\begin{array}{c} - & - & - & - & - & - & - & - & - & - $	$\begin{array}{c} + + + - + + + $
R.A. 198	h 14 02.4 05.5 14.8 33.1.3 38.4 40.7 57.3 8.4 40.7 57.3 8.4 57.3 57.3 57.3 8.4 57.3 8.4 57.3 8.4 57.3 57.3 8.4 57.3 8.4 57.3 8.4 57.3 8.4 57.3 8.4 57.3 57.4 57.4 57.4 57.4 57.4 57.4 57.4 57.4	15 01.2 10.8 10.8 10.8 10.8 17.1 17.1 20.1 17.1 20.1 20.1 20.2 20.3 20.1 20.3 20.1 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.9 20.
Star	$ \begin{array}{c} \beta \ {\rm Cen} \ AB \\ \pi \ {\rm Hya} \\ \theta \ {\rm Cen} \\ \alpha \ {\rm Boo} \\ \alpha \ {\rm Cen} \\ B \ {\rm Cen} \\ \alpha \ {\rm Cen} \\ B \ {\rm UMi} \\ \beta \ {\rm Lup} \\ \beta \ {\rm Lup} \\ \kappa \ {\rm Cen} \\ \kappa \ {\rm Cen} \end{array} $	β Boo σ Lib δ Lup A δ Lup A δ Lup A δ Lup δ Lup A δ Lup δ Lup A δ Lup A δ Lup A δ TrA δ Sco δ Sco δ Sco δ Sco δ Sco δ Sco δ Lup A δ δ Cup A δ δ Cup A δ δ δ δ δ δ δ δ δ δ δ δ δ δ δ δ δ δ δ

	4,	49 ^m 20′′	Amares	Atria		Sabik	ıs-Algethi		Shaula 1salhague
	4 2.78 ^m B 5.04 ^m 1′′, C 4.93 ^m 1	B CMa R 2.82-2.90, 0.25 ^d , B 8. B 8.7 ^m 6 ^{//} A 0 96m 1 00m D 6 07m 3 ^{//}	C	4 2.91m <i>B</i> 5.46m 1 ′′	Ecl. R 2.99–3.09, 1.4 ⁴	4 3.0 ^m <i>B</i> 3.4 ^m 1′′	$A 3.2^{m} \pm 0.3 B 5.4^{m} 5^{\prime\prime} R_{c}$	9 CM a var., 0.14 8 10m 18'' 8 11.49m 4''	3 CMa var., 0.21 ^d R.
R	km/sec -01.0 -19.9	- 14.3 - 14.3 - 14.3	-25.5 -25.5 -00.7 -19	- 69.9 + 08.3 - 03.6	$\begin{bmatrix} -02.5 \\ -25 \\ -55.6 \\ -06.0 \end{bmatrix}$	-14.1 -00.9 -28.4	-33.1 -41 -25.7	0.02 04 	+ 12.7 + 00 + 01.4
п	,, 0.027 0.156	0.030	0.105 0.030 0.022 0.022	0.608 0.097 0.044	$\begin{array}{c} 0.664 \\ 0.033 \\ 0.293 \\ 0.042 \end{array}$	0.026 0.097 0.293	0.032 0.164 0.029	0.017	0.031 0.260 0.012
D	1:y. 650 140	570	103 750 520	8233	520 90 90 90	620 69 52	410 96 710	1030 540 310 390	310 58 650
M_{V}	-3.7	+1.0 +4.4 +0.9	- +	+3.1 +2.1 -0.1	+0.7 -3.0 +0.1 +0.9	++1.3	- + - - 2.8 4.4		$-\frac{1}{3.3}$ +0.8 -4.6
π	,, 0.004 0.029	0.043	0.017	0.110 0.053 0.024	0.049 0.026 0.036	0.017 0.047 0.063	007 0.034 0.020	0.026 0.009	0.056
Type	νШ							יש _ם א משייים>	
	B0.5 M1	9222	89 89 2.60	852	B1.5 K2 K2 K2 K2 K2	B6 A2.5 F2	K3 W2	E E E E E E E E E E E E E E E E E E E	B1 A5 F0
B-V	-0.09 +1.59	+0.14	+0.92 +0.025 +0.00	+0.64 +1.43	+1.16 -0.20 +1.15 +1.61	-0.12 + 0.06 + 0.38	+1.41 +0.09 +1.43	-0.16: -0.22 -0.22 -0.22	-0.24 + 0.16 + 0.39
А	2.65	2.71 2.71	2.57	2.81 3.46	2.28 3.18 3.12	3.20 3.33 3.33	3.10v 3.14 3.13	2.90: 3.32 2.71 2.71	1.60v 2.09 1.86
80 Dec.	。 - 19 45 - 03 37	+ 61 33 + 61 33 - 25 32 - 25 - 25 - 25 - 25 - 25 - 25 - 25 - 2	$+21 \ 32 \ -28 \ 10 \ -10 \ 31$	+31 38 +31 38 +38 58 -68 60	-34 10 -38 01 +09 25 -55 57	+65 44 -15 42 -43 13	+14 24 +24 51 +36 49	-55 31 -56 22 +52 20 +52 20	$ \begin{array}{r} -37 \ 05 \\ +12 \ 35 \\ -42 \ 59 \\ \end{array} $
R.A. 19	h m 16 04.3 13.3	20.0 23.7	29.3 34.6 36.1	40.6 46.5 46.5	48.8 50.5 56.9	17 08.7 09.3 10.7	13.8 14.3 20 x	233.6 23.8 30.3 30.3 30.3	32.3 34.0 35.9
Star	B Sco AB 8 Oph	σ Sco A η Dra A Sco A	β Her ζ Oph	ζ Her AB η Her α TrA	ε sco μ ¹ Sco κ Oph ζ Ara	ζ Dra η Oph AB η Sco	α Her AB δ Her π Her Anh	β Ara Y Ara A υ Sco β Dra A	λ Sco α Oph θ Sco

	Eltanin		aus Australis Vega	.8 ^m 46'' Nunki		Albireo	Altair
	β CMa var., 0.20 ^d BC 9.78 ^m 33''	B 10 ^m 4′′	K	Eci. R 3.38–4.36, 12.9 ^d , <i>B</i> 7	A 3.3 ^m B 3.5 ^m < 1'' B 12 ^m 5'' A 3.7 ^m B 3.8 ^m C 6.0 ^m < 1''	B 5.11m 35'' A 2.91m B 6.44m 2''	
R	km/sec - 10 - 12.0 - 15.6 - 27.6 + 24.7 - 27.6 + 12.4	+22.1 +00.5 -20.0 +08.9	-11 -43.3 -13.9 +21.5	-17.8 -17.8 -11 -19.9 -21.5	+22 -26.3 -14 -09.8	+24.0 -29.9 -24.0 -21 -02.1	-26.3
п	" 0.031 0.160 0.811 0.004 0.064 0.026	$\begin{array}{c} 0.200\\ 0.218\\ 0.050\\ 0.894 \end{array}$	0.135 0.194 0.345 0.345	0.007	0.020 0.101 0.092 0.040	0.009 0.009 0.060 0.012	0.658
D	1.y. 470 124 33 3400 102 108 140	124 86: 84:	124 71 26.5	1300 300 370	140 160 86 250	53 53 270 340	16.5
$M_{\boldsymbol{\nu}}$	-3.4 +3.6 +3.6 +0.1 +0.7 +0.7 +0.2	+0.1 +1.1: +1.9	+ + 1.1 + 0.5 - 1.1	-2.7 +0.0 -2.1	++0.1	++	+2.2
π	" 0.023 0.108 0.013 0.013 0.017 0.017	$\begin{array}{c} 0.018\\ 0.038\\ 0.039\\ 0.054 \end{array}$	$\begin{array}{c} 0.015 \\ 0.046 \\ 0.123 \end{array}$	011 0.006 0.011	0.020 0.036 0.038 0.016 0.016	0.020 0.062 0.004 0.021 0.006	0.198
Type	.5 III VV III Ia	.5 III III III III-IV	5 111 11 V II		<i>IV</i> III-III III-III	5 III 5 III 1 III	IV-V
	9KK39KB	NKXX N N N N N N N N N N N N N N N N N N	K2. AQ		SERBAN	erzez	A7
B^-V	-0.21 +1.16 +0.75 +0.49 +1.18 +1.18 +1.62 +1.00	+1.00 +1.55 +1.39 +0.94	+0.02 $+1.05$ -0.00 $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.05$ $+1.0$	-0.05 -0.21 +1.18 -0.05	+0.08 +0.01 +0.10 +1.18 +0.35	+1.22	+0.22
V	2.39v 3.42 3.21 3.21 3.32 3.32	2.97 3.12 3.23	1.81 0.04 2.80	3.38v 3.38v 3.51 3.51	2.89 2.89 2.89 2.89	2.23	0.77
80 Dec.	- ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° °	-30 26 -36 47 -29 50 -02 54	- 34 24 - 25 27 - 37 46 - 37 01	+33 21 -26 19 -21 07 +32 40	$\begin{array}{c} -29 54 \\ +13 50 \\ -04 55 \\ -27 42 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 03 \\ -21 0$	+103 04 +27 55 +10 33	+08 49
R.A. 19	h m 17 41.1 42.5 45.7 46.2 48.4 56.1 58.0 58.0	18 04.5 16.3 19.7 20.2	22.9 26.7 36.2	54.0 58.2 58.2	19 01.3 04.5 05.2 05.7 08.6	24.5 29.9 44.3 45.3	49.8
Star	κ Sco β Oph ι' Sco G Sco γ Dra v Oph	Y Sgr n Sgr 8 Sgr n Ser	ε Sgr A Sgr A Lyr Ser	β Lyr σ Sgr γ Lyr	ζ Sgr AB ζ Aql A λ Aql τ Sgr ABC	δ Aql δ Cyg <i>A</i> δ Cyg <i>AB</i> γ Aql	ά Aql

	5'' Peacock Deneb	leramin Enif	ll Na'ir 1′′ nalhaut	Scheat Markab
	: B; <i>B</i> 5,97 ^m 20. <i>I</i>	6,0.19ª Ali	ر 5.4 ^d B6.19 ^m 4 <i>For</i>	r.
	Type gK0: + late	β CMa R 3.14–3.1 B 11 ^m 82″ Var. R 2.88–2.95	Cep. R 3.51-4.42, Var. R 2.11–2.23	Var. R 2.4–2.7
R	km/sec - 27.3 - 27.3 - 27.3 - 27.3 - 27.3 - 27.3 - 27.3 - 18.9 - 10.3 - 10.3 - 10.3 - 10.3	$\begin{array}{c} + 17.4 \\ - 10 \\ - 03.1 \\ + 06.5 \\ + 04.7 \\ - 00.2 \\ - 02.1 \end{array}$	$\begin{array}{c} + 07.5 \\ + 111.8 \\ - 18.4 \\ + 07.5 \\ + 011.6 \\ + 04.3 \\ + 06.5 \end{array}$	+08.7 -03.5 -42.4
ц	$\begin{array}{c} & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & &$	$\begin{array}{c} 0.056\\ 0.156\\ 0.156\\ 0.014\\ 0.017\\ 0.025\\ 0.392\\ 0.102\end{array}$	$\begin{array}{c} 0.016\\ 0.194\\ 0.015\\ 0.079\\ 0.077\\ 0.077\\ 0.077\\ 0.047\\ 0.047\\ 0.367\end{array}$	0.234 0.071 0.168
D	1.y. 330 330 330 330 330 310 160 160 160 160 160	390 52 980 780 540 540	1080 1240 1240 210 210 2280 360 322.6	210 109 51
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Ħ	× 0.008 0.005 − .006 0.039 0.039 0.013 0.026	$\begin{array}{c} 0.021\\ 0.063\\ 0.005\\ 0.006\\ 0.065\\ 0.065\\ 0.008\end{array}$	$\begin{smallmatrix} 0.003\\ 0.005\\ 0.003\\ 0.003\\ 0.003\\ 0.039\\ 0.144 \end{smallmatrix}$	0.015 0.030 0.064
Type	9.5 III 9.5 III 22.5 V 22.5 V 11 12 13 14 10 11 12 13 13 13 13 14 13 14 14 14 14 14 14 14 14 14 14 14 14 14	88 6 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	22 23 24 23 24 25 25 25 25 25 25 25 25 25 25	42 II-III 9.5 III 11 IV
B-V	H 10.00 H 10.0	++++0.224 +++0.224 +++0.224 ++-0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 +++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.237 ++0.2375 ++0.2375 ++0.2375 ++0.2375 ++0.20	++0.96 ++1.59 ++1.59 ++1.59 ++1.59 ++1.59 ++0.08 ++1.59 ++0.08 ++0.08 ++0.08	+1.02 B
7	3.24 3.26 3.11 3.15 3.45 3.45 2.45 2.45	3.19 2.44 2.38 2.38 3.00 3.00	$\begin{array}{c} 2.93\\ 1.76\\ 3.36\\ 2.87\\ 3.40\\ 2.17\\ 2.95\\ 3.28\\ 1.15\end{array}$	2.5 v 2.50 3.20
80 Dec.	 -00 52 -14 51 +40 11 -56 48 -47 21 +66 17 +66 17 +61 45 +33 53 	$\begin{array}{c} + 30 & 08 \\ + 62 & 31 \\ + 70 & 28 \\ - 05 & 40 \\ + 09 & 48 \\ - 16 & 13 \\ - 37 & 27 \\ \end{array}$	$\begin{array}{c} -00 & 25 \\ -47 & 047 \\ -60 & 216 \\ -60 & 216 \\ +10 & 448 \\ -115 & 566 \\ -29 & 44 \\ \end{array}$	$^{+27}_{+15}$ 58 $^{+15}_{+77}$ 30
R.A. 19	р 2010.3 21:5 221:5 221:5 221:5 44:9 45:9 45:9 45:9	21 12.1 18.2 28.4 28.5 30.5 45.9 52.7 52.7	22 04.7 10.1 17.1 28.5 28.5 440.5 53.6 53.6 56.5	23 02.8 03.8 38.5
Star	θ Aql β Cap A α Cyg α Cyg β Pav β Pav β Pav ε Cyg	ζ Cyg α Cep β Aqr β Aqr ε ε Peg A δ Cap	c Aqr c C Aqr c C Cep δ Cep A δ Aqr δ Aqr c PsA c PsA	β Peg α Peg γ Cep

DOUBLE AND MULTIPLE STARS

BY CHARLES E. WORLEY

Many stars can be separated into two or more components by use of a telescope. The larger the aperture of the telescope, the closer the stars which can be separated under good seeing conditions. With telescopes of moderate size and average optical quality, and for stars which are not unduly faint or of large magnitude difference, the minimum angular separation is given by 4.6/D, where D is the diameter of the telescope's objective in inches.

The following lists contain some interesting examples of double stars. The first list presents pairs whose orbital motions are very slow. Consequently, their angular separations remain relatively fixed and these pairs are suitable for testing the performance of small telescopes. In the second list are pairs of more general interest, including a number of binaries of short period for which the position angles and separations are changing rapidly.

In both lists the columns give, successively: the star designation in two forms; its right ascension and declination for 1975; the combined visual magnitude of the pair and the individual magnitudes; the apparent separation and position angle for 1976.0; and the period, if known.

Many of the components are themselves very close visual or spectroscopic binaries. (Other double stars appear in the table of The Brightest Stars and of The Nearest Stars.)

	Star	A.D.S.	R.A. 197 h m	Dec.	Magni comb.	tudes A B	P.A. 197	Sep. 76.0,	P (app.) years
$\lambda \\ \alpha \\ 33 \\ \Sigma \\ 55 \\ \Sigma \\ \epsilon^{1} \\ \epsilon^{2} \\ \pi \\ O\Sigma $	Cas Psc Ori 156 1338 Com 2054 Lyr† Lyr† Aql 500	434 1615 4123 5447 7307 8695 10052 11635 11635 11635 12962 16877	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccc} +54 & 24 \\ +02 & 39 \\ +03 & 16 \\ +18 & 13 \\ +38 & 18 \\ +21 & 23 \\ +61 & 45 \\ +39 & 38 \\ +39 & 38 \\ +31 & 44 \\ +44 & 18 \end{array}$	4.9 4.0 5.7 6.1 5.8 5.1* 5.6 5.1 4.4 5.6 5.9	5.5 $5.84.3$ $5.36.0$ $7.36.8$ $7.06.5$ $6.75.2$ $7.46.0$ $7.25.4$ $6.55.1$ $5.36.0$ $6.86.4$ 7.1	181 284 27 246 249 159 355 356 85 110 355	0.6 1.8 1.8 0.5 1.0 1.0 1.1 2.7 2.3 1.4 0.5	640 720 1100 400 500 1200 600 —
ηωγγΟααζζζογεγζεζτ7804 τ μς	Cas 186 And AB And BC 65 CMa Gem Cnc AB Cnc AB Cnc AC UMa Leo UMa Vir Boo Boo Her Oph Oph Oph Oph Cyg Aqr Cyg Cyg Cyg Cyg Cyg	671 1538 1630 2799 5423 6175 6650 7203 7724 8119 8630 9343 9413 10157 11005 11046 12880 14360 14367	00 47.4 01 54.5 02 02.1 02 02.1 03 48.9 06 44.1 07 33.0 08 10.8 09 08.2 10 18.6 11 16.8 12 40.4 14 50.2 16 40.4 18 01.6 19 44.2 20 50.1 21 13.7 21 43.0 22 75	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c} 3.5^{*}\\ 6.0\\ 2.1^{*}\\ 5.2\\ -1.4\\ .5.0\\ 5.2\\ 4.8^{*}\\ 1.6\\ 5.0\\ 4.8^{*}\\ 1.8\\ 3.8\\ 2.8\\ 3.8\\ 4.5\\ 2.8\\ 4.7\\ 4.0\\ 2.9^{*}\\ 6.0\\ 3.7\\ 4.6\end{array}$	3.5 6.81 5.8 5.84 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.84 2.55 5.90 2.22 2.64 5.52 2.22 2.64 5.52 2.22 2.64 5.52 2.22 2.64 2.52 2.22 2.64 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83 4.83	305 56 64 109 205 57 108 303 83 5 123 113 300 306 335 171 276 359 234 162 295	11.8 1.3 9.8 0.5 0.7 11.1 0.9 5.9 5.1 4.3 4.3 4.2 1.1 1.2 1.9 2.2 0.9 1.8	480 160 61 62 50 420 60 1100 620 60 1100 620 1100 620 5280 888 830 150 500 500 500 500 500 500 50

*There is a marked colour difference between the components.

†The separation of the two pairs of ε Lyr is 208".

(*Editor's Note* The co-ordinates of the stars in this table, which was extensively revised in 1975, are referred to the equinox of 1975, rather than to that of 1980, which is used elsewhere in this book.)

VARIABLE STARS

By JANET MATTEI

The systematic observation of variable stars is an area in which an amateur can make a valuable contribution to astronomy. For beginning observers, maps of the fields of four bright variable stars are given below. In each case, the magnitudes (with decimal point omitted) of several suitable comparison stars are given. Using two comparison stars, one brighter, one fainter than the variable, estimate the brightness of the variable in terms of these two stars. Record also the date and time of observation. When a number of observations have been made, a graph of magnitude versus date may be plotted. The shape of this "light curve" depends on the type of variable. Further information about variable star observing may be obtained from the American Association of Variable Star Observers, 187 Concord Ave., Cambridge, Mass. 02138.

In the tables the first column, the Harvard designation of the star, gives the 1900 position: the first four figures give the hours and minutes of R.A., the last two figures give the Dec. in degrees, italicised for southern declinations. The column headed *Max*, gives the mean maximum magnitude. The *Period* is in days. The *Epoch* gives the predicted date of the *earliest* maximum occurring this year; by adding the period to this epoch other dates of maximum may be found. The list of long-period variables has been prepared by the American Association of Variable Star Observers and includes the variables with maximum for several weeks. The second table contains stars which are representative of other types of variable. The data are taken from "The General Catalogue of Variable Stars" by Kukarkin and Parenago and for eclipsing binaries from *Rocznik Astronomiczny Obserwatorium Krakowskiego*, 1975, International Supplement.



LONG-PERIOD VARIABLE STARS

Variable	Max. m	Per d	Epoch 1976	Variable	Max. m	Per d	Epoch 1976
001755 T Cas 001838 R And 021143 W And 021403 o Cet 022813 U Cet 023813 R Tri 043065 T Cam 045514 R Lep 050953 R Aur 054920 U Ori 061702 V Mon 065355 R Lyn 070122aR Gem 070310 R CMi 072708 S CMi 07122aR Gem 070310 R CMi 07122aR Gem 070310 R CMi 072708 S CMi 07122aR Gem 070310 R CMi 072708 S CMi	$\begin{array}{c} 7.8\\ 7.04\\ 3.45\\ 7.5\\ 6.20\\ 8.67\\ 7.30\\ 7.5\\ 8.09\\ 7.10\\ 7.8\\ 7.8\\ 7.8\\ 7.8\\ 7.8\\ 7.8\\ 7.8\\ 7.8$	445 409 337 235 2266 374 432 459 372 335 370 338 332 257 288 372 257 288 372 257 213 313 302 317 355 257 146 226	Feb. 1 Dec. 19 Dec. 8 Jan. 28 Mar. 4 Jan. 29 Oct. 24 June 27 Nov. 30 Aug. 30 May 1 Apr. 23 July 1 Sept. 2 Apr. 26 Jan. 7 June 12 July 19 Mar. 26 Feb. 29 May 27 Apr. 21 Jan. 8 Dec. 13 Jan. 9 Jan. 24 Feb. 26	142539 V Boo 143227 R Boo 151731 S CrB 154639 V CrB 154615 R Ser 160625 RU Her 162119 U Her 162119 U Her 16212 V Oph 163266 R Dra 164715 S Her 170215 R Oph 171723 RS Her 180531 T Her 180531 T Her 180531 T Her 181136 W Lyr 183308 X Oph 190108 R Aql 191017 T Sgr 191019 R Sgr 193449 R Cyg 194048 RT Cyg 194048 RT Cyg 194048 RT Cyg 204405 T Aqr 210868 T Cep 213753 RU Cyg 230110 R Peg 230759 V Cas	$\begin{array}{c} 7.9\\ 7.3\\ 7.5\\ 6.9\\ 8.05\\ 7.5\\ 7.6\\ 7.9\\ 7.09\\ 7.09\\ 7.09\\ 7.09\\ 7.00\\ 7.5\\ 7.5\\ 7.5\\ 7.5\\ 7.5\\ 7.5\\ 7.5\\ 7.5$	258 223 361 358 357 484 406 228 245 307 302 219 165 196 334 300 302 269 426 190 407 465 202 330 234 378 228	July 5 July 7 Jan. 21 Oct. 15 Aug. 31 July 24 June 20 Mar. 3 Jan. 30 Mar. 8 Jan. 30 Mar. 8 Jan. 5 Aug. 1 Feb. 11 Apr. 29 Jan. 1 May 9 Jan. 1 June 28 Feb. 18 Apr. 9 Apr. 1 June 28 June 21 July 29 Apr. 6 Feb. 21 July 21
131546 V CVn 132706 S Vir 134440 R CVn 142584 R Cam	6.8 7.0 7.7 7.9	192 378 328 270	Jan. 24 Jan. 17 May 24 Apr. 11	231508 S Peg 2338/5 R Aqr 235350 R Cas 2357/5 W Cet	8.0 6.5 7.0 7.6	319 387 431 351	Oct. 17 Jan. 2 Mar. 12

OTHER TYPES OF VARIABLE STARS

$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Variable		Max. m	Min. m	Туре	Sp. Cl.	Period d	Epoch 1976 E.S.T.
222557 δ Cep 4.1 5.2 δ Cep F5-G2 5.366341 Jan. 4.70	005381 025838 030140 035512 060822 061907 065820 154428 171014 184205 184633 192242 194700 222557	U Cep ρ Per β Per λ Tau η Gem T Mon ζ Gem R Cr B α Her R Sct β Lyr R R Lyr η Aql δ Cep	6.7 3.3 2.1 3.5 3.1 6.4 4.4 5.8 3.0 6.3 3.4 6.9 4.1 4.1	9.8 4.0 3.3 4.0 3.9 8.0 5.2 14.8 4.0 8.6 4.3 8.0 5.2 5.2	Ecl. Semi R Ecl. Semi R δ Cep R Cr B Semi R RVTau Ecl. RR Lyr δ Cep δ Cep	B8+gG2 M4 B8+G B3 M3 F7-K1 F7-G3 cFpep M5 G0e-K0p B8 A2-F1 F6-G4 F5-G2	2.49302 33-55,1100 2.86731 3.952952 233.4 27.0205 10.15172 50-130, 6 yrs. 144 12.931163 0.5668223 7.176641 5.366341	Jan. 1.29* Jan. 2.22* Jan. 3.92* Jan. 6.43 Jan. 3.20 Jan. 3.34* Jan. 1.21 Jan. 2.75 Jan. 4.70

*Minimum.

BRIEF DESCRIPTION OF VARIABLE TYPES

Variables can be divided into three main classes; pulsating, eruptive and eclipsing binary stars as recommended by Commission 27 of the International Astronomical Union at its 12th General Assembly in Hamburg in 1964. A very brief and general description about the major types of variables in each class is given below.

I. Pulsating Variables

Cepheids: Variables that pulsate periodically with periods 1 to 70 days. They have high luminosity with amplitudes of light variations ranging from 0.1 to 2^{m} . Some of the group are located in open clusters, and they obey the well known period-luminosity relation. They are of F spectral class at maximum and G-K at minimum. The later their spectral class the greater is the period of light variation. Typical representative: δ Cephei.

RR Lyrae Type: Pulsating, giant variables with periods ranging from 0.05 to 1.2 and amplitude of light variation between 1 and 2^m. They are usually of A spectral class. Typical representative: RR Lyrae.

RV Tauri Type: Supergiant variables with light curves of alternating deep and shallow minima. The periods, defined as the interval between two deep minima, range from 30 to 150 days. The amplitude of light variations goes up to 3^{m} . Many show long term variations of 500 to 9000 days in their mean magnitude. Generally the spectral classes range from G to K. Typical representative: R Scuti.

Long period—Mira Ceti variables: Giant variables that vary with amplitudes from 2.5 to 5^{m} and larger with well defined periodicity, ranging from 80 to 1000 days. They show characteristic emission spectra of late spectral classes of Me, Ce and Se. Typical representative: \circ Ceti (Mira).

Semiregular Variables: Giants and supergiants showing appreciable periodicity accompanied by intervals of irregularities of light variation. The periods range from 30 to 1000 days with amplitudes not exceeding 1 to 2^m, in general. Typical representative: R Ursae Minoris.

Irregular Variables: Stars that show no periodicity or only a trace of it at times. Typical representative: ω Canis Majoris.

II. Eruptive Variables

Novae: Hot, dwarf stars with sudden increase in brightness, from 7 to 16^m in amplitude, in a matter of 1 to several to hundreds of days. After the outburst the brightness decreases slowly until its initial brightness is reached in several years or decades. Near the maximum brightness, spectra similar to A or F giants are usually observed. Typical representative: CP Puppis (Nova 1942).

Supernovae: Novae in a much larger scale, with sudden increase in brightness up to 20^m or more. The general appearance of their light curve is similar to novae. Typical representative: CM Tauri (central star of the Crab Nebula).

R Coronae Borealis Type: High luminosity variables with slow, non-periodic drops in brightness of amplitudes from about 1 to 9^m . The duration of minima varies from some dozen to several hundreds of days. Members of this type are of F to K and R spectral class. Typical representative: R Coronae Borealis.

U Geminorum Type: Dwarf novae that have long intervals of apparent quiesence at minimum with sudden rises to maximum. The range of outburst is from 2 to 6^m in light variations and ten to thousands of days between outbursts depending upon the star. It is a well established fact that most of the members are spectroscopic binaries with periods in order of hours. Typical representative: SS Cygni.

Z Camelopardalis Type: Variables similar to U Gem stars in their physical and spectroscopic properties. They show cyclical variations with intervals of constant brightness for several cycles, approximately one third of the way from maximum to minimum. Typical representative: Z Camelopardalis.

III. Eclipsing Binaries

Binary systems of stars with the orbital plane lying close to the line of sight of the observer. The components periodically eclipse each other, causing variations in the apparent brightness of the system, as is seen and recorded by the observer. The period of the eclipses coincides with the period of the orbital motion of the components. Typical representative: β Persei (Algol).

Editor's Note: In cooperation with the A.A.V.S.O., we plan to introduce one or two new variables to our readers each year. The following finding chart and light curve are for R CrB. Normally, it is 6°, but suddenly and unpredictably it may fade as much as 10 magnitudes, then slowly return to normal. It is an easy variable star to observe with binoculars.



THE NEAREST STARS

BY ALAN H. BATTEN

The accompanying table is similar to one that has been published in the HANDBOOK for several years. Like its predecessors, it is based on the work of Professor van de Kamp who has studied many of the nearest stars and published a revised list of them in 1969 in the Publications of the Astronomical Society of the Pacific. Since that list was published, four new stars have been found to have parallaxes of about 0''190 or greater and are therefore within the distance limit of about seventeen light years (or just over five parsecs) which has been arbitrarily set as the limit for this table. One of them, G158-27, has been included in the HANDBOOK since 1972; the other three, L725-32, B.D. -15° 6290, and B.D. 44° 2051, appear for the first time in the 1976 HANDBOOK. New determinations of the parallaxes of some of the stars in this list have also been published in the last few years. They have not been used because van de Kamp's discussion made use of all the data available for each star, and the inclusion of new data from single observatories for just a few stars would destroy the homogeneity of his list. The reader should remember, however, that new results may affect the order of stars in the list, and that the parallaxes of the new stars included will be relatively uncertain until more observations are available. The latest determination of the parallax of Stein 2051A and B is 0'.'179 and if this value is confirmed the stars should be dropped from the list.

Measuring the distances of stars is one of the most difficult and important jobs of an observational astronomer. As the earth travels around the sun each year, the directions of the nearer stars seem to change very slightly compared with those of more distant background stars. This change is called *annual parallax*; even for the nearest star it is less than one second of arc-the angle subtended by a penny about 2.5 miles away. That explains the difficulty of the task, and why results from different observatories are often slightly different. Parallax measurements are important because all our knowledge of the luminosities of stars, and hence of the structures of both the stars and the Galaxy, depends on the relatively few stellar distances that can be directly and accurately measured. The distances are so vast that new units are needed to describe them. Often we talk of light-years-the distance (nearly ten million million km or six million million miles) that light travels in a year-but in their own calculations astronomers use *parsecs*. One parsec is the distance of a star that has an annual parallax of one second of arc, and is equal to about 3.26 light years. The distance in parsecs is the reciprocal of the parallax expressed (as in the table) in seconds of arc.

The table gives the name and position of each star, the annual parallax π , the distance D in light-years, the spectral type, the proper motion μ in seconds of arc per year (that is the apparent motion of the star across the sky each year—nearby stars usually have large proper motions), the total space velocity W in km/sec (if known), the visual apparent magnitude and the luminosity in visible light in terms of that of the sun. In column 6, wd stands for white dwarf, and e indicates the presence of emission lines in the spectrum. Very few stars in our neighbourhood are brighter than the sun, and there are no very luminous or very hot stars at all. Most stars in this part of the galaxy are small, cool, and insignificant objects; we shall probably never be sure we have found them all.

The newest list contains 63 stars, including the Sun, thirty-one of which are single. There are eleven double-star systems and two triple systems. Earlier lists have emphasized the unseen companions believed to be associated with seven of the stars or systems. Recent work has called the reality of some of these into question—especially that of the supposed planetary companion of Barnard's star. The suspected companions are still indicated by asterisks in the table, but the evidence for several of them is no longer as clear as it appeared to be some years ago.

THE NEAREST STARS

	19	80							
Name	α.	δ	π	D	Sp.	μ	w	m	L
	hm	。,		l.y.		,,	km/sec		
Sun α Cen A	14 38	-60 46	0.760	4.3	G2 G2	3.68	32	-26.8 0.1	1.0
B C	14 28	-62 36			K5 M5e			1.5	0.36
Barnard's* Wolf 359	17 56	+04 36 +07 10	.552	5.9 7.6	M5 M6e	10.30	140	9.5	0.00044
Lal. 21185*	11 03	+3607	.402	8.1	M2	4.78	103	7.5	0.0052
B B	1 27	19 04	265	0.0	wd	2.25	10	7.2	0.008
Luy. 720-8A B	1 37	-18 04	. 303	0.9	M6e	5.55	32	12.5	0.00004
Ross 154 Ross 248	18 49 23 40	-23 50 +44 04	.345	9.4	M5e M6e	1.82	86	10.6	0.0004
ε Eri		-09 32	.305		K2	0.97	22	3.7	0.30
Ross 128	11 47	+0058	.301	10.8	M5	1.40	26	11.1	0.00033
61 Cyg A	21 06	+38 38	.292	11.2	K5	5.22	106	5.2	0.083
e Ind	22 03	-56 52	201	11.2	K7	4 67	86	6.0	0.040
Procyon A	07 39	$+05\ 17$.287	11.4	F5	1.25	21	0.3	7.6
B	10 40	1 50 20	204	11.5	wd	2 20	20	10.8	0.0005
2, 2398 A B	18 42	+ 59 30	. 284	11.5	M3.5 M4	2.29	39	8.9 9.7	0.0028
Groom. 34 A	00 18	+43 54	. 282	11.6	M1 M6	2.91	52	8.1	0.0058
Lacaille 9352	23 05	-35 59	.279	11.7	M2	6.87	117	7.4	0.012
τ Ceti	01 43	-16 03	.273	11.9	G8	1.92	37	3.5	0.44
BD+51008* L725-32		+05 27 -17 06	262	12.2	M4 M5e	1 31	/1	9.8	0.0014
Lacaille 8760	21 16	-38 58	.260	12.5	MI	3.46	67	6.7	0.025
Kapteyn's	05 11	-44 59	.256	12.7	M0	8.79	292	8.8	0.0040
Kruger 60 A	22 27	+57 36	. 254	12.8	M4 M6	0.87	31	9.7	0.0017
Ross 614 A	06 28	-02 48	. 249	13.1	M5e	0.97	30	11.3	0.0004
BD-12°4523	16 30	-12 36	.249	13.1	M 5	1.18	38	14.8	0.00002
van Maanen's	00 48	+05 19	.234	13.9	wdF	2.98	270	12.4	0.00017
Wolf 424 A	12 33	+09 09	. 229	14.2	M6e	1.87	39	12.6	0.00014
CD - 37°15492	00.04	-37 27	225	14 5	M6e M3	6 00	130	12.6	0.00014
G158 27	00 04	-07 38	.223	14.5	IVI 5	2.1	150	13.8	0.00005
Groom. 1618	10 10	+49 33	.217	15.0	M0	1.45	40	6.6	0.040
$CD - 46^{\circ}11540$	17 28	-46 53	.216	15.1	M4	1.15		9.4	0.0030
$CD = 49^{\circ}13313$ $CD = 44^{\circ}11909$	17 37	-49 11	213	15.2	M5	1 14		11 2	0.00063
Luy. 1159-16	01 59	+13 00	.212	15.4	(M7)	2.08		12.3	0.00023
Lal. 25372	13 44	+15 01	. 208	15.7	M3.5	2.30	55	8.5	0.0076
AOe 17415-6*	17 37	+68 22	.207	15.7	M3.5	1.31	34	9.1	0.0044
Ross 780	22 52	-1422	.200	15.8	M5	1.17	28	10.2	0.0016
o ² Eri A	04 14	-07 41	. 205	15.9	K0	4.08	104	4.4	0.33
B					wdA			9,9	0.0027
$BD = 15^{\circ}6290$	22 52	-14 22	205	15.9	M5	1.16		10.2	0.0016
BD+20°2465*	10 19	+19.58	.202	16.1	M4.5	0.49	15	9.4	0.0036
BD+44°2051	11 05	+43 36	.199	16.4	M2e	4.40		8.8	0.0063
Aitair 70 Oph A	19 49	+0849 +0231	196	16.0		0.00	29	4 2	0.44
B	10 05	102 51		10.7	K6		, .	6.0	0.083
AC+79°3888	11 46	+78 47	. 194	16.8	M4	0.87	121	11.0	0.0009
BD+43°4305*	22 46	+44 14	. 193	16.9	M5e (M5)	0.84	21		0.0021
B	04 30	+ 38 3/	. 192	17.0	wd	2.37		12.4	0.0003
	1		1	L	L		L		1

*Star may have an unseen component.

GALACTIC NEBULAE

By René Racine

The following objects were selected from the brightest and largest of the various classes to illustrate the different types of interactions between stars and interstellar matter in our galaxy. *Emission regions* (HII) are excited by the strong ultraviolet flux of young, hot stars and are characterized by the lines of hydrogen in their spectra. *Reflection nebulae* (Ref) result from the diffusion of starlight by clouds of interstellar dust. At certain stages of their evolution stars become unstable and explode, shedding their outer layers into what becomes a *planetary nebula* (P1) or a *supernova remnant* (SN). Protostellar nebulae (PrS) are objects still poorly understood; they are somewhat similar to the reflection nebulae, but their associated stars, often variable, are very luminous infrared stars which may be in the earliest stages of stellar evolution. Also included in the selection are four *extended complexes* (Compl) of special interest for their rich population of dark and bright nebulosities of various types. In the table S is the optical surface brightness in magnitude per square second of arc of representative regions of the nebula, and m* is the magnitude of the associated star.

			α 1980 δ			Sine	S		Dist.	
NGC	м	Con	h m	。,	Туре	Size	sq'	*	l.y.	Remarks
650/1 IC348 1435 1535 1952	76 1	Per Per Tau Eri Tau	01 40.9 03 43.2 03 46.3 04 13.3 05 33.3	$ \begin{array}{r} +51 & 28 \\ +32 & 07 \\ +24 & 01 \\ -12 & 48 \\ +22 & 05 \end{array} $	Pl Ref Ref Pl SN	1.5 3 15 0.5 5	20 21 20 17 19	17 8 4 12 16v	15 0.5 0.4 4	Nebulous cluster Merope nebula "Crab" + pulsar
1976 1999 ζ Ori 2068 IC443	42 78	Ori Ori Ori Ori Gem	05 34.3 05 35.5 05 39.8 05 45.8 06 16.4	$\begin{array}{r} -05 \ 25 \\ -06 \ 45 \\ -01 \ 57 \\ +00 \ 02 \\ +22 \ 36 \end{array}$	HII PrS Comp Ref SN	30 1 2° 5 40	18 20	4 10v	1.5 1.5 1.5 1.5 2	Orion nebula Incl. "Horsehead"
2244 2247 2261 2392 3587	97	Mon Mon Gem UMa	06 31.3 06 32.1 06 38.0 07 28.0 11 13.6	$\begin{array}{r} +04 \ 53 \\ +10 \ 20 \\ +08 \ 44 \\ +20 \ 57 \\ +55 \ 08 \end{array}$	HII PrS PrS Pl Pl	50 2 0.3 3	21 20 18 21	7 9 12v 10 13	3 3 4 10 12	Rosette neb. Hubble's var. neb. Clown face neb. Owl nebula
ρOph θOph 6514 6523 6543	20 8	Oph Oph Sgr Sgr Dra	16 24.4 17 20.7 18 01.2 18 02.4 17 58.6	$\begin{array}{r} -23 & 24 \\ -24 & 59 \\ -23 & 02 \\ -24 & 23 \\ +66 & 37 \end{array}$	Comp Comp HII HII Pl	4° 5° 15 40 0,4	19 18 15	11	0.5 3.5 4.5 3.5	Bright + dark neb. Incl. "S" neb. Trifid nebula Lagoon nebula
6611 6618 6720 6826 6853	16 17 57 27	Ser Sgr Lyr Cyg Vul	18 17.8 18 19.7 18 52.9 19 44.4 19 58.6	$-13 \ 48 \\ -16 \ 12 \\ +33 \ 01 \\ +50 \ 28 \\ +22 \ 40$	HII HII Pl Pl Pl	15 20 1.2 0.7 7	19 19 18 16 20	10 15 10 13	6 3 5 3.5 3.5	Horseshoe neb. Ring nebula Dumb-bell neb.
6888 γCyg 6960/95 7000 7009		Cyg Cyg Cyg Cyg Aqr	20 11.6 20 21.5 20 44.8 20 58.2 21 03.0	$^{+38\ 21}_{+40\ 12}_{+30\ 38}_{+44\ 14}_{-11\ 28}$	HII Comp SN HII Pl	15 6° 150 100 0.5	22 16	12	2.5 3.5 3	HII + dark neb. Cygnus loop N. America neb. Saturn nebula
7023 7027 7129 7293 7662		Cep Cyg Cep Aqr And	21 01.4 21 06.4 21 42.5 22 28.5 23 25.0	$^{+68}_{+42} \begin{array}{c} 05 \\ +42 \\ 09 \\ +65 \\ 00 \\ -20 \\ 54 \\ +42 \\ 25 \end{array}$	Ref Pl Ref Pl Pl	5 0.2 3 13 0.3	21 15 21 22 16	7 13 10 13 12	1.3 2.5 4	Small cluster Helix nebula

MESSIER'S CATALOGUE OF DIFFUSE OBJECTS

This table lists the 103 objects in Messier's original catalogue. The columns contain: Messier's number (M), the number in Dreyer's New General Catalogue (NGC), the constellation, the 1970 position, the integrated visual magnitude (m_v) , and the class of object. OC means open cluster, GC, globular cluster, PN, planetary nebula, DN, diffuse nebula, and G, galaxy. The type of galaxy is also indicated, as explained in the table of external galaxies. An asterisk indicates that additional information about the object may be found elsewhere in the *Handbook*, in the appropriate table.

M NGC	Con	α 1	980 δ	m _v	Туре	M	NGC	Con	α	198	30	δ	m⊮	Туре
1 1952 2 7089 3 5272 4 6121 5 5904	Tau Aqr CVn Sco Ser-	5 33.3 21 32.4 13 41.3 16 22.4 15 17.5	$^{+22}_{-00} \begin{array}{c} 01 \\ -00 \\ +28 \\ 29 \\ -26 \\ 27 \\ +02 \\ 11 \end{array}$	11.3 6.27 6.22 6.07 5.99	DN* GC* GC* GC* GC*	56 57 58 59 60	6779 6720 4579 4621 4649	Lyr Lyr Vir Vir Vir	19 18 12 12 12	15.8 52.9 36.7 41.0 42.6	+30 +33 +11 +11 +11 +11	0 08 01 56 47 41	8.33 9.0 9.9 10.3 9.3	GC PN* G-SBb G-E G-E
6 6405 7 6475 8 6523 9 6333 10 6254	Sco Sco Sgr Oph Oph	17 38.9 17 52.6 18 02.4 17 18.1 16 56.0	$ \begin{array}{r} -32 & 11 \\ -34 & 48 \\ -24 & 23 \\ -18 & 30 \\ -04 & 05 \\ \end{array} $	6 5 7.58 6.40	OC* OC* DN* GC GC*	61 62 63 64 65	4303 6266 5055 4826 3623	Vir Sco CVn Com Leo	12 16 13 12 11	20.8 59.9 14.8 55.7 17.8	+04 -30 +42 +21 +13	36 0 05 08 48 3 13	9.7 7.2 8.8 8.7 9.6	G-Sc GC G-Sb* G-Sb* G-Sa
11 6705 12 6218 13 6205 14 6402 15 7078	Sct Oph Her Oph Peg	18 50.0 16 46.1 16 41.0 17 36.5 21 29.1	$\begin{array}{r} -06 \ 18 \\ -01 \ 55 \\ +36 \ 30 \\ -03 \ 14 \\ +12 \ 05 \end{array}$	7 6.74 5.78 7.82 6.29	0C* GC* GC* GC GC*	66 67 68 69 70	3627 2682 4590 6637 6681	Leo Cnc Hya Sgr Sgr	11 8 12 18 18	19.1 50.0 38.3 30.1 42.0	+13 + 11 - 26 - 32 - 32	07 54 38 23 218	9.2 7 8.04 7.7 8.2	G-Sb OC* GC GC GC GC
16 6611 17 6618 18 6613 19 6273 20 6514	Ser Sgr Sgr Oph Sgr	18 17.8 18 19.7 18 18.8 17 01.3 18 01.2	$-13 \ 48 \\ -16 \ 12 \\ -17 \ 09 \\ -26 \ 14 \\ -23 \ 02$	7 7 7 6.94	OC* DN* OC GC DN*	71 72 73 74 75	6838 6981 6994 628 6864	Sge Aqr Aqr Psc Sgr	19 20 20 1 20	52.8 52.3 57.8 35.6 04.9	$+18 \\ -12 \\ -12 \\ +13 \\ -21$	44 39 44 5 41 59	6.9 9.15 9.5 8.31	GC GC OC G-Sc GC
21 6531 22 6656 23 6494 24 6603 25 4725†	Sgr Sgr Sgr Sgr Sgr	18 03.4 18 35.2 17 55.7 18 17.3 18 30.5	$\begin{array}{r} -22 & 30 \\ -23 & 55 \\ -19 & 00 \\ -18 & 27 \\ -19 & 16 \end{array}$	7 5.22 6 6 6	OC GC* OC* OC OC*	76 77 78 79 80	650 1068 2068 1904 6093	Per Cet Ori Lep Sco	1 2 5 5 16	40.9 41.6 45.8 23.3 15.8	+51-00+00-24-22	28 0 04 0 02 4 32 2 56	11.4 9.1 7.3 7.17	PN* G-Sb DN GC GC
26 6694 27 6853 28 6626 29 6913 30 7099	Sct Vul Sgr Cyg Cap	18 44.1 19 58.8 18 23.2 20 23.3 21 39.2	$\begin{array}{r} -09 & 25 \\ +22 & 40 \\ -24 & 52 \\ +38 & 27 \\ -23 & 15 \end{array}$	9 8.2 7.07 8 7.63	OC PN* GC OC GC	81 82 83 84 85	3031 3034 5236 4374 4382	UMa UMa Hya Vir Com	9 9 13 12 12	54.2 54.4 35.9 24.1 24.3	+69 +69 -29 +13 +18	09 047 046 00 00 18	6.9 8.7 7.5 9.8 9.5	G-Sb* G-Irr* G-Sc* G-E G-SO
31 224 32 221 33 598 34 1039 35 2168	And And Tri Per Gem	0 41.6 0 41.6 1 32.8 2 40.7 6 07.6	$ \begin{array}{r} +41 & 09 \\ +40 & 45 \\ +30 & 33 \\ +42 & 43 \\ +24 & 21 \end{array} $	3.7 8.5 5.9 6 6	G-Sb* G-E* G-Sc* OC OC*	86 87 88 89 90	4406 4486 4501 4552 4569	Vir Vir Com Vir Vir Vir	12 12 12 12 12	$25.1 \\ 29.7 \\ 30.9 \\ 34.6 \\ 35.8 \end{cases}$	+13 + 12 + 14 + 12 + 13	03 2 30 32 32 40 3 16	9.8 9.3 9.7 10.3 9.7	G-E G-Ep G-Sb G-E G-Sb
36 1960 37 2099 38 1912 39 7092 40 —	Aur Aur Aur Cyg UMa	5 35.0 5 51.5 5 27.3 21 31.5	+34 05 + 32 33 + 35 48 + 48 21	6 6 6 6	OC OC* OC OC 2 stars	91 92 93 94 95	6341 2447 4736 3351	Her Pup CVn Leo	17 7 12 10	16.5 43.6 50.1 42.8	+43 -23 +41 +11	- 3 10 3 49 1 14 1 49	6.33 6 8.1 9.9	M58? GC* OC G-Sb* G-SBb
41 2287 42 1976 43 1982 44 2632 45 —	CMa Ori Ori Cnc Tau	6 46.2 5 34.4 5 34.6 8 38.8 3 46.3	$\begin{array}{r} -20 \ 43 \\ -05 \ 24 \\ -05 \ 18 \\ +20 \ 04 \\ +24 \ 03 \end{array}$	6 4 2	OC* DN* DN OC* OC*	96 97 98 99 100	3368 3587 4192 4254 4321	Leo UMa Com Com Com	10 11 12 12 12	45.6 13.7 12.7 17.8 21.9	+11 + 55 + 15 + 15 + 15 + 15 + 15 + 15	56 5 08 5 01 4 32 5 56	9.4 11.1 10.4 9.9 9.6	G-Sa PN* G-Sb G-Sc G-Sc
46 2437 47 2422 48 2548 49 4472 50 2323	Pup Pup Hya Vir Mon	7 40.9 7 35.6 8 12.5 12 28.8 7 02 0	-14 46 -14 27 -05 43 +08 07 -08 19	7 5 6 8.9 7	OC* OC OC G-E*	101 102 103	5457 581	UMa Cas	14 1	02.5 31.9	+54 + 60	27 35	8.1 7	G-Sc* M101? OC
51 5194 52 7654 53 5024 54 6715 55 6809	CVn Cas Com Sgr Sgr	13 29.0 23 23.3 13 12.0 18 53.8 19 38.7	$\begin{array}{r} +47 & 18 \\ +61 & 29 \\ +18 & 17 \\ -30 & 30 \\ -31 & 00 \end{array}$	8.4 7 7.70 7.7 6.09	G-Sc* OC GC GC GC*									

STAR CLUSTERS

By T. SCHMIDT-KALER

The star clusters for this list have been selected to include those most conspicuous. Two types of clusters can be recognized: open (or galactic), and globular. Globulars appear as highly symmetrical agglomerations of very large numbers of stars, distributed throughout the galactic halo but concentrated toward the centre of the Galaxy. Their colour-magnitude diagrams are typical for the old stellar population II. Open clusters appear usually as irregular aggregates of stars, sometimes barely distinguished from random fluctuations of the general field. They are concentrated to the galactic disk, with colour-magnitude diagrams typical for the stellar population I of the normal stars of the solar neighbourhood.

The first table includes all well-defined open clusters with diameters greater than 40' or integrated magnitudes brighter than 5.0, as well as the richest clusters and some of special interest. NGC indicates the serial number of the cluster in Dreyer's New General Catalogue of Clusters and Nebulae, M, its number in Messier's catalogue, a and δ denote right ascension and declination, P, the apparent integrated photographic magnitude according to Collinder (1931), D, the apparent diameter in minutes of arc according to Trumpler (1930) when possible, in one case from Collinder; m, the photographic magnitude of the fifth-brightest star according to Shapley (1933) when possible or from new data, in italics; r, the distance of the cluster in kpcs (1 kpc = 3263 light-years), usually as given by Becker and Fenkart (1971); Sp, the earliest spectral type of cluster stars as a mean determined from three colour photometry and directly from the stellar spectra. The spectral type indicates the age of the cluster, expressed in millions of years, thus: O5 = 2, B0 = 8, B5 = 70, A0 = 400, A5 = 1000, F0 = 3000 and F5 = 10000.

The second table includes all globular clusters with a total apparent photographic magnitude brighter than 7.6. The first three columns are as in the first table, followed by *B*, the total photographic magnitude; *D*, the apparent diameter in minutes of arc containing 90 per cent of the stars, and in italics, total diameters from miscellaneous sources; *Sp*, the integrated spectral type; *m*, the mean blue magnitude of the 25 brightest stars (excluding the five brightest); *N*, the number of known variables; *r*, the distance in kpcs (absolute magnitude of RR Lyrae variables taken as $M_B = +0.5$); *V*, the radial velocity in km/sec. The data are taken from a compilation by Arp (1965); in case no data were available there, various other sources have been used, especially H. S. Hogg's Bibliography (1963).

		α 19	80 δ							
NGC	h	m	o	,	Р	D	m	r	Sp	Remarks
188 752 869 884 Perseus Pleiades Hyades 1912 1976/80 2099 2168 2232 2244 2264 2287 2362	00 01 02 03 03 04 05 05 06 06 06 06 06 06 06 07	42.0 56.6 17.6 21.0 45.9 19 27.3 34.4 51.1 07.6 25.5 31.3 39.9 46.2 18.0	$\begin{array}{r} + 85 \\ + 37 \\ + 57 \\ + 57 \\ + 48 \\ + 24 \\ + 15 \\ + 35 \\ - 05 \\ + 32 \\ + 24 \\ - 04 \\ + 09 \\ - 20 \\ - 24 \end{array}$	14 35 04 02 32 04 35 49 24 32 21 44 53 54 43 54	9.3 6.6 4.3 4.4 2.3 1.6 0.8 7.0 2.5 6.2 5.6 4.1 5.2 4.1 5.0 3.8	14 45 30 240 120 400 18 50 24 29 20 27 30 32 7	14.6 9.6 9.5 5 4.2 1.5 9.7 5.5 9.7 9.0 7 8.0 8.0 8.8 9.4	$\begin{array}{c} 1.55\\ 0.38\\ 2.15\\ 2.48\\ 0.17\\ 0.125\\ 0.040\\ 1.41\\ 1.28\\ 0.87\\ 0.49\\ 1.62\\ 0.72\\ 0.66\\ 1.64\\ \end{array}$	F2 A5 B1 B0 B1 B6 A2 B5 O5 B8 B5 B3 O5 O8 B4 O9	oldest known h Per χ Per, M supergiants moving cl., α Per M45, best known moving cl. in Tau* Trapezium, very young M37 M35 Rosette, very young S Mon M41 τ CMa
2422	07	34.7	-14	21	4.3	30	9.8	0.48	83	

OPEN CLUSTERS

*Basic for distance determination.
		α 19	80 δ												
NGC	h	m	0	'	Р	D	n	1	1	r .	Sp		Ren	narks	
2437 2437 2451 2516 2546 2632 11C2395 2682 3114 1C2602 Tr 16 3532 3766 Coma 4755 6067 6231 Tr 24 6405 1C4665 6475 6494 6523 6611 1C4725 1C4756 6705 Mel 227 1C1396 7790	07 44 07 4 07 4 07 5 08 1 08 3 08 4 08 4 10 0 10 4 10 4 10 4 10 4 11 0 11 3 12 2 5 16 1 16 5 17 3 17 4 17 5 17 5 18 0 18 1 18 3 18 3 18 5 7 20 0 21 3 23 5	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$		$\begin{array}{cccccccccccccccccccccccccccccccccccc$		$\begin{array}{c} 27\\ 37\\ 50\\ 45\\ 90\\ 45\\ 20\\ 18\\ 37\\ 65\\ 10\\ 55\\ 12\\ 300\\ 12\\ 16\\ 60\\ 26\\ 50\\ 27\\ 45\\ 8\\ 35\\ 50\\ 12.5\\ 60\\ 4.5\\ \end{array}$	$\begin{array}{c} 10 \\ 6 \\ 10 \\ 7 \\ 7 \\ 3 \\ 10 \\ 7 \\ 6 \\ 10 \\ 8 \\ 5 \\ 7 \\ 10 \\ 7 \\ 8 \\ 7 \\ 7 \\ 10 \\ 7 \\ 10 \\ 9 \\ 8 \\ 12 \\ 9 \\ 8 \\ 11 \\ \end{array}$.8 .1 .5 .1 .5 .1 .1 .5 .5 .1 .1 .5 .5 .3 .3 .42 .5 .5 .7	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	66 30 37 84 15 90 83 85 15 95 42 79 83 85 15 95 42 77 645 33 23 44 56 69 60 44 724 71 16	B8 B5 B8 B0 A0 B4 B2 E5 B1 B3 B3 O5 B8 B3 B3 O5 B8 B3 B3 O5 B8 B3 B3 O5 B8 B3 B3 O5 B8 B3 B3 O5 B8 B3 B3 O5 B3 B3 B3 B3 B3 B3 B3 B3 B3 B3 B3 B3 B3	M46 Prae M67 θ Ca η Ca Very κ Cr G an O suy M6 M7 M23 M8, ve N4 M16 M25 M11 Tr 3 C Ca	sepe,] , old c r r and spars u, "je d K s pergia Lago ry you GC65. , nebu , cepl , very 7 pph: C	M44 El. Nebula e cl. wel box upergia nts, WF oon neb ing cl. 30 ila heid, U rich cl. CEa, CE	a ints R-stars o. and Sgr Eb,
	-L				GI	OBULA	R (Clu	L Ste	RS					
			α 198	30 δ											
NGC	М	h	 m		• •	В		Γ	>	Sp		m	N	r	l v
104 4 1851 2808 5139 6 5272 5904 6121 6205 6218 6254 6341 6397 6546 6723 6752 6809 70789 7089	47 Tuc 0 Cen 3 5 4 13 12 10 92 22 55 15 2	00 05 09 13 15 16 16 16 16 16 17 17 18 18 18 19 19 21	23.1 13.3 11.5 25.6 41.3 17.5 22.4 41.0 46.1 56.0 16.5 39.2 06.5 35.1 58.3 09.1 38.8 29.1	-7 -4 -4 +2 +00 +4 -2 +30 -00 +4 -5 -44 -22 -33 -6 -33 +10	2 11 10 02 14 42 17 12 18 29 12 10 16 28 16 28 16 30 1 55 3 40 3 40 3 40 3 56 6 39 0 01 0 59 2 55	4.35 7.72 7.4 4.5 6.86 6.69 7.05 6.43 7.58 7.26 6.94 6.9 7.55 6.15 7.37 6.8 6.72 6.96		44 11 18 65 9 10 22 12 12 12 12 12 12 12 12 23 26 11 41 21 19 23 26 11 21 21 21 21 21 21 21 21 21	.5843769523 2279140	G2 F7 F8 F7 F7 F6 G0 F6 G0 F6 F8 G1 F1 F5 F6 F7 G4 F5 F6 F5 F2	3 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	13.54 5.09 3.01 4.07 3.21 3.85 4.07 4.17 3.96 2.71 3.45 3.73 4.32 3.36 3.68 4.47	111 3 4 165 189 97 43 10 1 3 16 3 1 24 19 1 6 103 22	5 14.0 9.1 5.2 10.6 8.1 4.3 6.3 7.4 6.2 7.9 2.9 4.0 3.0 7.4 5.3 6.0 10.5 12.3	$\begin{array}{r} -24\\ +309\\ +101\\ +230\\ -153\\ +49\\ +65\\ -241\\ -118\\ +111\\ -118\\ +111\\ -148\\ +11\\ -148\\ +170\\ -39\\ +170\\ -107\end{array}$

EXTERNAL GALAXIES

BY S. VAN DEN BERGH

Among the hundreds of thousands of systems far beyond our own Galaxy relatively few are readily seen in small telescopes. The first list contains the brightest galaxies. The first four columns give the catalogue numbers and position. In the column *Type*, E indicates elliptical, I, irregular, and Sa, Sb, Sc, spiral galaxies in which the arms are more open going from a to c. Roman numerals I, II, III, IV, and V refer to supergiant, bright giant, subgiant and dwarf galaxies respectively; p means "peculiar". The remaining columns give the apparent photographic magnitude, the angular dimensions and the distance in millions of light-years.

The second list contains the nearest galaxies and includes the photographic distance modulus $(m - M)_{pg}$, and the absolute photographic magnitude, M_{pg} .

NGCor		α 198	80 δ			Dimen-	Distance
name	М	h m	• •	Туре	m_{pg}		of l.y.
55 205 221 224 247	32 31	00 14.0 00 39.2 00 41.6 00 41.6 00 46.1	$\begin{array}{r} -39 \ 20 \\ +41 \ 35 \\ +40 \ 46 \\ +41 \ 10 \\ -20 \ 51 \end{array}$	Sc or Ir E6p E2 Sb I–II S IV	7.9 8.89 9.06 4.33 9.47	30 × 5 12 × 6 3.4 × 2.9 163 × 42 21 × 8.4	7.5 2.1 2.1 2.1 7.5
253 SMC 300 598 Fornax	33	00 46.6 00 52.0 00 54.0 01 32.8 02 38.7	$\begin{array}{r} -25 & 24 \\ -72 & 56 \\ -37 & 48 \\ +30 & 33 \\ -34 & 36 \end{array}$	Sep Ir IV or IV–V Sc III–IV Sc II–III dE	7.0: 2.86 8.66 6.19 9.1:	$22 \times 4.6216 \times 21622 \times 16.561 \times 4250 \times 35$	7.5 0.2 7.5 2.4 0.4
LMC 2403 2903 3031 3034	81 82	05 23.7 07 34.9 09 31.0 09 53.9 09 54.4	$ \begin{array}{r} -69 & 46 \\ +65 & 39 \\ +21 & 36 \\ +69 & 09 \\ +69 & 47 \end{array} $	Ir or Sc III–IV Sc III Sb I–II Sb I–II Scp:	0.86 8.80 9.48 7.85 9.20	$432 \times 432 \\ 22 \times 12 \\ 16 \times 6.8 \\ 25 \times 12 \\ 10 \times 1.5$	0.2 6.5 19.0 6.5 6.5
4258 4472 4594 4736 4826	49 104 94 64	12 18.0 12 28.8 12 38.8 12 50.0 12 55.8	+47 25 +08 06 -11 31 +41 13 +21 48	Sbp E4 Sb Sbp II: ?	8.90 9.33 9.18 8.91 9.27	19 × 7 9.8 × 6.6 7.9 × 4.7 13 × 12 10 × 3.8	14.0 37.0 37.0 14.0 12.0:
4945 5055 5128 5194 5236	63 51 83	13 04.1 13 14.8 13 24.2 13 29.0 13 36.0	$ \begin{array}{r} -49 & 22 \\ +42 & 08 \\ -42 & 54 \\ +47 & 18 \\ -29 & 46 \end{array} $	Sb III Sb II EOp Sc I Sc I–II	8.0 9.26 7.87 8.88 7.0:	$20 \times 48.0 \times 3.023 \times 2011 \times 6.513 \times 12$	14.0
5457 6822	101	14 02.4 19 43.8	+54 26 -14 49	Sc I Ir IV–V	8.20 9.21	$\begin{array}{c} 23 \times 21 \\ 20 \times 10 \end{array}$	14.0 1.7

THE BRIGHTEST GALAXIES

			α 198	80 δ						Dist.
Name	NGC	h	m	0	'	m _{pg}	$(m-M)_{pg}$	M_{pg}	Туре	of l.y.
M31	224	00	41.6	+41	10	4.33	24.65	-20.3	Sb I–II	2,100
Galaxy				-				?	Sb or Sc	
M33	598	01	32.8	+30	33	6.19	24.70	-18.5	Sc II–III	2,400
LMC		05	23.7	- 69	46	0.86	18.65	-17.8	Ir or SBc III–IV	160
SMC		00	52.0	-72	56	2.86	19.05	-16.2	Ir IV or IV–V	190
NGC	205	00	39.2	+41	35	8.89	24.65	-15.8	E6p	2.100
M32	221	00	41.6	+40	46	9.06	24.65	-15.6	E2	2,100
NGC	6822	19	43.8	-14	49	9.21	24.55	-15.3	Ir IV–V	1,700
NGC	185	00	37.8	+48	14	10.29	24.65	-14.4	E0	2,100
IC1613		01	04.0	+02	01	10.00	24.40	-14.4	Ir V	2,400
NGC	147	00	32.0	+48	14	10.57	24.65	-14.1	dE4	2,100
Fornax		02	38.7	- 34	36	9.1:	20.6:	-12:	dE	430
And I		00	44.4	+ 37	56	13.5:	24.65	-11:	dE	2,100
And II		01	15.3	+33	20	13.5:	24.65	-11:	dE	2,100
And III		00	34.3	+36	24	13.5:	24.65	-11:	dE	2,100
Leo I		10	07.4	+12	24	11.27	21.8:	-10:	dE	750:
Sculptor		00	58.9	- 33	49	10.5	19.70	-9.2:	dE	280:
Leo II		11	12.4	+22	16	12.85	21.8:	-9:	dE	750:
Draco		17	19.8	+ 57	56		19.50	?	dE	260
Ursa Minor		15	08.5	+ 67	11		19.40	?	dE	250

THE NEAREST GALAXIES

MAXIMA OF DELTA CEPHEI

A finding chart for this famous pulsating variable is given on p. 98. The magnitudes (minus decimal point) of non-variable comparison stars are marked; the magnitude of δ Cep can be estimated relative to these. Observation of this star, or of Algol, is a good introduction to serious variable star observing, and is a good project for the amateur or student.

Times given are E.S.T., rounded off to the nearest 10 minutes, and are based on the ephemeris J.D. (max) = 2436075.445 + 5.366341 E.

Date	Time	Date Time	Date Time	Date	Time
Jan. 5	15 ^h 50 ^m	Apr. 5 21 ^h 20 ^m	July 6 2 ^h 50 ^m	Oct. 5	8 ^h 20 ^m
11	0 40	11 6 10	11 11 40	10	17 10
16	9 30	16 15 00	16 20 30	16	1 50
21	18 20	21 23 50	22 5 10	21	10 40
27	3 00	27 8 30	27 14 00	26	19 30
Feb. 1	11 50	May 2 17 20	Aug. 1 22 50	Nov. 1	4 20
6	20 40	8 2 10	7 7 40	6	13 00
12	5 30	13 11 00	12 16 20	11	21 50
17	14 10	18 19 40	18 1 10	17	6 40
22	23 00	24 4 30	23 10 00	22	15 30
28	7 50	29 13 20	28 18 50	28	0 10
Mar. 4	16 40	June 3 22 10	Sept. 3 3 30	Dec. 3	9 00
10	1 25	9 6 50	8 12 20	8	17 50
15	10 10	14 15 40	13 21 00	14	2 40
20	19 00	20 0 30	19 6 00	19	11 20
26	3 50	25 9 20	24 14 40	24	20 10
31	12 40	30 18 00	29 23 30	30	5 00

RADIO SOURCES

By John Galt

Although several thousand radio sources have been catalogued most of them are only observable with the largest radio telescopes. This list contains the few strong sources which could be detected with amateur radio telescopes as well as representative examples of astronomical objects which emit radio waves.

	α (19	80) δ						
Name	h m	• •	Remarks					
Tycho's s'nova	00 24.6	+64 01 +41 09 +62 01 -23 14 +41 26	Remnant of supernova of 1572					
Andromeda gal.	00 41.5		Closest normal spiral galaxy					
IC 1795, W3	02 23.9		Multiple HII region, OH emission					
PKS 0237–23	02 39.1		Quasar with large red shift $Z = 2.2$					
NGC 1275, 3C 84	03 18.5		Seyfert galaxy, radio variable					
Fornax A	03 21.6	$ \begin{array}{r} -37 & 15 \\ +54 & 29 \\ +22 & 00 \\ +22 & 00 \\ +01 & 54 \end{array} $	10th mag. SO galaxy					
CP 0328	03 31.3		Pulsar, period = 0.7145 sec., H abs'n.					
Crab neb, M1*	05 33.2		Remnant of supernova of 1054					
NP 0532	05 33.2		Radio, optical & X-ray pulsar					
V 371 Orionis	05 32.7		Red dwarf, radio & optical flare star					
Orion neb, M42	05 34.3	$\begin{array}{r} -05 \ 24 \\ +22 \ 36 \\ +04 \ 53 \\ -20 \ 42 \\ +02 \ 10 \end{array}$	HII region, OH emission, IR source					
IC 443	06 16.1		Supernova remnant (date unknown)					
Rosette neb	06 30.9		HII region					
YV CMa	07 22.2		Optical var. IR source, OH, H ₂ O emission					
3C 273	12 28.0		Nearest, strongest quasar					
Virgo A, M87*	12 29.8	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	EO galaxy with jet					
Centaurus A	13 24.2		NGC 5128 peculiar galaxy					
3C 295	14 10.7		21st mag. galaxy, 4,500,000,000 light years					
Scorpio X-1	16 18.8		X-ray, radio optical variable					
3C 353	17 19.5		Double source, probably galaxy					
Kepler's s'nova	17 27.6	$\begin{array}{r} -21 & 16 \\ -28 & 56 \\ -16 & 10 \\ +09 & 05 \\ +21 & 50 \end{array}$	Remnant of supernova of 1604					
Galactic nucleus	17 44.3		Complex region OH, NH ₃ em., H ₂ COabs'n.					
Omega neb, M17	18 19.3		HII region, double structure					
W 49	19 09.4		HII region s'nova remnant, OH emission					
CP 1919	19 20.8		First pulsar discovered, $P = 1.337$ sec.					
Cygnus A*	19 58.7	+40 41	Strong radio galaxy, double source					
Cygnus X	20 21.9	+40 19	Complex region					
NML Cygnus	20 45.8	+40 02	Infrared source, OH emission					
Cygnus loop	20 51.4	+29 36	S'nova remnant (Network nebula)					
N. America	20 54.4	+43 59	Radio shape resembles photographs					
3C 446 Cassiopeia A* Sun* Moon Jupiter*	22 24.7 23 22.5	-05 04 + 58 42	Quasar, optical mag. & spectrum var. Strongest source, s'nova remnant Continuous emission & bursts Thermal source only Radio bursts controlled by Io					

*Could be detected with amateur radio telescopes.



The above map represents the evening sky on the dates and times shown. For earlier (or later) dates, add (or subtract) two hours per month. For instance, the map represents the early morning sky in late October at 4 a.m. The map is drawn for latitude 45° N. but is useful for latitudes several degrees north or south of this.

The centre of the map is the *zenith*, the point directly overhead; the circumference of the map is the *horizon*. To identify the stars, hold the map in front of you so that the part of the horizon which you are facing (north, for instance) is downward.



The above map represents the evening sky on the dates and times shown. For earlier (or later) dates, add (or subtract) two hours per month. For instance, the map represents the early morning sky in late December at 4 a.m. The map is drawn for latitude 45° N, but is useful for latitudes several degrees north or south of this.

The centre of the map is the *zenith*, the point directly overhead; the circumference of the map is the *horizon*. To identify the stars, hold the map in front of you so that the part of the horizon which you are facing (north, for instance) is downward.



The above map represents the evening sky on the dates and times shown. For earlier (or later) dates, add (or subtract) two hours per month. For instance, the map represents the early morning sky in late February at 4 a.m. The map is drawn for latitude 45° N, but is useful for latitudes several degrees north or south of this.

The centre of the map is the *zenith*, the point directly overhead; the circumference of the map is the *horizon*. To identify the stars, hold the map in front of you so that the part of the horizon which you are facing (north, for instance) is downward.



The above map represents the evening sky on the dates and times shown. For earlier (or later) dates, add (or subtract) two hours per month. For instance, the map represents the early morning sky in late April at 4 a.m. The map is drawn for latitude 45° N, but is useful for latitudes several degrees north or south of this.

45° N, but is useful for latitudes several degrees north or south of this. The centre of the map is the *zenith*, the point directly overhead; the circumference of the map is the *horizon*. To identify the stars, hold the map in front of you so that the part of the horizon which you are facing (north, for instance) is downward.



The above map represents the evening sky on the dates and times shown. For earlier (or later) dates, add (or subtract) two hours per month. For instance, the map represents the early morning sky in late June at 4 a.m. The map is drawn for latitude 45° N, but is useful for latitudes several degrees north or south of this.

The centre of the map is the *zenith*, the point directly overhead; the circumference of the map is the *horizon*. To identify the stars, hold the map in front of you so that the part of the horizon which you are facing (north, for instance) is downward.



The above map represents the evening sky on the dates and times shown. For earlier (or later) dates, add (or subtract) two hours per month. For instance, the map represents the early morning sky in late August at 4 a.m. The map is drawn for latitude 45° N, but is useful for latitudes several degrees north or south of this.

The centre of the map is the *zenith*, the point directly overhead; the circumference of the map is the *horizon*. To identify the stars, hold the map in front of you so that the part of the horizon which you are facing (north, for instance) is downward.

CALENDAR

Jai	nua	ary					Fe	bri	uar	y				Ma	irc	h					Ap	ril					
S	Μ	Ť	w	Т	F	S	S	M	T	W	Т	F	S	S	M	Т	W	T	F	S	ŝ	M	Т	w	Т	F	S
				1	2	3	1	2	3	4	5	6	7		1	2	3	4	5	6					1	2	3
4	5	6	7	8	9	10	8	9	10	11	12	13	14	7	8	9	10	11	12	13	4	5	6	7	8	9	10
11	12	13	14	15	16	17	15	16	17	18	19	20	21	14	15	16	17	18	19	20	11	12	13	14	15	16	17
18	19	20	21	22	23	24	22	23	24	25	26	27	28	21	22	23	24	25	26	27	18	19	20	21	22	23	24
25	26	27	28	29	30	31	29							28	29	30	31				25	26	27	28	29	30	
Ma	iv				a.		Jur	ne						Jul	v						Au	gu	st				
S	M	Т	w	Т	F	S	S	M	Т	w	Т	F	S	S	M	Т	w	Т	F	S	S	M	Т	w	Т	F	S
						1			1	2	3	4	5					1	2	3	1	2	3	4	5	6	7
2	3	4	5	6	7	8	6	7	8	9	10	11	12	4	5	6	7	. 8	9	10	8	9	10	11	12	13	14
9	10	11	12	13	14	15	13	14	15	16	17	18	19	11	12	13	14	15	16	17	15	16	17	18	19	20	21
16	17	18	19	20	21	22	20	21	22	23	24	25	26	18	19	20	21	22	23	24	22	23	24	25	26	27	28
23	24	25	26	27	28	29	27	28	29	30				25	26	27	28	29	30	31	29	30	31				
30	31																										
Se	pte	em	bei	r			Oc	to	bei	r				No	ve	m	bei				De	ce	mt	ber			
S	M	Т	w	Т	F	S	S	M	Т	W	Т	F	S	S	Μ	Т	w	Т	F	S	S	Μ	Т	w	Т	F	S
			1	2	3	4						1	2		1	2	3	4	5	6				1	2	3	4
5	6	7	8	9	10	11	3	4	5	6	7	8	9	7	8	9	10	11	12	13	5	6	7	8	9	10	11
12	13	14	15	16	17	18	10	11	12	13	14	15	16	14	15	16	17	18	19	20	12	13	14	15	16	17	18
19	20	21	22	23	24	25	17	18	19	20	21	22	23	21	22	23	24	25	26	27	19	20	21	22	23	24	25
26	27	28	29	30			24	25	26	27	28	29	30	28	29	30			Res .		26	27	28	29	30	31	
200			23				31				22.0					S. C.					1963				1		

CALENDAR

January	February	March	April
SMTWTFS	SMTWTFS	SMTWTFS	SMTWTFS
1	1 2 3 4 5	1 2 3 4 5	. 1 2
2 3 4 5 6 7 8	6 7 8 9 10 11 12	6 7 8 9 10 11 12	3 4 5 6 7 8 9
9 10 11 12 13 14 15	13 14 15 16 17 18 19	13 14 15 16 17 18 19	10 11 12 13 14 15 16
16 17 18 19 20 21 22	20 21 22 23 24 25 26	20 21 22 23 24 25 26	17 18 19 20 21 22 23
23 24 25 26 27 28 29	27 28	27 28 29 30 31	24 25 26 27 28 29 30
30 31			
May	June	July	August
SMTWTFS	SMTWTFS	SMTWTFS	SMTWTFS
1 2 3 4 5 6 7	1 2 3 4	1 2	1 2 3 4 5 6
8 9 10 11 12 13 14	5 6 7 8 9 10 11	3 4 5 6 7 8 9	7 8 9 10 11 12 13
15 16 17 18 19 20 21	12 13 14 15 16 17 18	10 11 12 13 14 15 16	14 15 16 17 18 19 20
22 23 24 25 26 27 28	19 20 21 22 23 24 25	17 18 19 20 21 22 23	21 22 23 24 25 26 27
29 30 31	26 27 28 29 30	24 25 26 27 28 29 30	28 29 30 31
		31	
September	October	November	December
SMTWTFS	SMTWTFS	SMTWTFS	SMTWTFS
1 2 3	1	1 2 3 4 5	1 2 3
4 5 6 7 8 9 10	2 3 4 5 6 7 8	6 7 8 9 10 11 12	4 5 6 7 8 9 10
11 12 13 14 15 16 17	9 10 11 12 13 14 15	13 14 15 16 17 18 19	11 12 13 14 15 16 17
18 19 20 21 22 23 24	16 17 18 19 20 21 22	20 21 22 23 24 25 26	18 19 20 21 22 23 24
25 26 27 28 29 30	23 24 25 26 27 28 29	27 28 29 30	25 26 27 28 29 30 31
	30 31		

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